

### US National Gemini Office (NGO) Program Update Letizia Stanghellini NGO



NGO mission

- The US NGO supports Gemini users in the various phases of science program planning and execution, from proposal preparation to data analysis.
  - Support US users with telescope application
  - Support NOAO TAC with technical reviews as needed
  - Support TAC-approved programs
  - Liaise between Gemini and the US PIs in preparation for the ITAC to facilitate scheduling
  - Support post-data acquisition analysis
  - Inform the US community of Gemini opportunities



#### NGO supports post-data acquisition activities

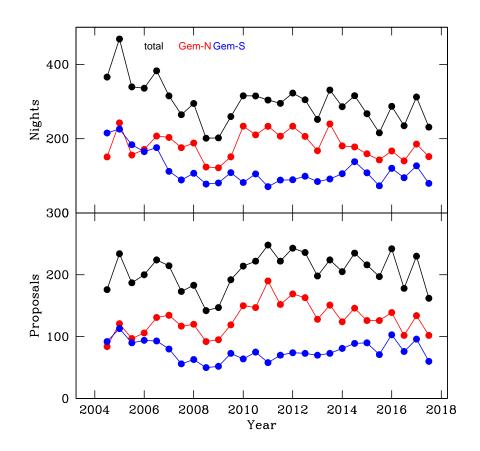
NGO does not provide Phase II support, which is done by Gemini

Recent and current activities

- Update the GMOS Data Reduction Manual
- Update the NGO webpages with community-wide news, tools, and information useful for US Gemini users (and all users)
- A draft MIRI data reduction manual is now available
- AAS Mini-workshop series continues. January 2017 *Mining Data Archives* has been successful (see extra slides, including Gemini plans)
- 2018 AAS Mini-workshop is being planned at this time
- TAC support
- Supporting Phoenix as Gemini-S visitor instrument in 2017A/B



#### Gemini submission to NOAO TAC



These numbers exclude NOAO Surveys, and LLP Gemini programs.

Submission has remained healthy at both telescopes.



The NGO webpage is a gateway for all Gemini users

It is updated bi-weekly

It includes links to all data analysis tools available to date

News Gemini Science Instruments and data reduction links · Overview - general information and useful links · OCTOCAM has been selected as the Gen4#3 instrument. It is . How to get your data - The Gemini Archive scheduled to be available at GS for LSST followup. The Feasibility · Gemini data format Study (205 pages) is available here. · Gemini data processing software · Hamamatsu CCDs have been installed at Gemini North. First use was March 26, 2017. Final characterization is underway. Users · GMOS with Hamamatsu data should keep in touch with their contact. NiRi scientist GNIRS · Phase I/II driving you crazy? Become proactive. A rewrite NIES project has started. Watch fo opportunities to submit input or send email to Bryan Miller at Gemini. · GRACES PROGRAMS PAY OVERHEAD CHARGES IN 20178. Details can GPI be found in the 20178 call for proposals. The PIT calculates • E2 the overhead charge. · GSAOI ITAC meets june 1. email concerning TAC decisions follows Phoenix by a few weeks. 2015 AAS Mini-workshop "IR data reduction" 2016 AAS Mini-workshop "AO data reduction" · 2017 AAS Mini-workshop "Mining data from the Public Archives" Proposal preparation sites Post-data acquisition sites Current NOAO Call for Proposals Gemini Science Archive Current Gemini Call for Proposals Gemini Publications Current Fast Turnaround Call Integration Time Calculator Gemini Helpdesk

## NGO webpage

#### **US National Gemini Office**



The US NGO main goal is to support US Gemini Users in the phases of the astronomical observing cycle, from proposal preparation through data analysis. This page is the main portal to all information needed in the different phases. explaining and connecting to specific Gemini web pages, where users can find definitive contents on instruments and modes. In addition, this page contains tools for data reduction and analysis for the entire Gemini community. All Gemini Users are encouraged to contact us with questions and comments.





Letizia Stanghellini Jane Price (Istanghellini at noao.edulorice at noao.edu) GMOS administration

NOAO Users Committee 2015 Tucson

Dave Bell at NOAO supports Phase I.

Who We Are

Verne Smith

**Gemini Office** 

(vsmith at noao.edu)

Head, US National

Ken Hinkle

(hinkle at noao.edu)

instruments, archive, Phase II, Phoenix PI

Current and retired

Dara Norman

GMOS

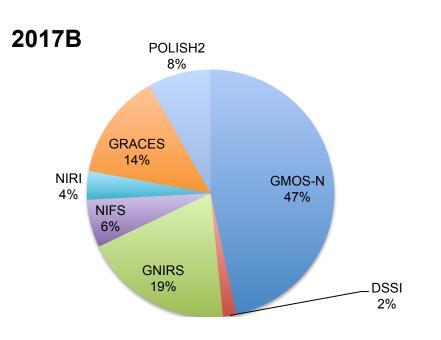
(dnorman\_at\_noao.edu)

Guide for Planning GMOS Observations

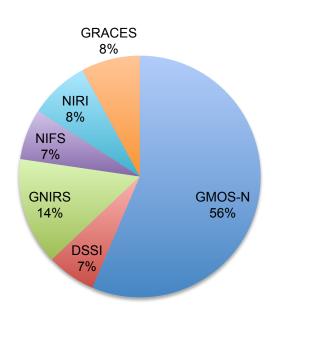


#### Gemini-N Instrument requests (by nights)

Regular NOAO TAC New programs only (excluding Subaru exchange, LLP, surveys, FT, DD)



2017A





#### Gemini-N Instrument requests (by nights)

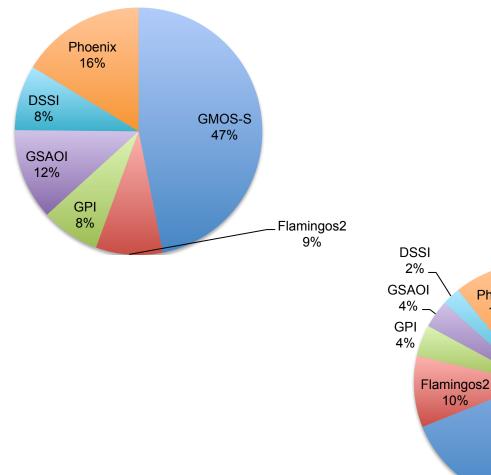
2017B

GMOS-S 69%

Phoenix

11%

2017A

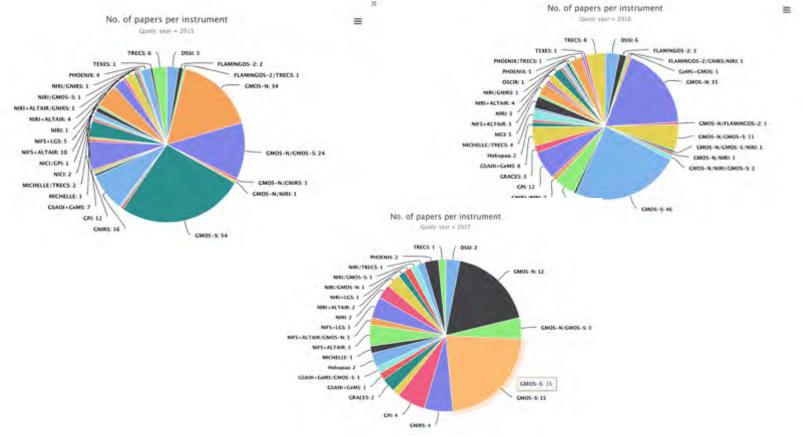


#### Regular NOAO TAC New programs only (excluding LLP, surveys, FT, DD)



# Publications from NOAO TAC allocated programs

- The number of US Gemini refereed publications has remained constant though the last few years (2015: 103; 2016: 90; 2017: 29 so far)
- Graph shows paper by instrument in 2015, 2016, 2017





#### **Backup slides**

#### 2017 AAS Mini-workshop on *Mining Data* Archives

 (1) Andy Adamson and Andre-Nicolas Chene: *Gemini Observatory Publications, Statistics, and Archive* (2) Harry Teplitz: *Astrophysics Archives at IPAC* (3) Scott Fleming: *Sailing the Archival Seas with MAST* (4) Knut Olsen: *The NOAO Data Lab Project*



# Gemini Observatory Publications, Statistics and Archive

## A Adamson

A-N Chené



#### Abstract

We describe the statistics of Gemini refereed publications, including relative productivity and impact of instruments and observing modes, and overall statistics such as the total publication count. We identify factors which may influence the probability of a publication emerging from a given observing program. At present, only a small fraction of publications arise purely from archival data. We present some of our plans for post-observing community support, and solicit input on various options for increased productivity in archival research.

