Report from Breakout Group on Exoplanet Searches

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Science Drivers

- Census of Exoplanets
- Habitable Planets
- Planet Characterization
- Direct Detection

Detecting Planets (Ground, O/IR)



- Methods
 - Transits (Timing)
 - Microlensing
 - RV
 - Astrometry
 - Direct Detection

Census: More of the Same



- Increase sample of planets
- 10 m/s on fainter stars
- Multiplexed RV

Census: Terrestrial and Below



Direct

Transits

Microlensing

RV

$$K \approx 10 \text{ cm s}^{-1} \left(\frac{M_p \sin i}{M_{\circ \acute{U}}}\right) \left(\frac{M_*}{M_{\circ}}\right)^{-1/2} \left(\frac{a}{\text{AU}}\right)^{-1/2}$$

Habitable Planets



- A stars -- SIM
- G stars -- Kepler
- Late G/K -- Optical PRV
- Late K/M -- NIR PRV

Direct Detection



- Young Systems
- Many planned instruments
- Path laid out
- Developing New Tech.
 - Broadening AODP to include high contrast imaging

Planet Characterization

1.3m telescope



(Winn et al. 2006)

Transit Follow-Up

- IR Photometry
- Astrometry
- Optical/NIR Spectroscopy
- Precision Photometry
 - Radii
 - Transit Timing
 - Moons/Rings

Planet Characterization



• 10⁻⁴ in 30-minutes



• High-precision, High Cadence on LT

Kepler Follow-up



(Gaudi & Winn 2006)

- PRV on LT/ELT
- Lots of Time!

Recommendations - I

Programmatic Suggestions

- Scheduling Flexibility
 - ToO, both routine and disruptive.
 - Non-traditional scheduling and access
- Intensive, Dedicated & Coordinated
 - More time for a census
 - Coordinated efforts to maximize science
 - 4-6m facility for RV planet-hunting

Recommendatations - II

Instrumental Capabilities

- "Diversity of Instrumentation"
- Near-IR Spectrometers for PRV
- High-speed, ultra-precise photometers

 QUOTA, ODI
- LT/ELT Polarimeters
- LT/ELT High-resolution Spectrometer

Recommendatations - III

Direct Imaging

- Path laid out
- Develop new technologies
- Broadening AODP to include high-contrast imaging

Astrometry

• Integrating Scientists with Experts