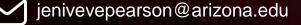
The Search for Failed Supernovae: Past, Present, and Future

Jeniveve Pearson

University of Arizona

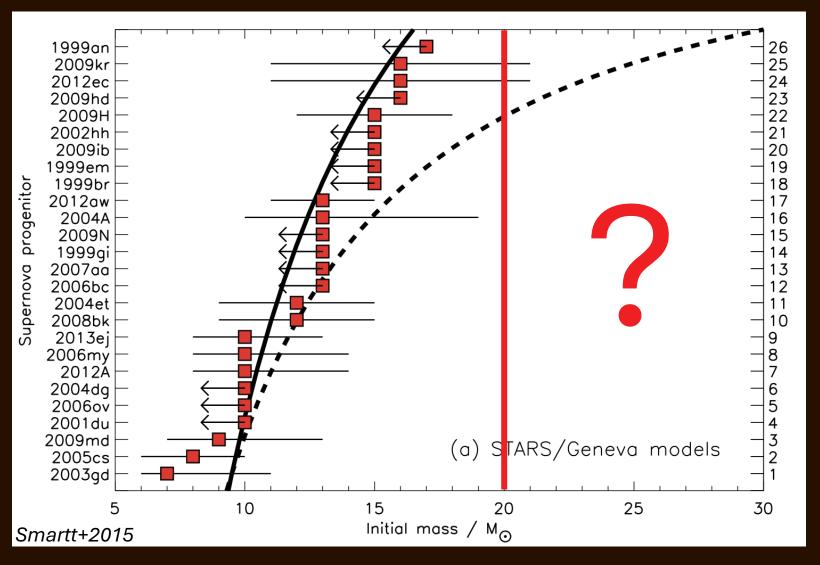
In collaboration with: David Sand, Jacob E. Jencson, Jay Strader, Jennifer E. Andrews, Nathan Smith, Emma Beasor, K. Azalee Bostroem, Griffin Hosseinzadeh, and Manisha Shrestha





Rare Gems in Big Data May 22nd, 2024

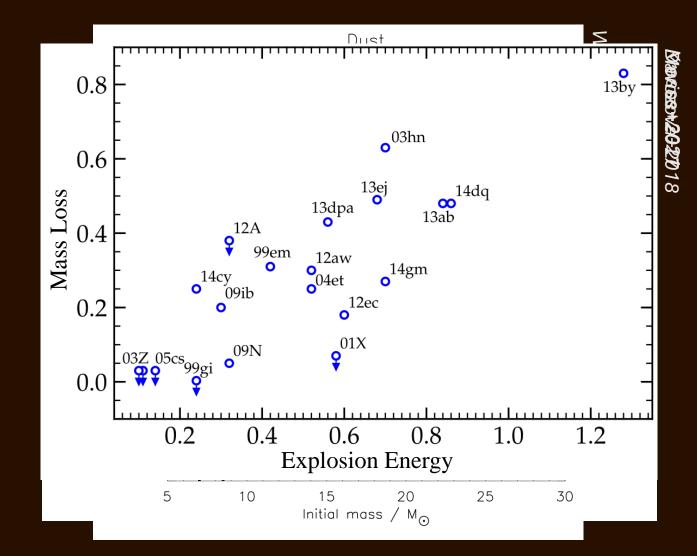
The Red Supergiant Problem?



- Red supergiants explode as type IIP/L supernovae
- RSGs have been observed up to 30 M_{\odot}
- No SNe II with progenitors >20 M_{\odot}
- Missing high mass RSGs?

Solutions to the Red Supergiant Problem

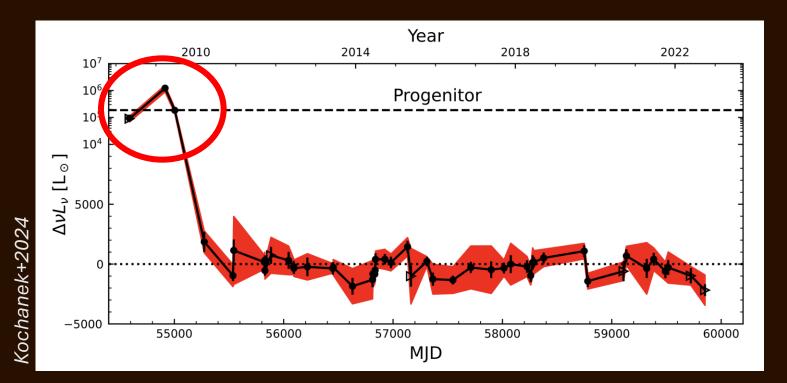
- Just haven't seen any explode yet
- The progenitor masses are incorrect
 - Dust attenuation
- Mass loss
- Direct implosion
 - Failed Supernova
 - Provides pathway for lowmass black hole formation