Notes in italics on each slide are the comments added to the information in the course of the talk.



# Outline

Show some of the more interesting statistics relating to publications from Gemini observations - AJA

Show some of the ways in which we may be able to increase the publication rate, and solicit input on options - ANC





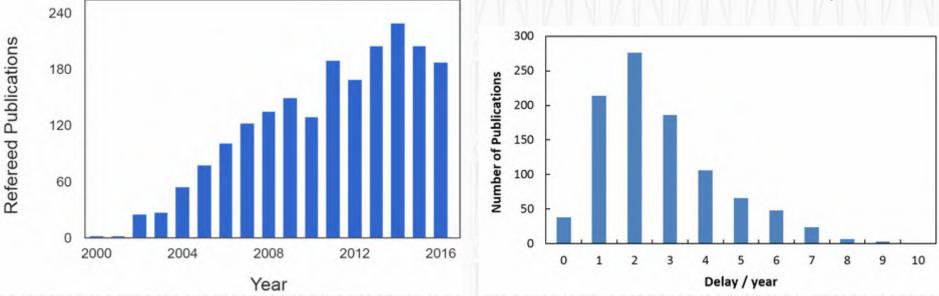
https://www.gemini.edu/library/gemini-publication-inclusion-criteria

## **Overall Statistics**

Exploring the Universe. GEA Sharing its Wonders Public/Media About/Contact Images Science Jobs clusien Search Search Results About 314 misuits (0.26 seconds) Gemini Publication Inclusion Criteria Gemini Publication Inclusion Criteria, by admin. To qualify as a Gemini publication, a paper must employ in an original way an image, spectrum, or data set

Publication count has now plateaued - about 200/yr, or a couple of papers per week at each telescope.

Lead time is about 2 years.



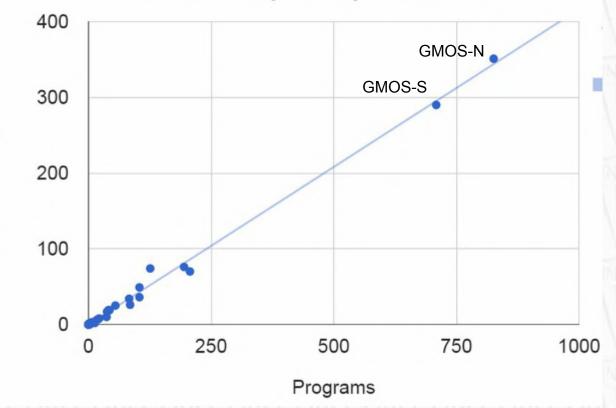
Noted in discussion: Gemini PIs across the whole partnership publish about two papers a week per telescope. Looks like we may now be at a plateau, and the influence of Large/Long programs is yet to be seen.



Published programs

## **Publications vs Instrument**

**Publication/Completion per Instrument** 



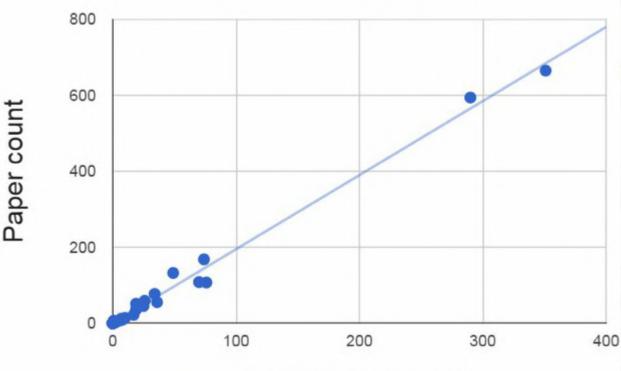
- >80% complete programs
- 2005-2014
- Each point is an instrument
- Single-instrument papers only

Noted in discussion: There's really great variation between instruments. All lie on the same line, with a slope of about 40-50% (meaning there's room to grow in terms of publication count).



## Multiple publications

Paper count vs. Published programs



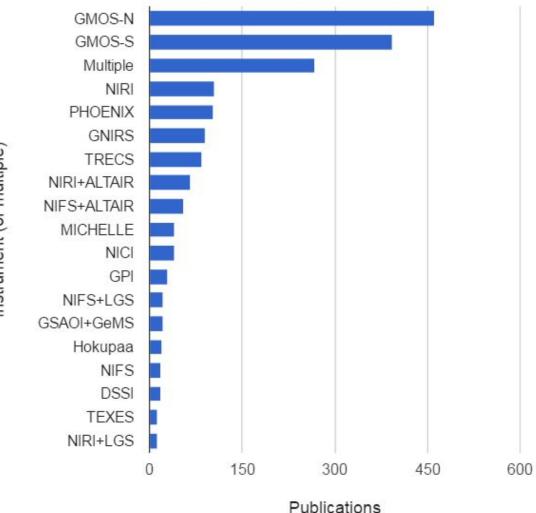
Published programs

- A program that publishes, tends to publish twice
- Again, instrument-invariant



## Publications by instrument

#### Publications per Instrument



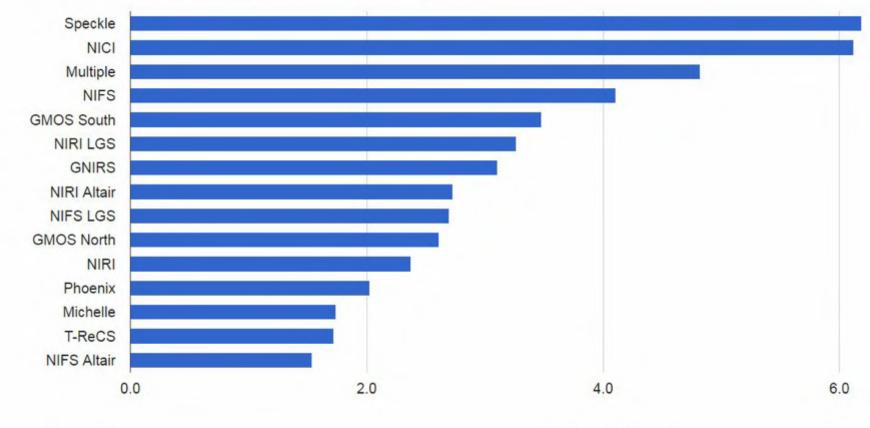
#### 2005-2014

Note added in discussion: GMOSs being at the top is not surprising (they dominate the program count). "Multiple" instrument publications coming just below that is more interesting, probably reflects the many possible combinations of more than one instrument, and programs over more than one semester contributing to a publication.

# Exploring the Universe, Sharing its Wonders

#### <Impact> per instrument

#### Average impact per instrument

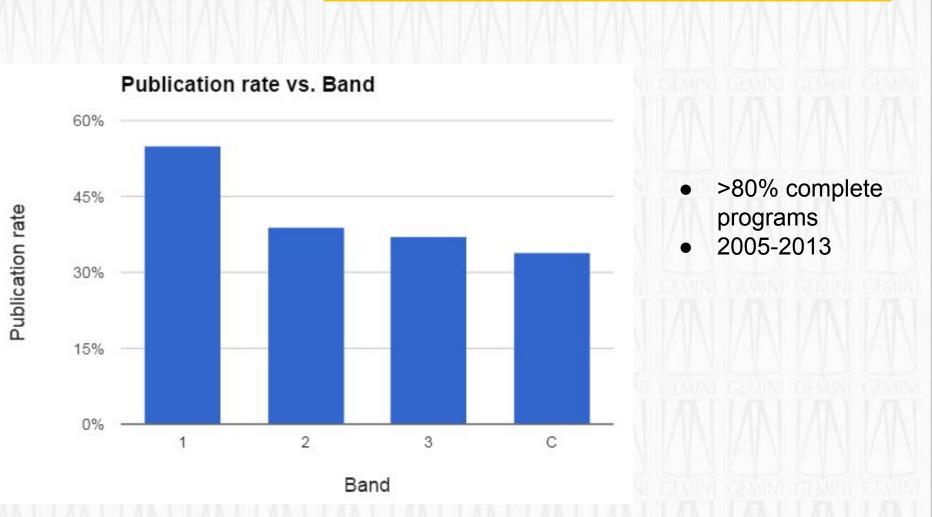


<impact>

Note in discussion: interesting that the cheapest instrument ever on the telescope has the highest current impact (but remember this is a snapshot). Also the fact that NICI - another exoplanet instrument - comes second shows that this chart is about subject fields, not instruments per se.



Publications vs band



Note added in discussion: since this is all to do with programs at the same minimum completion rate, the indication is that the TACs can spot a winner.



