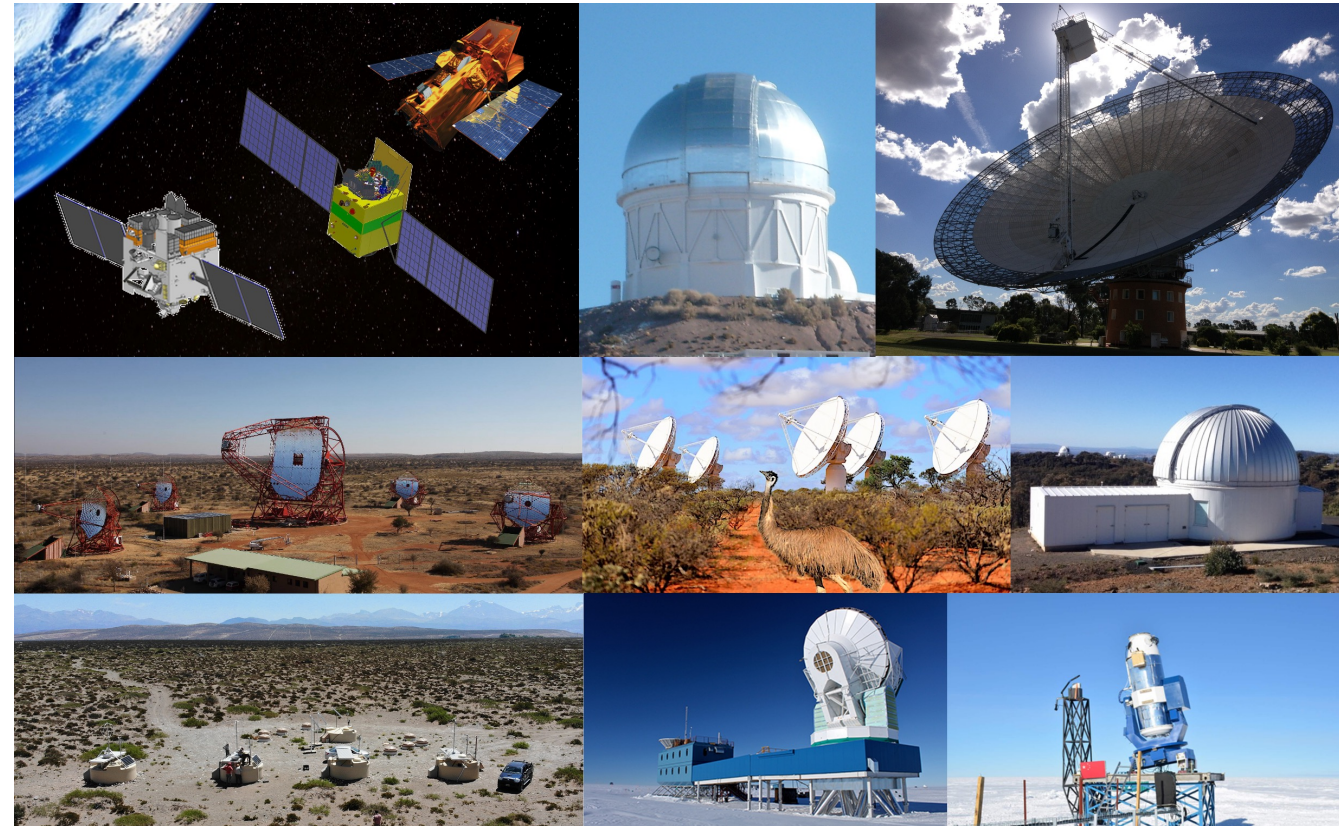