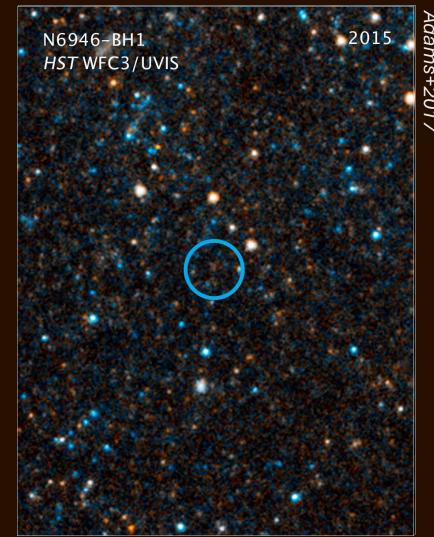


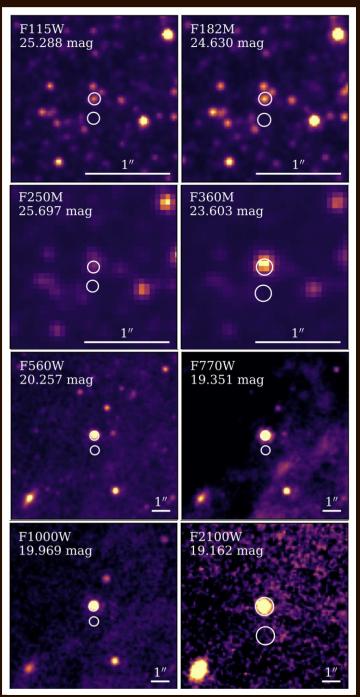
Finding Failed Supernovae

Three candidates

- NGC6946-BH1 is the most convincing
- Still hasn't brightened in the optical







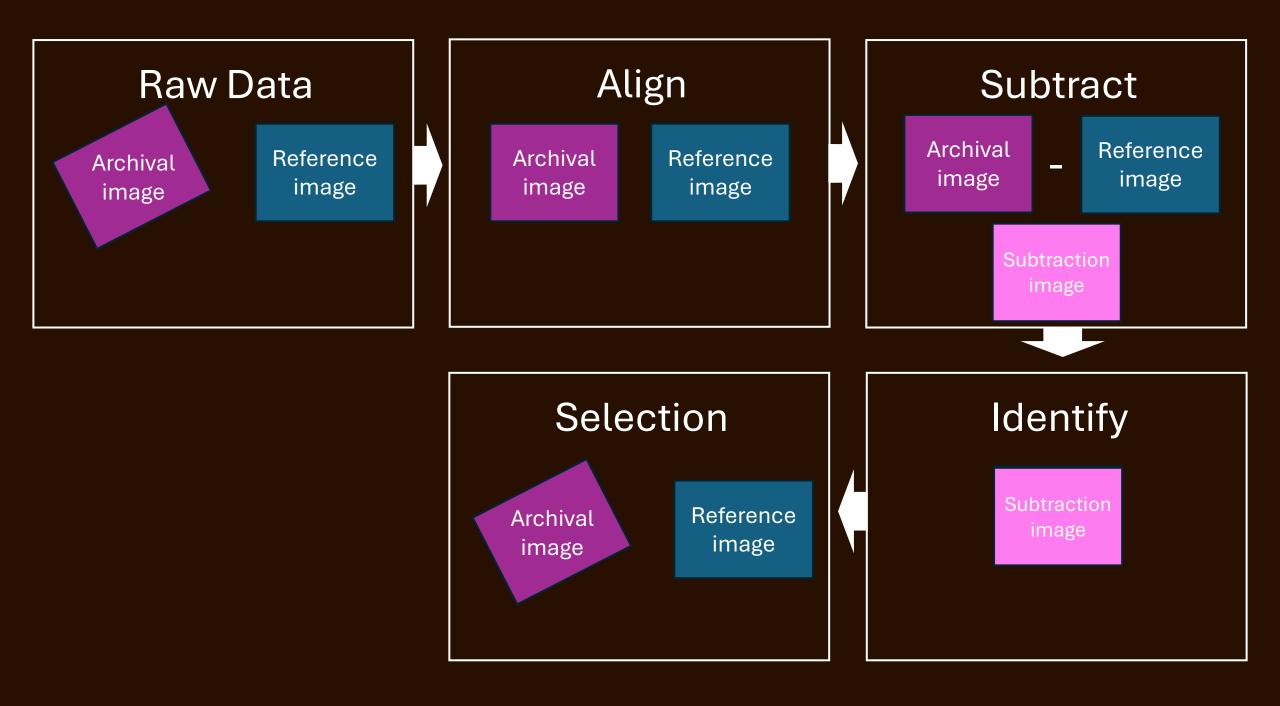
JWST Observations of N6946-BH1

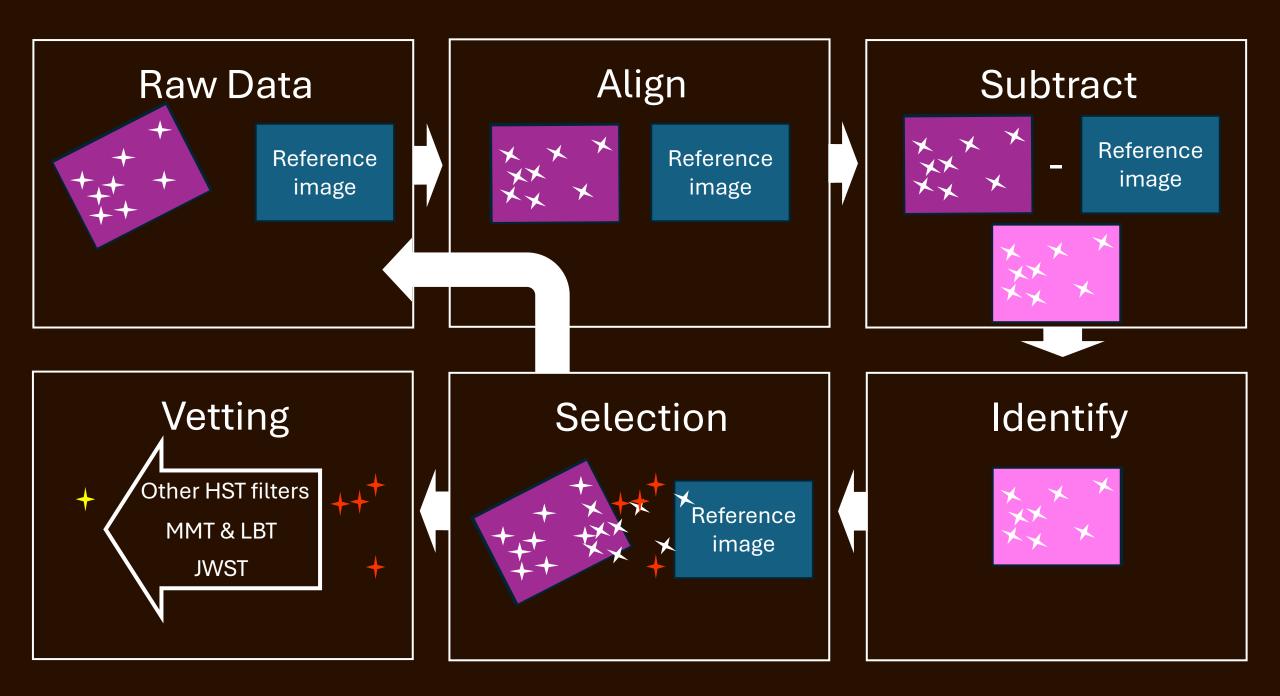
- If failed supernova \rightarrow no source in IR
- Source is clearly detected
- Possibilities:
 - Failed Supernova with accretion
 - ~15 years since disappearance
 - Outburst forming a dust shell
 - Fainter than pre-disappearance star
 - Merger event?
- Follow-up JWST spectra needed

The HST Failed Supernovae Project

- Observed a sample of 31 galaxies
 - High core-collapse supernovae rates (cycle 26; PI D. Sand)
- >2 epochs of F814W imaging per galaxy
 - Coverage in other HST filters
 - ~10 galaxies with JWST imaging
- 30+ year time baseline
- Based on current predictions, we expect to find ~5-20 new candidates