#### <Impact> vs band

#### Average impact per Band (single-program papers) 4.0 3.0 <impact> 2.0 1.0 0.0 Band 1 Band 2 Band 3 Band 4 Classical LLP Band

Single-program papers only, programs 2005-2015. Joints included.

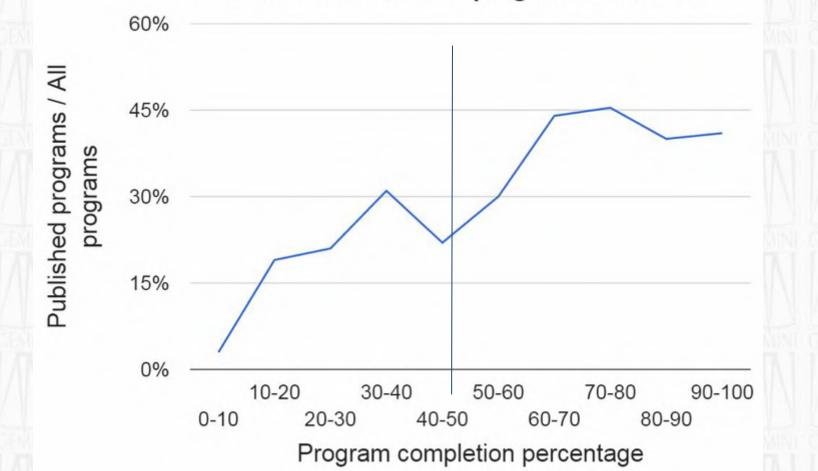
Impact data from 2000-2014 publications. Updated Oct 2016

Nett coverage: publications between 2005 and 2013.

*In discussion: within Band 1, Joint programs have by far the highest impact when published.* 

### **Publication vs completion**

Published fraction of all programs 2006-2013



Note in discussion: once a program gets above 60-70% complete, the chance that its data will appear in the literature does not increase. This point has been discussed by Gemini and its Governance and has affected queue execution policy to some extent.



Feedback from Users

# **Data Gathering (NGOs):** 2013-2015, with two-year lead time

#### Main Themes (consistent across partnership):

- Low priority data set <- longer programs</li>
- Circumstances changed <- fast turnaround</li>
- Postdoc left <- fast turnaround</li>
- Didn't detect the target

Note in discussion: the most typical responses were not what had been expected. Changes in our offered observing modes may address some of these, as indicated in red.



**Archive Publications** 

## Papers emerging from purely archival data amount to only a few percent of all Gemini publications.



# Gemini Science User Support, Data reduction and Archive possibilities

#### A-N Chené



Science User Support

# 4.6 FTE working on DR software 1 FTE focused on support

#### Avenues:

- Helpdesk
- Contact Scientists
- DR Forum
- Contacts page
  - Sus\_inquiries
  - NGO contact info



### **Our Archive**

Help - About - Misc Files

▲ achene -

#### So you have seen it at least once

#### https://archive.gemini.edu/

#### Semini Observatory Archive

			(leave blank for Any)	(Name of Target)
cave empty for Any)	(YYYYMMDD or start - end)	Resolver:	None ‡ R	esolve (Name Resolver)
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any :	(help)	Dec:	(leave blank for Am)	([+-]DD:MM:SS.ss or decimal degrees)
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Set at least one of the search criteria above to search for data. Mouse over the (text in brackets) to see more help for each item.



Reducing archive data

Two possible approaches:

• People can figure it out...

• Ain't nobody got time for that!!!



#### **Resources currently available**

Gemini reduction packages

- Gemini IRAF packages
- AstroConda!
  - set of astronomy-related packages for Anaconda
  - testing channel is now available
  - main Astroconda channel at STScI later this month

(One can contact James Turner from Gemini for any question about AstroConda.)

## **Reduction cookbook**

http://ast.noao.edu/sites/default/files/GMOS\_Cookbook/



GMOS Cookbook



But we could do more!!!

Publishing more cookbooks
 That is on the way (F2)

- Adding reduced data to archive
  - Ready for GRACES and GNIRS XDmode!
  - Reduced data provided by PIs
- Making "pipelines"
  - Would require a tremendous additional effort







# **Astrophysics Archives at IPAC**

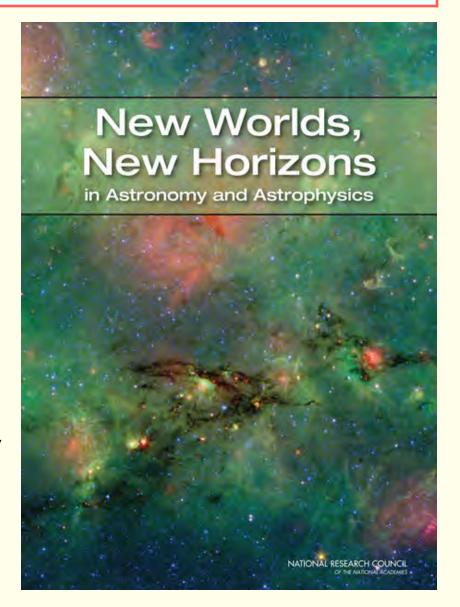
# Harry Teplitz

Mining Observatory Archives 1

#### **NASA's Commitment to Astrophysics Data Archives**

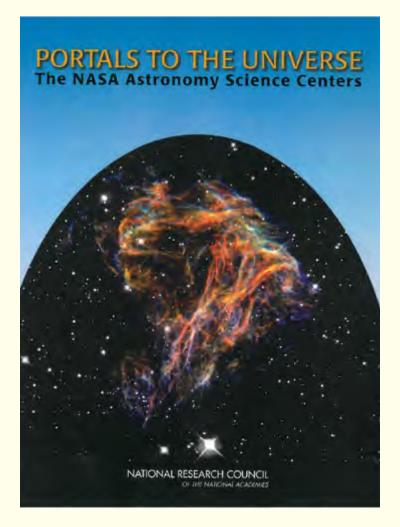
NASA has regarded data handling and archiving as an integral part of space missions."

"This support now provides the major return on the considerable investment the agency made... over the past 20 years."



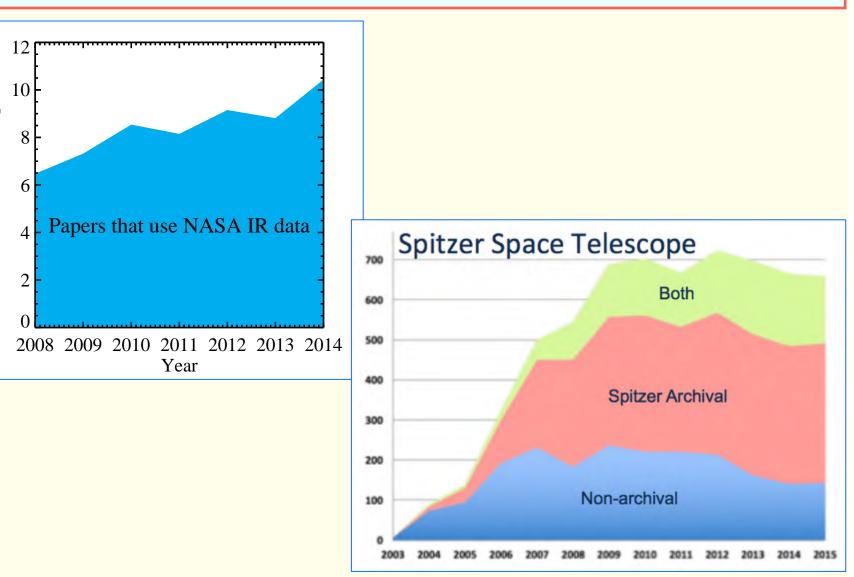
#### "A Sustainable Archive"

- Continually facilitates production of new scientific results
- Has a strategic goal to enable more and better science
- Contains high-quality, reliable data
- Provides simple and useful tools to a broad community
- Provides user support to the novice as well as to the power user
- Has many diverse uses (and users)
- Adapts and evolves in response to community input





#### **Archives Double the Number of Papers from the Observatory**



Mining Observatory Archives 4

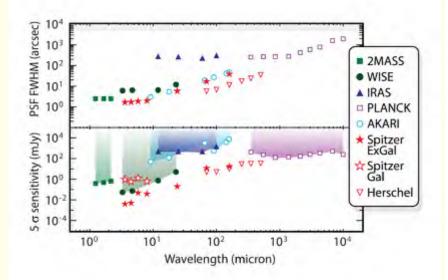
## **IRSA : NASA's IR/sub-mm archive**

# IRSA ensures the legacy of NASA's "golden age" of IR

- Enable research that has not yet been envisioned.
- Priorities set by missions and the community
- Support future flight missions

#### IRSA is continuing rapid expansion

- Since 2011, holdings more than doubled (now > 1 PB);
- # table rows increased by factor of 15 (>100 billion)
- Almost 40 million queries in 2016



- All-sky 20 photometric bands from 1 micron to 1 cm
- About 40% of approved ADAP programs involve analysis of IR data sets



### NED : managing complex data sets Overview – NED in a Nutshell Published:

Data are continually integrated from the literature, mission archives, and surveys ...

resulting in a comprehensive and current census of the universe

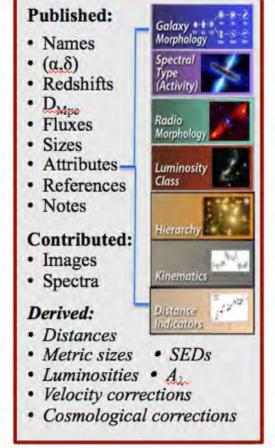
#### NED is where you find ...