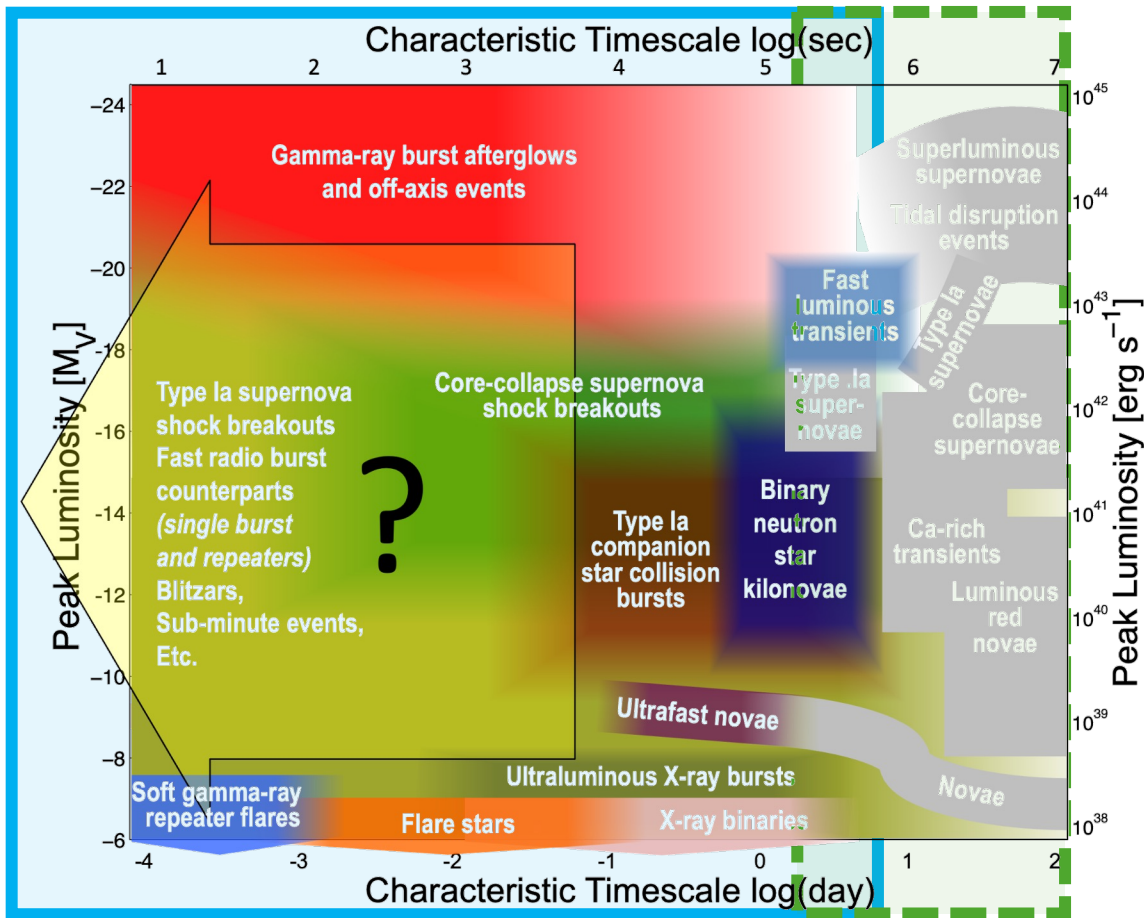