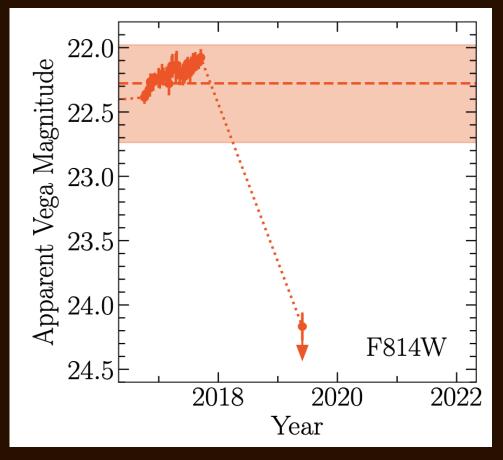


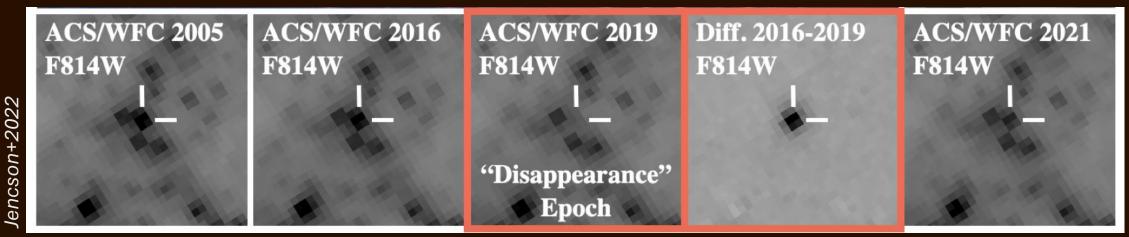


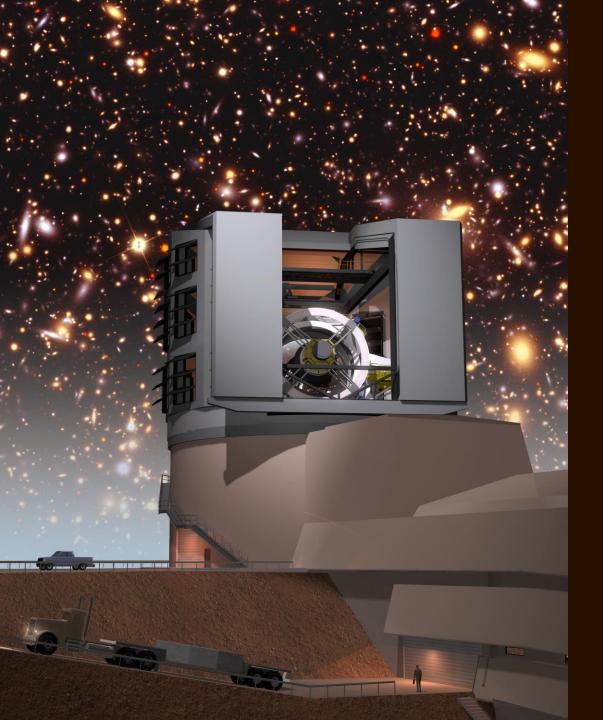


What We've Found

- M51-DS1 ("Disappearing Star 1")
- yellow or red supergiant that underwent an exceptional dimming in 2019
 - very massive (>20 $M_{\odot})$
- large mass-loss event
 - more extreme version of the "Great Dimming" of Betelgeuse







In the Era of LSST

- HST search will place upper limit on FSNe rate
- Detect faint outbursts
- In nearby galaxies (D≲15 Mpc):
 - Observe disappearing sources
 - Significant imposters
 - Cadence will allow for rebrightening checks
 - i, z, and y bands are ideal
- Need IR follow-up to confirm
 - JWST and Roman