- > Objects with z > 2.0 and available GALEX NUV flux
- Most precise z-independent distance measurement to M82
- > SED, spanning gamma-rays through radio, for quasar 3C 279

Lanse Sky

Spiral galaxies with stellar bars and Type 2 AGNs

Extragalactic papers have grown to 3,500 per year, with unique measurements for millions of objects

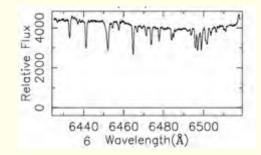


#### **The Keck Observatory Archive**





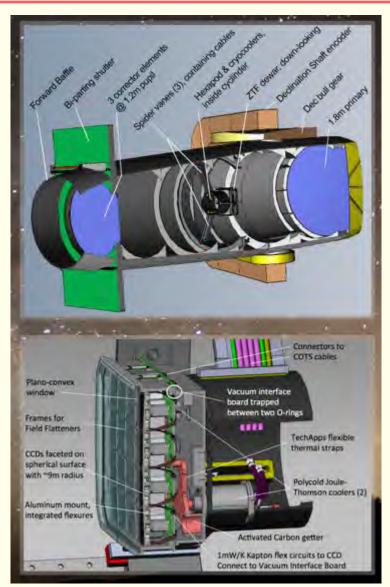
- NASA-funded collaboration between WMKO and IPAC/NExScI.
- Started with HIRES
- Systematic/automated capture of metadata ensures efficiency
- Now data from all ten instruments since their dates of commissioning
   *decommissioned instruments Summer 2015.*
  - proprietary period of at least 18 mo.
- KOA creates browse products for three instruments by automating pipelines.



KOA creates extracted HIRES browse spectra for every order of each object raw frame. Shown: T Tau. (PI: Reipurth).







- Palomar Observatory
  - Zwicky Transient Facility (2017+)
  - intermediate Palomar Transient Factory (iPTF; 2013-2016)
  - Palomar Transient Factory (2009-2012)
- Fully automated wide-field survey with 1.2 m Oschin telescope
- Publicly accessible survey data products available at IPAC
  - single frame exposures for selected regions of the sky,
  - source catalog files for those same regions.



#### An Archive's Job

#### Ingest new data

#### Maintain/serve vital repository of irreplaceable data

Support for observation planning

Resource for original science

High level science products

# Enable cutting-edge research API and Virtual Observatory

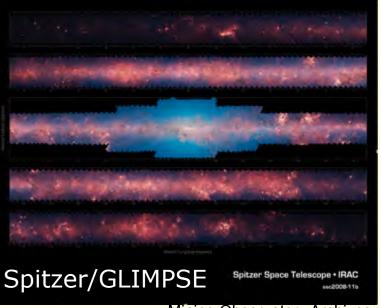
\* User support by experts

New/enhanced services

#### **High Level Science Products**

- Greatly enhance the science return of the archives Hubble Legacy HLSP are used 10x as much as typical pipeline products
- Make complex data sets accessible to a wider audience of researchers
- Expand the use of large, coherent projects
  - Herschel Key Projects
  - Spitzer Legacy and Exploration Science
- Generated by the community or by the archive





#### **Technical Synergy and Innovation**



Infrared Processing and Analysis Center

6

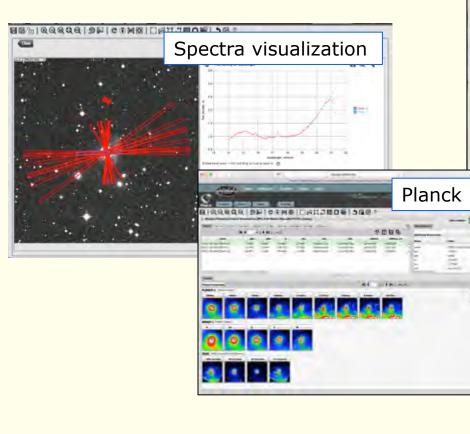
- IRSA implemented innovative indexing techniques for NEO/ WISE, optimized to meet the required use cases for database queries
  - single position spatial searches (using a recursively subdivided triangular mesh)
  - simultaneous matching of large user-supplied lists of positions (using a file-based index outside of the database).
- Challenges presented by WISE were used as opportunities to extend IPAC's capabilities
  - ZTF will require databases that are at least an order of magnitude larger

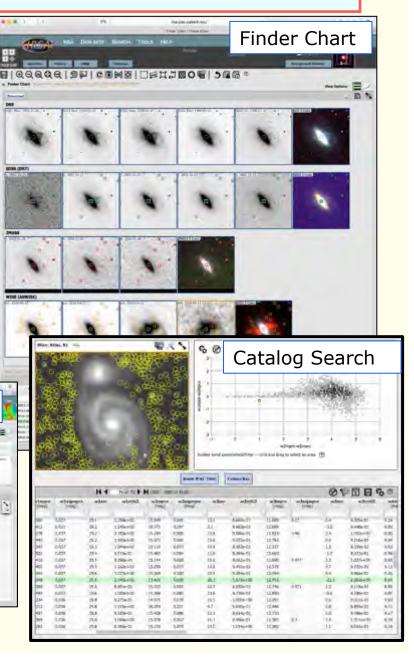
#### **Science User Support**

- Helpdesk some tickets are simple, others extremely complex
- Documentation
  - tools/data releases
  - updates in response to tickets
  - Handouts
- Demos
  - Live (AAS, ADASS, DPS)
  - Video tutorials (IRSA has > 60 videos)
- The complexity of Science User needs increases with time.

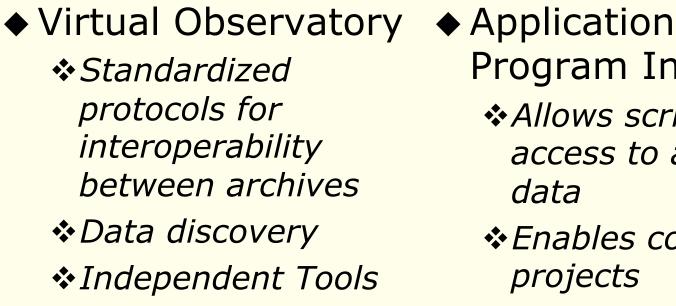
#### **Data Exploration and Visualization Services**

- Search & display can be tailored to various instrument/science contexts, using reusable visualization components
- Combine images, plots, tables, spectra
- Supports observation planning
- Firefly by IPAC





#### **VO broadens audience;** API supports diverse users

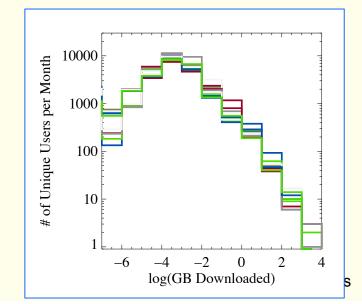


TOPCAT

**Program Interface** 

\*Allows scripted access to archive data

Enables complex projects



#### **Data Analysis Tools**

#### Analysis "near the data"

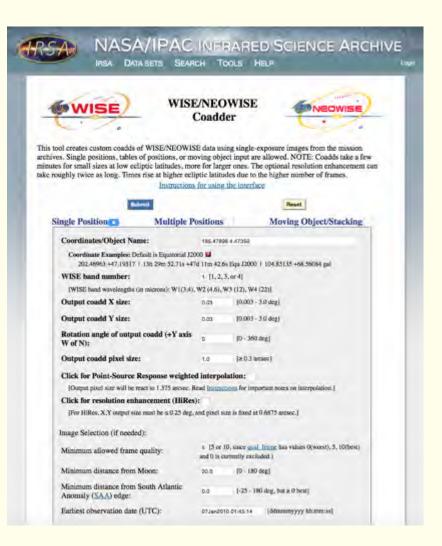
As data sets grow beyond local resources, researchers look to data centers to provide computing power and tools