# Deeper, wider, faster program

Coordinated simultaneous deep, wide-field, fast-cadenced multiwavelength/messenger fast transient search & follow up

Fast transients (millisecond-to-hours duration) occurs at all wavelengths. Examples at optical wavelengths shown below.

Fast-cadenced observations on the same fields at the same time. 1-2 telescopes each wavelength regime + particles. Real-time data processing and transient IDs.



Telescopes wide-field coordinated simultaneously include - high-energy particles: Pierre Auger Observatory, HAWC; gamma-ray: H.E.S.S., Swift; X-ray: HXMT, Astrosat; UV: Astrosat; optical: CTIO DECam, Subaru HSC, KMTNet, AST3-2; mm/sub-mm: South Pole Telescope; radio: Parkes, ASKAP, MeerKAT, MWA; (also GW: LIGO/Virgo/KAGRA when online).

## DWF fast transients

Millisecond, seconds and minute cadence  
Multi-wavelength and particle observations

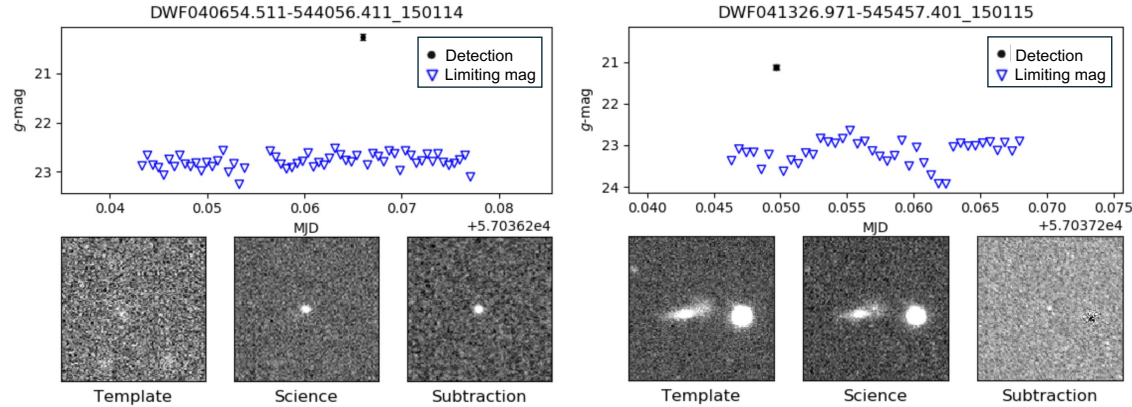
## Conventional transients

Day (or days) cadence  
Typically one wavelength

# Deeper, wider, faster program

Coordinated simultaneous deep, wide-field, fast-cadenced multiwavelength/messenger fast transient search & follow up

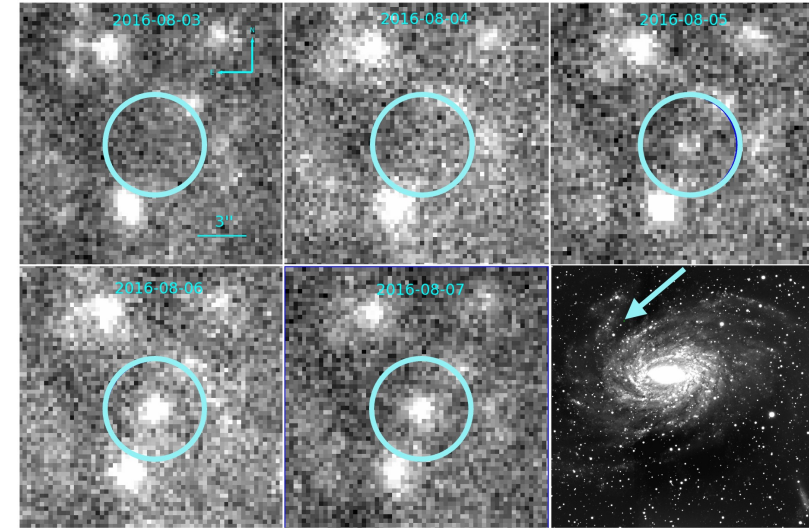
Example optical fast transients and potential new classes of events (Rare Gems) mined from the DWF data



