## **Galaxy Breakout**

- Key Science
- Requirements
- Preparatory Work
- Recommendations

# **Priority Science**

- BH Mass / Maggorian M-σ<sub>\*</sub> stellar orbits through astrometry (unlikely – faint!) spectroscopy (Vr unlikely do dispersion σ<sub>\*</sub>)
- Structure of inner AGN
- CMD of diverse populations
- Geometric Distances light echoes of GRB, SNe, AGN Baader-Wesselink (sp?), M31 rotations

## **Additional Science**

- Structure of nuclei distribution of stars, core clusters, etc.
- Microlensing what are MACHOs ?
- Surface Brightness Fluctuations
- Interferometric Deep Field ?

#### Requirements (most projects assume K band)

	Baseline	N dish	FoV*	Sens.	DynRange	Spec R	Time M
$M \sigma_*$	km	3-5	1 X	Check	No	2000	No
+evol.							
Structure	few km	5-10	1 X	Moderate	Yes high?	150, 2000	Yes, daily
of AGN				K 10-12			
CMD	Knut	Many	1 X several pointings	High	No (?)	10	No
Geomet. Distance	10-40 km	5-10	1 X	M31 – K~15	No	10 500(AGN)	Yes, weekly?
Structure of Nuclei			1 X				No
uLensing		3-5 (?)	1 X				Yes
SBF		MANY					No
IDF			few arcmin				No

\* Note: X = lambda/baseline x spectral resolution

## Preparatory Work (incomplete)

Structure of nuclei -

M  $\sigma_*$  - Theoretical work on brightness of cores.

Structure of inner AGN – Brightest / nearest with current facilities. More intense reverberation studies. Going to NIR.

CMD of diverse populations -

Microlensing -

Surface Brightness Fluctuations

Geometric Distances – Classical Novae. Studies of asymmetry in SNe and AGN (GRBs?)

Interferometric Deep Field -

### Recommendations