#### Interactive tools

- ✤ WISE Coadder
- Planck map making
- ✤ IRAS tools
- \* Spitzer imaging and spectra

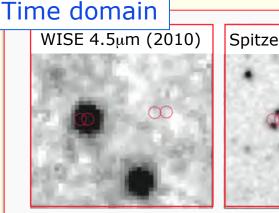
#### Looking Forward

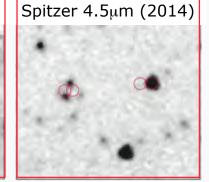
Big data means you can't move it all", suggesting the analysis must move to the data



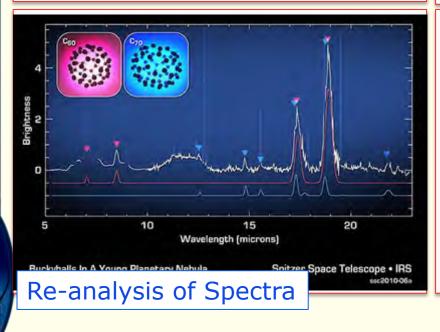
# **IR Science Highlights**





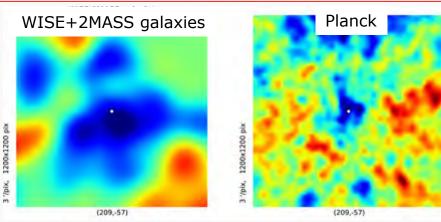


WISE+Spitzer discover the coldest brown dwarf (Luhman 2014, ApJL 786, L18)



Follow-up Observations

*WISE morphological study of Wolf-Rayet nebulae*, Toala et al. (A&A 2015, arXiv:1503.06878)



WISE+2MASS+PanSTARRS data may reveal supervoid in CMB cold spot seen by Planck; (Szapudi et al. 2015, MNRAS, 450, 288)

Combination of Surveys

#### Lessons Learned

 Long-term, stable archives greatly increase the return on observatory investment

Robust support for both expert and novice users pays off
 User support by instrument experts is crucial
 Standardization of tools within an archive increases

 Standardization of tools within an archive increases efficiency

Integrity of science data as obtained must be maintained
 Interoperability between archives benefits everyone

High level data products can expand the reach of large data sets

### Sailing the Archival Seas with MAST







# Scott W. Fleming STScl on behalf of the MAST team

# Outline – Mining Archival Data

- Overview of MAST missions
- The MAST Discovery Portal
  - Cone Search
  - Advanced Search
  - Cross-Matching
  - Access to the Virtual Observatory
- MAST-led Enhanced Data Products
  - gPhoton: GALEX photon events
  - HSC: The Hubble Source Catalog

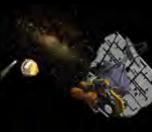
### MAST Missions: Past, Present and Future

#### Legacy Missions





Copernicus



EPOCh



EUVE



FUSE



GALEX



IUE



Kepler

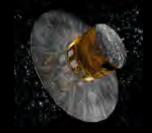


ORFEUS



VLA - FIRST

#### Active Missions



Gaia



Hubble



К2







Swift UVOT



XMM OM

# MAST Missions: Past, Present and Future

#### **Upcoming Missions**



TESS – Mar. 2018



James Webb – Oct. 2018



WFIRST – mid-2020s

# **Cross-Mission Discovery**

# MAST Portal: Cone Search

All MAST Observations About Collections			M60		andom Sear	ch Advanced Sean	Search db				
Upload Target List				U	ser Manual/H	Help I Leave Feedb	ack I About Th	nis Site			
lome Page 🖌 🔏 MAST: M60 👘 🖌 MAST: M60	3 81	MAST:	M60	MAST:	M60						AstroView
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ACS/WFC (30 of 30)	-	9			HST	ACS/WFC	HST	F475W	OPTICAL	NGC464	
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# MAST Portal: Advanced Search

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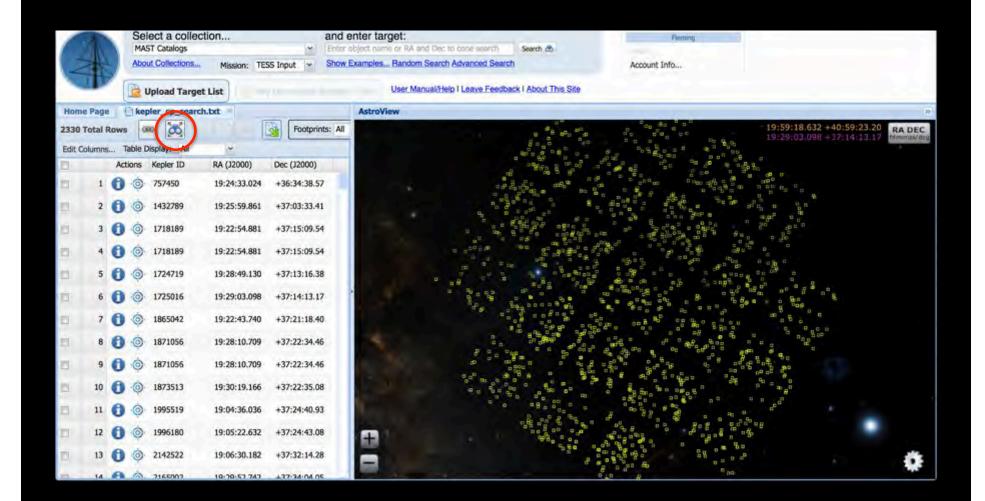
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#### MAST Portal: Access to the Virtual Observatory

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# Enhancing Data Products: gPhoton

#### gPhoton: A 1.1 Trillion Row Database of GALEX Photon Events

#### **GALEX** Overview

- Launched 28 April 2003
- FUV + NUV (simultaneously)
- FUV = 1350-1750 Å
- Phot. (5-6" resolution)
- Microchannel Plate Detector

Retired 28 June 2013 1.2 deg. diameter FoV

NUV = 1750-2750 Å

Spec. (R ~ 100-250)

77% of the Sky at Diff. Depths

#### gPhoton

- Databse of 1.1 trillion photon events
- Open-source Python software: create cal. images and light curves
- Photon events meas. with 5 millisecond precision
- Light curves with few second sampling for any GALEX image

gPhoton page: <u>https://archive.stsci.edu/prepds/gphoton/</u> Paper: Million, C; Fleming, S.W., Shiao, B., et al. 2016, *ApJ*, 833, 292