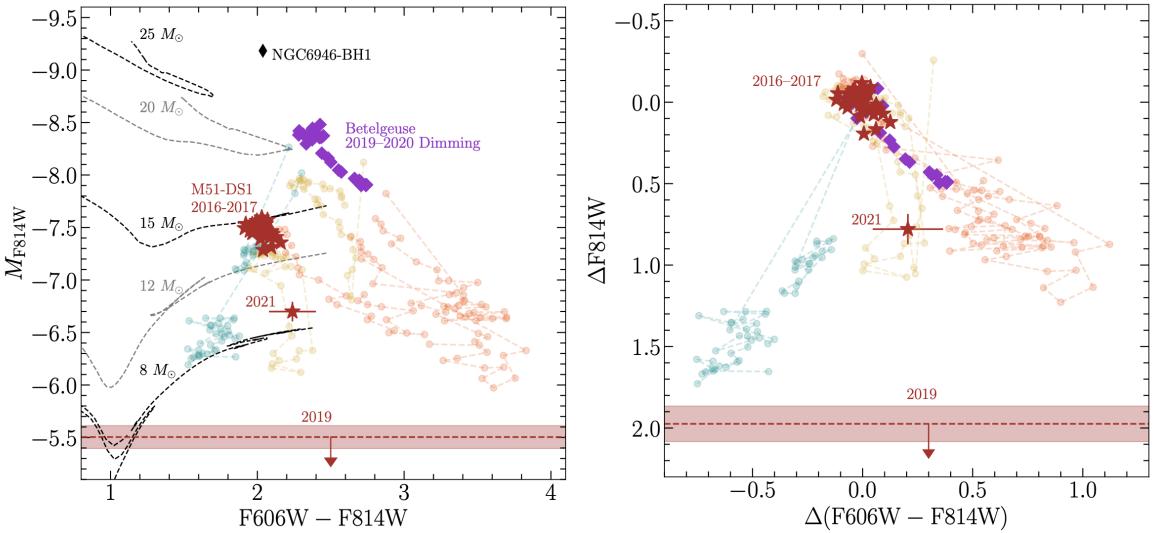
In Summary

- Failed supernovae are an enticing solution to the red supergiant problem
 - Not confirmed to exist
- HST archival data will provide strong upper limit
- Failed supernovae searches are great at uncovering variables
- LSST will detect faint transients and may find individual disappearing stars
- We don't know what failed supernovae look like so vetting will require human element



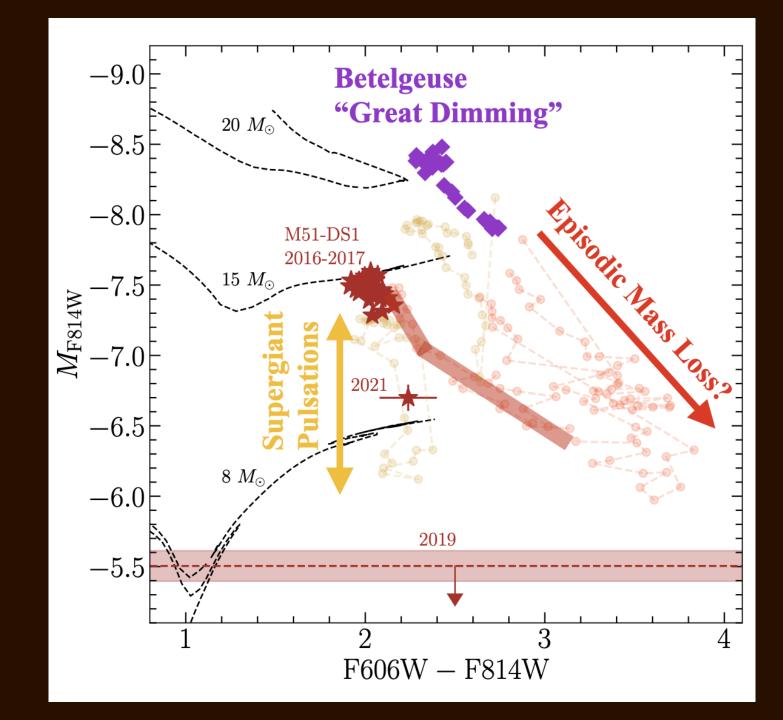
• Pros and cons of with doing this with HST

