

Outlier Detection in Spectroscopy & Photometry

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Astrophysical Sciences & CSML

Spectroscopic Analysis: Spender

Code: https://github.com/pmelchior/spender

Melchior et al. (2023), AJ, 166, 74



Forward Model

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Analysis Model

Forward Model

Low-Dimensional Galaxy Spectrum Model



Melchior et al. (2023), AJ, 166, 74

Super-Resolution Model

- * Internally × 2 oversampling
- * Deconvolve from the (unknown) SDSS-I LSF



The Latent Space

Liang et al. (2023) AJ, 166, 75





Yan Liang

Latent Space Distribution

Liang et al. (2023) AJ, 166, 75



Outliers in SDSS-I Main Galaxy Sample





in DESI

- 1. Stars
- 2. Blends
- 3. Velocity structure
- 4. weird QSOs

Liang et al. (2023) ApJ, 956, L6

Feature Attribution for Outliers

Shen & Melchior (2023) arXiv:2310.20012





Jeff Shen

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Jeff Shen

Galaxy Photometry: PROVABGS

Code: https://github.com/changhoonhahn/provabgs

Stellar population synthesis

$$f_{\rm csp}(\lambda, t) = \int_{t'=0}^{t'=t} f_{\rm ssp}(t', Z(t')) \operatorname{SFR}(t') e^{-\tau_{\rm dust}(t')}$$

Name

 $\log M_*$

 β_4

fburst

t_{burst}

 γ_1, γ_2

 $au_{\rm BC}$

 $\tau_{\rm ISM}$

*n*_{dust}

 f_{fiber}

 $\beta_1, \beta_2, \beta_3,$

- Expensive evaluation
- Traditionally O(h)/galaxy
- Not feasible for O(10⁹)
 galaxies!
- * Normalizing flow for $p(\theta \mid \mathscr{D})$

 Table 1

 Parameters of the PROVABGS SPS Model and Their Priors Used for Joint SED Modeling of DESI Photometry and Spectroscopy

Description

log galaxy stellar mass

NMF basis coefficients

for SFH

fraction of total stellar mass

formed in starburst event

time of starburst event

NMF basis coefficients

for ZH

Birth cloud optical depth

diffuse-dust optical depth

Calzetti (2001) dust index spectrum fiber-aperture effect normalization	uniform over $[-2, 1]$ Gaussian $\mathcal{N}(\hat{f}_r^{\text{fiber}}, \frac{f_r^{\text{fiber}}}{f_r}\sigma_r)$
Hahn et al. (2023), ApJ, 945, 16	



ChangHoon Hahn

Prior

uniform over [7, 12.5]

Dirichlet prior

uniform over [0, 1]

uniform over [10 Myr,

13.2 Gyr]

log uniform over $[4.5 \times 10^{-5}]$,

 1.5×10^{-2}]

uniform over [0, 3]

uniform over [0, 3]



SEDFlow

Code: https://github.com/changhoonhahn/SEDflow

 $p(\theta \mid \{g, r, i, z\}, \{\sigma_g, \sigma_r, \sigma_i, \sigma_z\})$

- O(1s) instead ofO(100h)!
- Done withbroadbandphotometry
- Application to spectra coming
- * Built explicit model $p(\{\sigma_x\} \mid \{\max_x\})$



Outliers in Photometry: Shroom

Code: https://github.com/astro-data-lab/shroom



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- * **Self-evident**: Detection hard, interpretation harder
- **Boring:** (ideally)
 Success of outlier detection measured by our surprise

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SCHMIDT FUTURES



