

SURVEY

The Dark Energy Survey Collaboration

~300 scientists US support from DOE+NSF

Fermilab, UIUC/NCSA, University of Chicago, LBNL, NOAO, University of Michigan, University of Pennsylvania, Argonne National Lab, Ohio State University, Santa-Cruz/SLAC/Stanford, Texas A&M





The Dark Energy Survey

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- DES uses 4 complementary techniques to measure acceleration of the Universe
 - I. Cluster Counts
 - II. Weak Lensing
 - III. Large-scale Structure (BAO) IV. Supernovae
- Two multiband imaging surveys: 5000 deg² grizY to 24th mag 30 deg² repeat griz (SNe)
- Built DECam, a 3 deg² FOV camera for the Blanco 4m telescope at CTIO

Survey 2013-2018 (525 nights) Facility instrument for astronomy community (DES uses 30% time).







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The Dark Energy Survey: Galaxy Cluster Counts

X-ray Cosmology in DES





The Dark Energy Survey: Cluster Counts

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4



- Spatially coherent shear pattern, ~1% distortion
- Radial distances depend on *expansion history* of Universe
- Foreground mass distribution depends on growth of structure



DES:Weak Lensing Shear Tomography







The Dark Energy Survey: BAO and Large-Scale Structure

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> Galaxy angular power spectrum in photo-z bins (relative to model without BAO)

> Photometric surveys provide angular measure

Radial modes require spectroscopy (MS-DESI)



Fosalba & Gaztanaga



DES: Type Ia Supernovae Light-curves





DES: Cosmological Goals Combined Probes





DES: Affiliated Surveys South Pole Telescope Survey(s)

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DES survey area encompasses SPT Sunyaev-Zel'dovich Cluster Survey 11



DES: Affiliated Surveys VISTA Hemisphere Survey





DES: Affiliated Surveys OZDES

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DES Timeline and DECam

- 2003 2007 planning
- 2008-2011 construction & testing
- 2011-2012 Installation at CTIO
- Sept. 2012 "1st Light"
- Oct. 2012 Feb. 2013 commissioning and "Science Verification"
- August 2013 Feb. 2014 "Year 1"
- Sept 2014 "Year 2" begins
- Sept 2015 "Year 3" begins



The Dark Energy Camera on the Blanco Telescope



Dark Energy Survey: Early Science





DES Early Results: Photometric Redshifts: Sanchez at al. 1406.4407





DES Early Results: Galaxy Clusters: Preliminary

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Red-sequence photo-z's are unbiased with ~1-2% scatter



Sensitivity: z=1.26 X-ray cluster is easy to see in DES-SV data



Miller et al., under DES review



DES Early Results: Cluster Weak Lensing: Melchior et al. 1405.4285





Cluster name	M_{200c}
RXC J2248.7-4431	$17.6^{+4.5}_{-4.0}$
1E 0657-56	$14.2^{+10.0}_{-6.1}$
SCSO J233227-535827	$10.0^{+3.7}_{-3.4}$
Abell 3261	$8.6^{+8.6}_{-3.9}$



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DES Early Results: Super-luminous Supernova: Papadopoulos et al 1501.07232



Radioactive decay of 56Ni? Or Magnetar?



DES Early Results: Synergy with VHS: Banerji et al 1407.3801



DES Early Results: Trans Neptunian Objects: Gerdes et al. in prep

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TNOs live in the Lagrange points at Neptune First discovered in 2001. Could be 10x greater in number than Jupiter Trojans Might help explain Solar System formation



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DES Early Results: Trans Neptunian Objects





•DES has just started!
•Science analysis of DES Science Verification data and Y1 data underway
• first results are coming out NOW
•Raw DES data available after 12 months via NOAO
•Planned public releases of processed/coadd
data/catalogs from 1st 2 seasons and from full survey