Transient phenomena and variations in comets, asteroids, centaurs and trans-Neptunian objects

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Focus of talk

- Restricted to *physical transient* phenomena
- Variations: mostly lightcurves

Some Reminders

- Small solar system objects are different: they move wsrpt to stars, the Sun, and the Earth.
- Visible only due to reflected light or scattered light and fluorescence (comets)

Some Definitions

- Comets: objects with elliptical (periodic) or hyperbolic orbits around the sun. Need to show activity somewhere in their orbit
- Asteroids: objects in the solar system that do not show activity. These objects are further distinguished dynamically:









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- Near-Earth objects (might contain dead comets) in Earth-crossing orbits (perihelia < 1.3 AU)
- Main belt asteroids (between Mars and Jupiter)
- Trojans around Jupiter, Neptune, and Mars (at respective Lagrange points)
- Centaurs, transient orbits with semi-major axis between 5.5 and 30.1 AU. Some show cometary activity
- Trans-Neptunian Objects: larger semi-major axis than Neptune







Calm and boring solar system?

- Solar system is dynamically pretty stable
- But there is still a lot going on!
- Example: Comet Shoemaker-Levy 9's impact into Jupiter in 1994. Once in a life-time event? 15 years later another impact into Jupiter!









Jupiter • July 23, 2009 Hubble Space Telescope Wide Field Camera 3



NAGA, ESA, H. Hammel (Space Science Institute), and the Jupiter Impact Team

What happened here?

- P/2010 A2 discovered by LINEAR Jan 6, 2010
- Small object about 150m in diameter in the main belt.
- Comet-like appearance but HST follow-up shows a very unusual pattern in the dust.



Possible explanations

- It is a main-belt comet (it has a comet designation) --> does not look like any coma or tail we have seen in comets.
- It is an asteroid collision --> need detailed modeling of the dust pattern to confirm this
- It is an asteroid/(dead) comet break-up: spin-up due to YORP or jets past the break-up limit --> dust pattern should be very different from collision.

More transient phenomena

- Unexpected outburst in comets (far from the sun e.g. Halley at 14 AU outbound), (immense increase in brightness in a short time, e.g. Holmes at about 2.5 AU, from mag 17 to about 3)
- Turn-on points for comets and centaurs (brightening in magnitude, exhibiting non-stellar appearance)
- Catastrophic disintegration of comets (e.g. C/ 1999 S4 LINEAR)
- Intermittent activity in centaurs
- Intermittent activity in main-belt comets

Why do we care?

- Interior structure
- Impact dynamics
- Evolution of small solar-system bodies
- Detailed understanding of solar system formation
- Death of comets/asteroids
- Frequency of transient phenomena

Observations

- Often serendipitous
- Dependent on limiting magnitude
- Unpredictable frequency
- Recognizable?

Variations

- Mostly rotation period
- Could also be due to albedo or color variations
- Why? Spin state is a basic parameter (often needed to explain other observations in comets)
- Distribution of spin states or rotation periods among and between the different solar system objects. --> Inference for internal structure, evolution of the objects.
- For comets: NPA paradox
- For comets: spin-up due to jet action (has been observed)
- For small asteroids spin-up due to YORP effect (has been observed)



From Pravec and Harris 2000 (Icarus, 12-20)

Spin states of other objects

- Periods of only about 25 comets known. Why? Comets are very faint (4% albedo) when far from the sun when inactive
- Need a lot more to populate the diagram as for asteroids.
- Spin states for TNOs even sparser. Only known for large TNOs. (Interesting object Haumea, fast rotator, cigar shaped).

How are spin states acquired?

- The old fashioned way: lightcurves covering several cycle of the entire rotation period at several different geometries ---> Laborious, time & telescope intensive.
- Some new methods: AO ---> less time intensive but needs large telescopes and objects.



Example: Spin Acceleration of Comet 9P/ Tempel 1



Observations needed

- Surveys to get initial determinations of a rough rotation period. How easy and feasible is this?
- Surveys feasible for detailed studies like for comet Tempel 1?
- A lot of observations covering the entire rotation period over a long period of time. Combination of lightcurves over time need to take geometry into account.