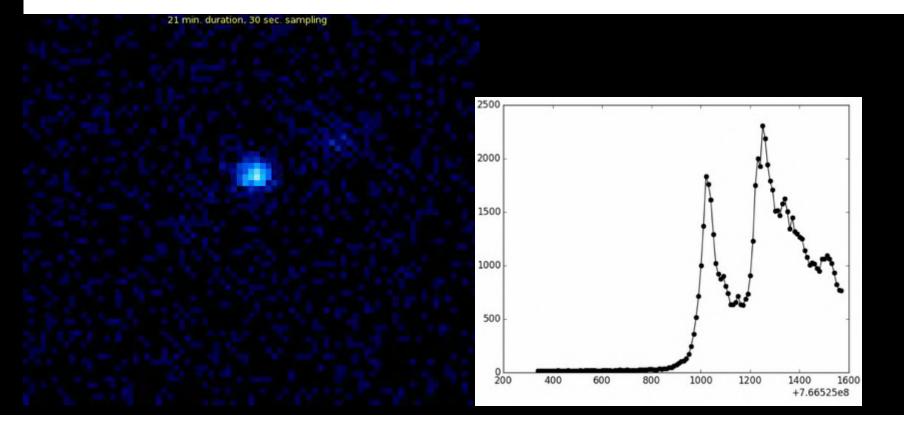
### gPhoton Design Considerations

#### import gPhoton

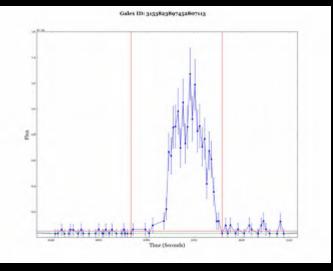
def main():

```
gPhoton.gMap(band='NUV', skypos=[176.91975, 0.25561], stepsz=30., skyrange=[0.0333, 0.0333], cntfile='gj_3685a_movie.fits')
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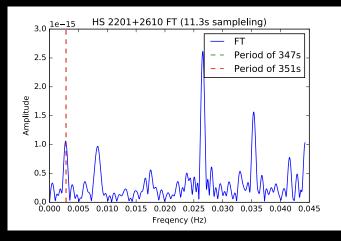
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if __name__ == '__main__':
    main()
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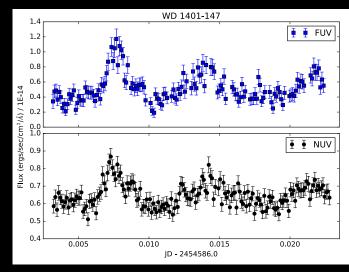
### gPhoton – Early Science Examples



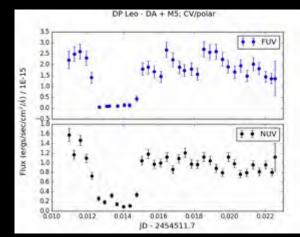
#### Stellar flares at low energies, w/ Rachel Osten and Clara Brasseur



sdB pulsation survey, with Thomas Boudreaux (2016 REU)



# White dwarf pulsations, with Michael Tucker (2015 REU)



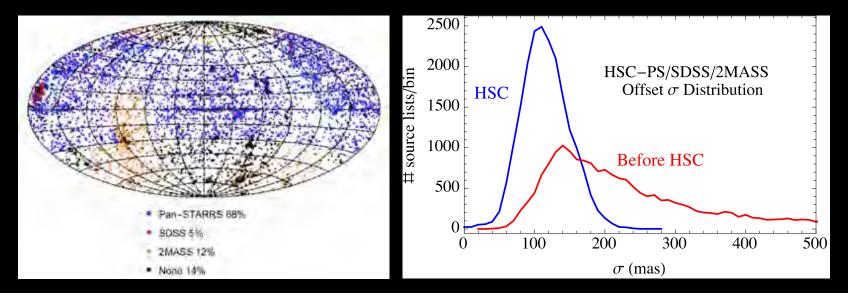
Eclipsing objects, including polars and WD exoplanets

### Enhancing Data Products: Hubble Source Catalog

# The Hubble Source Catalog (HSC)

# https://archive.stsci.edu/hst/hsc/ Paper: Whitmore, B.C, et al. 2016, AJ, 151, 134

- 1. Combines tens of thousands of SourceExtractor HLA source lists into a single master catalog. Uses matching algorithm from Budavari and Lubow 2012.
- 2. Includes WFPC2, ACS/WFC, and WFC3.
- 3. Absolute astrometry is good to ~100 mas (calibrated using PanSTARRS and 2MASS). This can eventually be improved to ~10 mas using Gaia observations).



### **HSC: Science Use Case**

# There are several Science Use Cases with step-by-step instructions and screen shots on the HSC webpage.

#### 2. Are there Use Cases available for the HSC?

Yes. We have a variety of Use Cases:

HSC Use Case #1 - Using the Discovery Portal to Query the HSC - (Stellar Photometry in M31 - Brown et al. 2009)

HSC Use Case #2 - Using CASJOBS to Query the HSC - (Globular Clusters in M87 and a Color Magnitude Diagram for the SMC)

HSC Use Case #3 - Using the Discovery Portal to search for Variable Objects in the HSC - (Time Variability in the dwarf irregular galaxy IC 1613)

HSC Use Case #4 - Using the Discovery Portal to perform cross-matching between an input catalog and the HSC - (Search for the Supernova 2005cs progenitor in the galaxy M51)

NOTE: This use case was made using version 1. However, most of the changes are relatively minor, hence it is still quite useful.

HSC Use Case #5 - Using the Discovery Portal and CasJobs to search for Outlier Objects in the HSC - (White dwarfs in the Globular Cluster M4)

HSC Use Case #6 - Using the Discovery Portal to study the Red Sequence in a Galaxy Cluster - (The Red Sequence in the Galaxy Cluster Abell 2390) NOTE: This use case was made using version 1. However, most of the changes are relatively minor, hence it is still quite useful.

HSC Use Case #7 - Comparing HSC "Sloan" filter magnitudes and SDSS magnitudes - (using the field around GRB110328A)< NOTE: This use case was made using version 1. However, most of the changes are relatively minor, hence it is still quite useful.

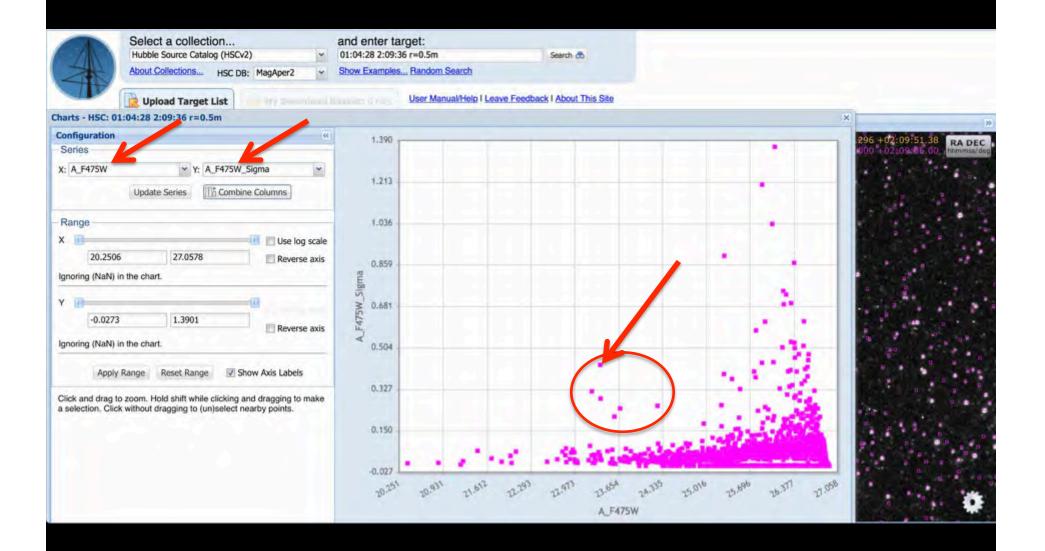
HSC Use Case #8 - Combining HSC magnitudes and HST spectra to Investigate Objects in the HSC (using objects in the LMC Cluster R136)

HSC Use Case #9 - Searching for Objects with both HST Imaging and Spectroscopic Data

#### HSC: Science Use Case, Variables in IC 1613 Select a collection ... and enter target: Hubble Source Catalog (HSCv2) 01:04:28 2:09:36 r=0.5m Search 🚓 About Collections... HSC DB: MagAper2 Show Examples... Random Search User Manual/Help I Leave Feedback I About This Site 😥 Upload Target List Home Page HSC: 01:04:28 2:09:36 r=0.5m HSC Matches: 87416231 AstroView Footprints: All 2416 Total Rows of Objects Equatorial Coord 01:04:28 +2:09:36 J2000, radius: 0.00833° 690 RA DEC Filters List View -00 Edit Columns... Table Display: All Clear Filters Edit Filters... Help... Match ID Distance Match RA Match Dec Actions - Keyword/Text Filter E 87416231 0.0209300446210454 01:04:27.958 +02:09:34. Filter All Columns ×Р 13 87418391 0.0227336760824004 01:04:27.916 +02:09:35. Match ID 87416241 01:04:27.897 +02:09:37. 0.0379670283468321 87416171 0.0412776078346797 01:04:27.949 +02:09:33. 1 87428772 0.042604457732627 01:04:27.861 +02:09:34. 1 87415845 0.0441243272632123 01:04:27.973 +02:09:33. 87411055 87595960 87418399 0.0454911604334835 01:04:27.866 +02:09:34. Distance 87416631 0.0459757290961777 01:04:27.910 +02:09:33. 0 a marile day bare and by blacketters 0 87416173 0.0460838457631922 01:04:27.943 +02:09:33. 1 10 87595655 0.0505050289109621 01:04:27.816 +02:09:34. 10 11 87415859 0.0530502911429743 01:04:27.894 +02:09:33. 0.0209 0.5 門 87417121 0.0553675651103514 01:04:27.809 +02:09:37. 12 Long. 47 A 0350555A .03.00.35 01.04.77

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# HSC: Science Use Case, Variables in IC 1613