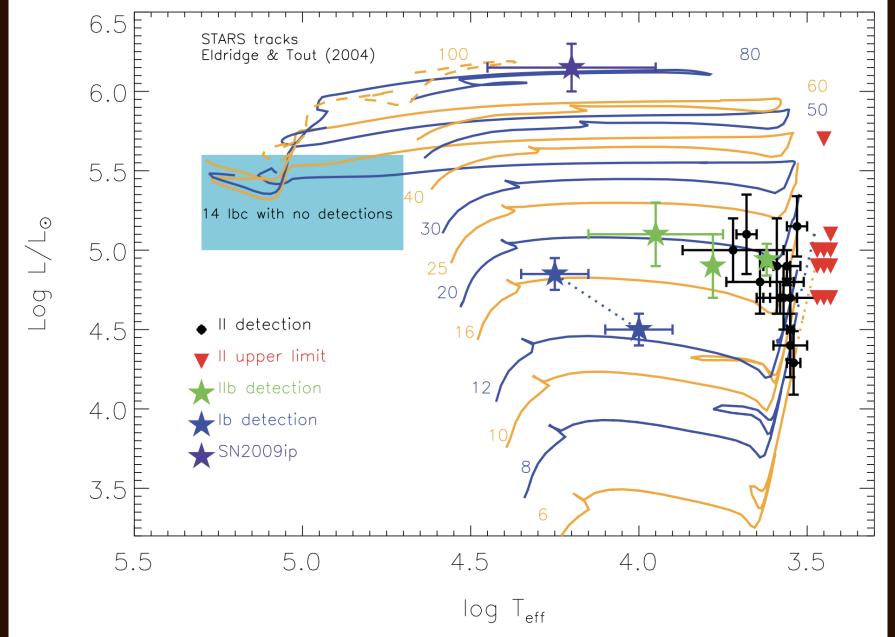
- Allows for resolving individual stars
- Challenges -> how do we do robust statistics to figure out a rate?
- Needs to be a targeted search
- Inhomogeneous sampling
 - Somewhat mitigated by additional archival data



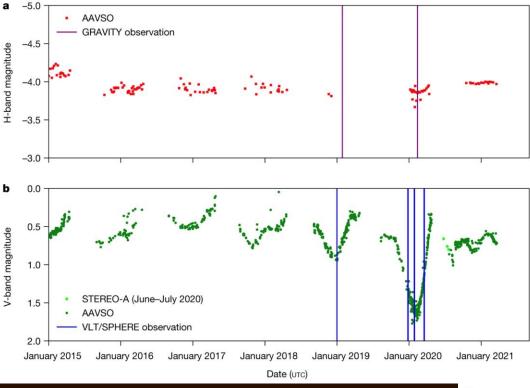
Jencson+2022

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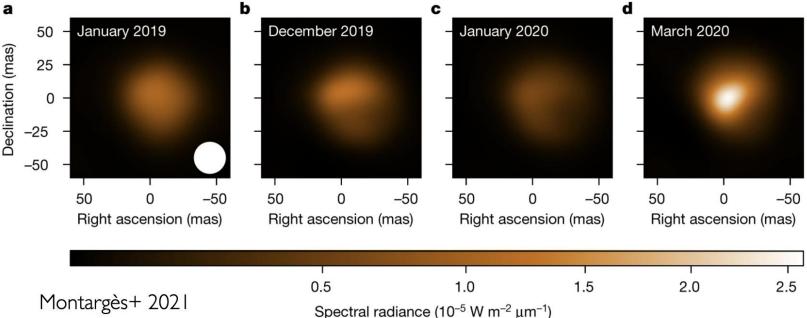
Smartt+2015



The Great Dimming of Betelgeuse

Favored model:

• Episode of mass loss that forms dust and obscures star. Coincides favorably with pulsation cycle and possibly the surfacing of a giant, cool convective cell.



The most common class of core-collapse supernovae, SNe IIP, are the result of The most common class of core-collapse supernovae, sive IIP, are the result of gravitational core collapse in red supergiants. Despite the existence of red supergiants with initial masses >18 solar masses, no SN IIP with a progenitor mass greater than ~18 solar masses has been observed. The apparent absence of high mass progenitors may indicate that massive red supergiants end their lives as failed supernovae, collapsing directly into black holes without producing a supernova. Proving the existence of failed supernovae, however, is a difficult observational challenge. Overcoming this requires frequent images of nearby galaxies to identify stars that disappear and remain absent in subsequent epochs. In this talk, I will discuss the ongoing effort to discover failed supernovae candidates in archival HST images of pearby star-forming galaxies. This sample enables strong observational images of nearby star-forming galaxies. This sample enables strong observational constraints on the frequency of failed SNe, sheds light on the origins of black holes, and assembles one of the most comprehensive datasets to date of extreme variables. Entering the era of LSST, this methodology will be vital for discovering the rare cases of failed supernovae and exceptional variable stars amongst the millions

of massive stars in nearby galaxies.