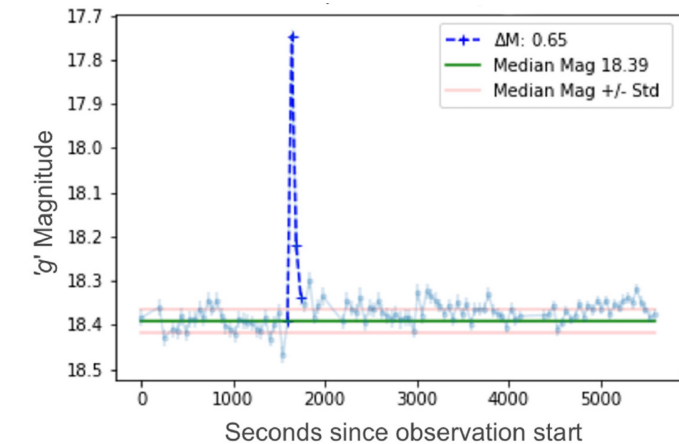
Multi-wavelength coverage of single-burst **fast radio bursts** sets limits (Zhang et al. 2024).

Multiple years detecting **extragalactic novae** is placing new rate constraints.

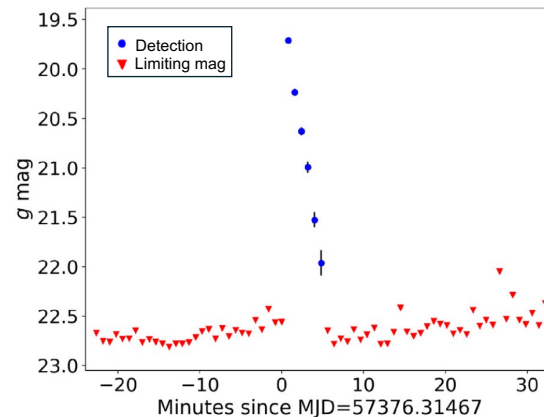
DWF catalogs the **faint transient and variable sky** ( $m \sim 23$ , *minute cadence*,  $m \sim 26$ , *nightly cadence*) that will help Rubin.



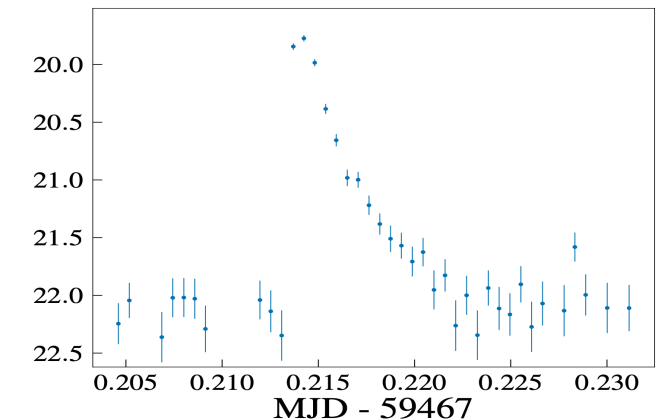
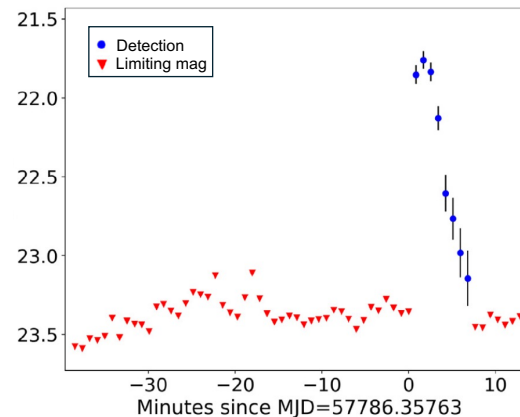
Two of the **sub-minute optical bursts** detected in the DWF dataset from machine-learning and human inspection (Goode et al. 2024).



**Flare stars** discovered with bursts as short as  $\sim 3$  minutes (Webb et al. 2021)



Example **fast extragalactic transient candidates**. DWF probes a faster-cadenced deeper regime than others (Andreoni et al. 2020)



**GRB orphan afterglow candidate** from a deep DWF search (Freeburn et al. 2024)