# HSC: Science Use Case, Variables in IC 1613

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# Summary

- MAST supports more than 20 missions: cross-mission discoverability is a key driver in our development efforts.
- The MAST Portal allows for searching across MAST missions and catalogs, the Virtual Observatory, and CDS catalogs.
- Beyond archving mission data, MAST also creates enhanced data products that enable new science. Recent examples are:
  - gPhoton: A time-tagged databse of 1.1 trillion GALEX photon events.
  - HSC: A master catalog of unique sources across all HST observations.
- Newly created STScI Data Mission Office, led by Arfon Smith, will begin to expand MAST capabilities into modern technological arenas, including cloud storage, "bring-code-to-the-data" environments, improved access to high performance computing resources, better support for API access to MAST resources.





# The NOAO Data Lab Project Introduction

## Knut Olsen for the Data Lab team



1

# Data Lab Team



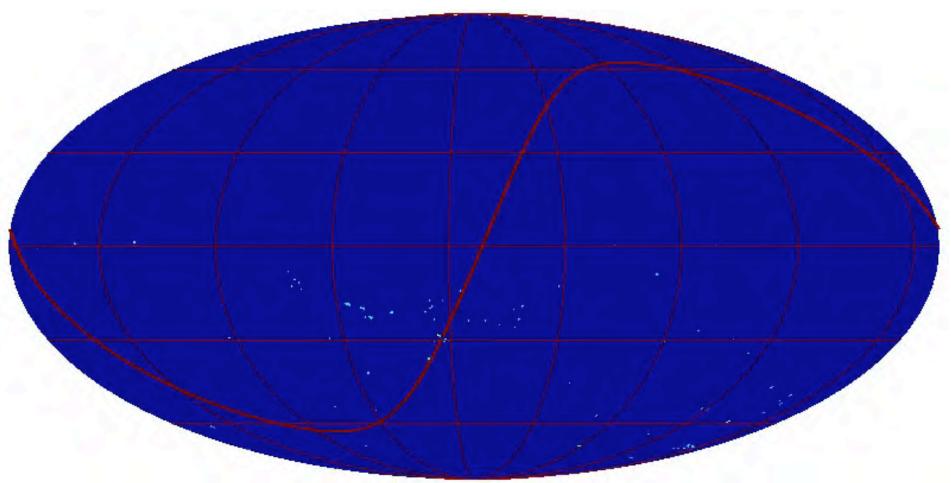
Current team:

- Mike Fitzpatrick, Lead Developer
- Matthew Graham, Scientist/Developer
- •Wendy Huang, Software Engineer
- •Stephanie Juneau, Data Scientist
- David Nidever, Data Scientist
- Robert Nikutta, Data Scientist
- •Pat Norris, Test Engineer
- •Knut Olsen, Project Scientist
- Steve Ridgway, Scientist
- •Pete Wargo, System Administrator

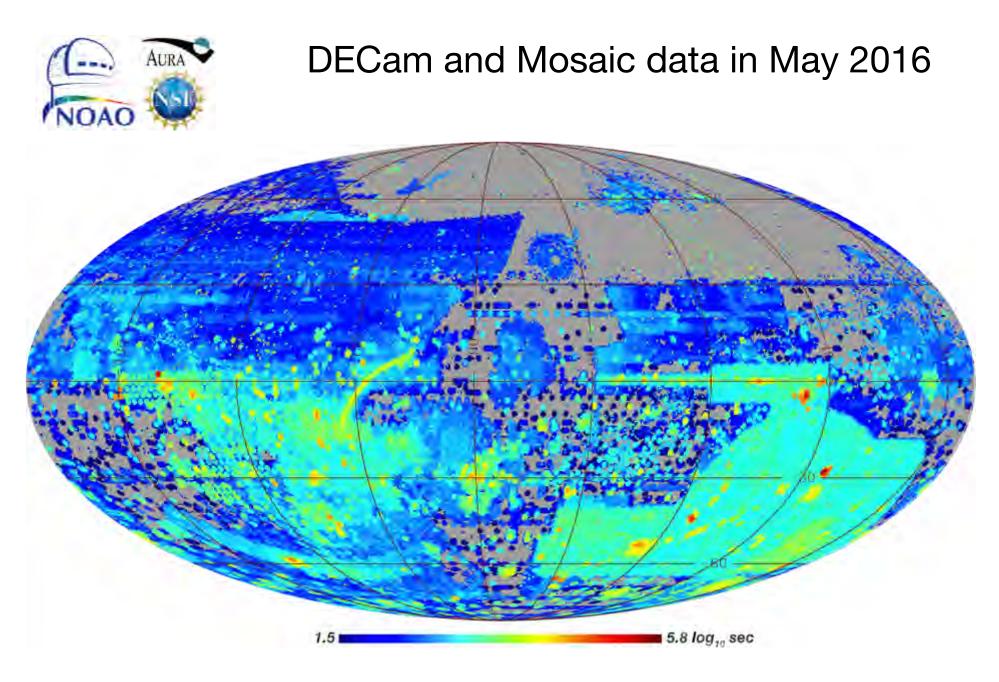




# NOAO wide field imaging data over time











500 TB (January 2017) of on-target imaging data ( $t_{exp}$ >30s) currently from:

- Dark Energy Survey
- Legacy Surveys for DESI Targeting
- Community DECam and Mosaic programs and surveys Hundreds of TB more coming
   Total holdings of several PB

Large catalogs coming:

- Dark Energy Survey 45 TB
- DESI Targeting Survey ~5 TB
- Community programs and surveys up to several TB each



# NOAO Data Lab



- Goal:
- Efficient exploration and analysis of the large datasets being generated by instruments on NOAO wide-field 4-m telescopes
- Approach:
  - Catalogs and images linked to catalog objects
  - Data discovery
  - Developing intuition through interaction with selected catalog and image set of known objects
  - Automation of analysis to aid discovery of unknown objects





## Data Lab in a Nutshell

Large Catalogs – Data Lab will serve TB-scale databases

- Pixel Data Data Lab will connect users to images and spectra in NOAO Science Archive
- Virtual Storage Minimizes data transfer
- Visualization Data Lab will enable data exploration
- Compute Processing Data Lab will allow workflows to run close to the data
- Additional features Access to published datasets and external data services, data publication, exportable workflows, distributable software



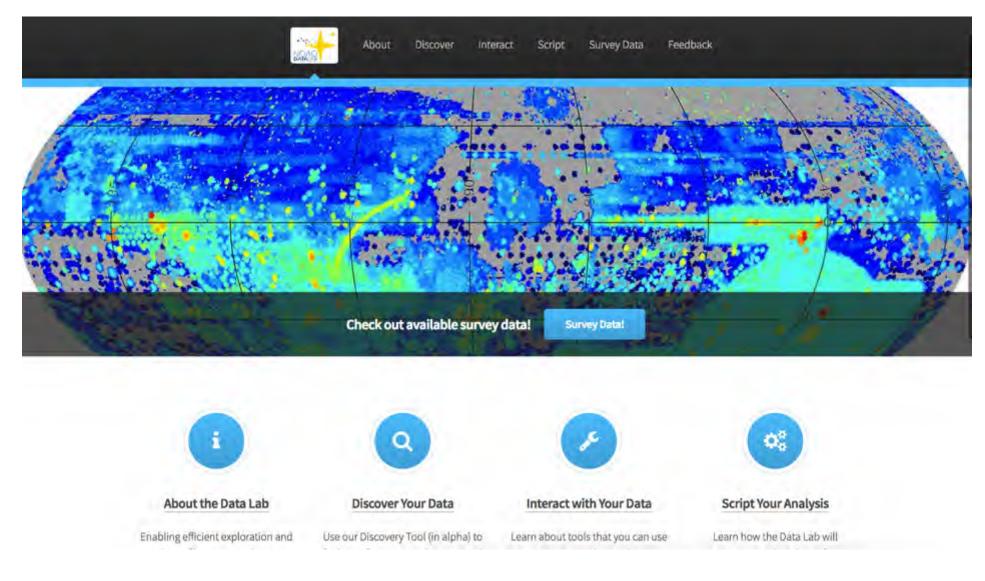
# Timeline



- March 2015: Conceptual Design Review
  - Lisa Storrie-Lombardi (Chair), Severin Gaudet, Zeljko Ivezic, Connie Rockosi, Beth Willman reviewed Science Case & Requirements, System Architecture, Operations Concept & Requirements, and Schedule
- Fall 2015 hiring campaign
- June 2016 San Diego AAS Demo
- August 2016 Interim Review
  - Lisa Storrie-Lombardi (Chair), Severin Gaudet, Zeljko Ivezic, Ed Olszewski, Beth Willman, and Dennis Zaritsky reviewed progress and Year 2 plan
- January 2017 AAS SMASH DR1 and DECaLS DR3
- Summer 2017 first public release
- End 2017/Early 2018 DES DR1

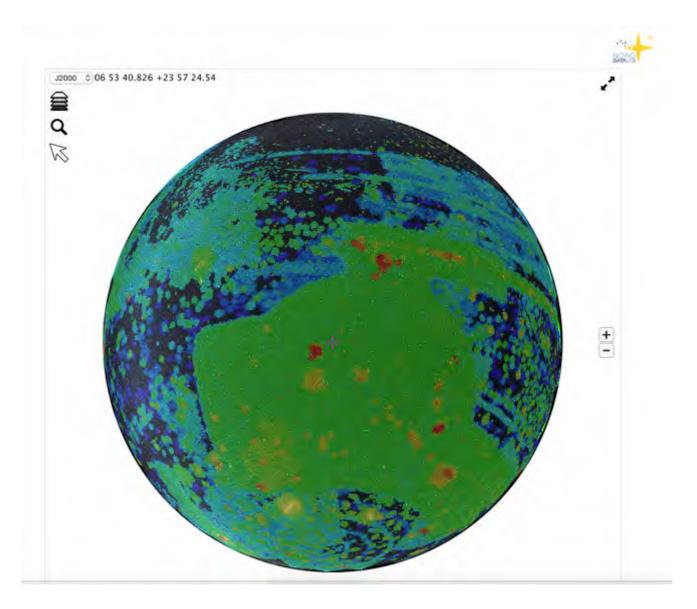


## datalab.noao.edu



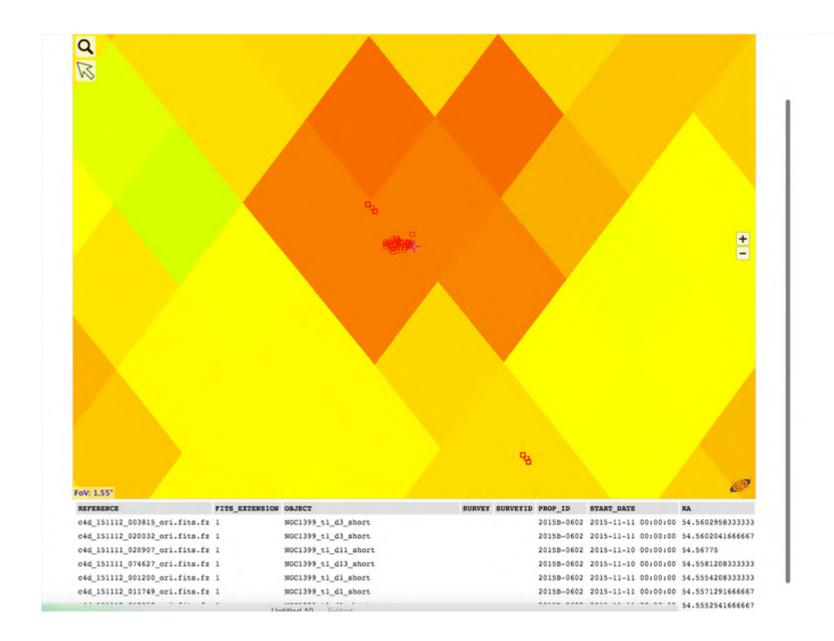


# Data discovery





### Data discovery





### Survey data



About Discover Interact Script

ript Survey Data

y Data Feedback

### SMASH

### Survey Coverage

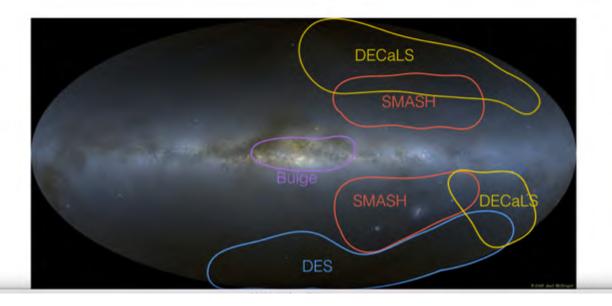
#### DECaLS

DES (coming soon)

Bulge

(coming soon)

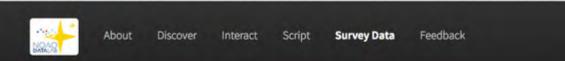
The map below shows the areas covered by surveys with catalog data currently available through the NOAO Data Lab (SMASH, DECaLS) and those expected within approximately a year (DES, Bulge). Hover over an outline to see the survey name or click on an outline or on the sidebar links to go to the page for that survey. For pure image data from these and other observing programs, visit the NOAO Science Archive.





### SMASH DR1

### Survey of the Magellanic Stellar History (SMASH)



#### Description

- Overview
- Goals
- First Data Release
- Data Reduction and Calibration

#### **Data Access**

Analysis

Explore

Results



### The SMASH Survey

### Overview

The Survey of the Magellanic Stellar History (SMASH) is using DECam to map 480 square degrees of sky to depths of ugriz~24 with the goal of identifying broadly distributed, low surface brightness stellar populations associated with the stellar halos and tidal debris of the Magellanic Clouds. It will eventually contain measurements of approximately 250 million objects distributed in discrete fields spanning an area of about 2400 square degrees. The first data release (DR1) contains ~100 million objects from 61 observed fields. Browse these pages to learn more about SMASH and to access the data. The <u>SMASH</u> overview paper (Nidever et al. 2017) describes the survey in detail, including its goals, survey strategy, reduction, and calibration.

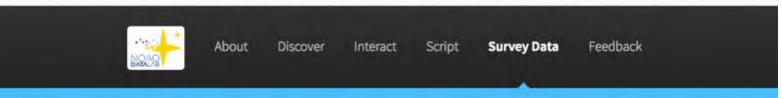


## SMASH DR1 Data Access

Description	Data Access
Data Access	The SMASH data are accessible by a variety of means:
Analysis	
Explore	Data Lab Table Access Protocol (TAP) service TAP provides a convenient access layer to the SMASH catalog database. TAP-aware clients (such as <u>TOPCAT</u> ) can point to <i>http://datalab.noao.edu/tap</i> , select the <i>smash_dr1</i> database, and see the database tables and descriptions.
Results	smash_dr1 contains six tables: chip, exposure, field, object, source, and xmatch. These are described in the schema page.
	Data Lab Query Manager
	The Query Manager is available as part of the prototype Data Lab software distribution. The Query Manager client provides a Python API to Data Lab database services. For the SMASH DR1 release, these services include only anonymous access through synchronous queries of the catalog made directly to the database. The full public release of the Data Lab Query Manager in the summer of 2017 will include authenticated access, synchronous and asynchronous queries, TAP queries, personal database storage, and storage through the Data Lab VOSpace.
	Image cutouts
	The Data Lab Simple Image Access (SIA) service provides a fast way to retrieve cutouts from SMASH images. For an example of how to use the SIA service, see this Jupyter notebook.
	Jupyter Notebook Server
	The Data Lab Jupyter Notebook server contains examples of how to access and visualize the SMASH catalog.

# SMASH DR1 Data Analysis

### Survey of the Magellanic Stellar History (SMASH)



Description	Analysis
Data Access	
	Jupyter Notebook Server
Analysis	
	We have set up a public Jupyter Notebook server to allow anonymous access and exploration of the SMASH catalog
Explore	and images. By clicking this link, you will start an instance of this server running. You can make changes to the example
Exprore	notebooks, but note that these changes will disappear once you close the page or the browser.
Results	
	Example notebooks in the Data Lab Notebook server
	You can view static versions of the example notebooks contained on the Jupyter Notebook server by selecting a
	notebook from the list below:
	<ul> <li>Basic access (field list, avg. photometry of a field, single-source light curve)</li> </ul>
	<ul> <li>Interactive filtering and plotting (Hydra II dwarf galaxy discovery demonstration)</li> </ul>
	Making an interactive source density map
	<ul> <li>Identifying ugr dropout candidates (Simple Image Access search and retrieval)</li> </ul>
	<ul> <li>Demonstrating criteria for separating stars and galaxies in the SMASH catalog (visualization of millions of points)</li> </ul>



### Identifying *r*-dropouts

#### **Catalog query**

For our query, we will look for objects that are undetected or have large errors in u, g, and r, but are detected and have small errors in i and z. We will only keep objects that have a match in the ALLWISE catalog. Using subqueries to limit the object and xmatch tables using indexed columns makes the query run much faster than it would otherwise.

#### In [2]: 11time

```
dbl='smash drl.object' # the SMASH object table with average magnitudes
dblsel='dbl.fieldid,dbl.id,dbl.ra,dbl.dec,dbl.umag,dbl.gmag,dbl.rmag,dbl.imag,'+\
    'dbl.zmag,dbl.werr,dbl.gerr,dbl.rerr,dbl.ierr,dbl.zerr,dbl.depthflag' 🖋 select ID, coordinates, and mags
db2='smash drl.xmatch' # the SMASH cross-match table, which contains cross-matches to ALLWISE
db2sel='db2.wise id,db2.wise wlmag,db2.wise wlerr,db2.wise wZmag,db2.wise w2err' # ALLWISE W1&W2 mags
dblwhere='(dbl.ndetu=0 or dbl.uerr>0.3) and ' + \
         (dbl.ndetg=0 or dbl.gerr>0.3) and ' + \
        ' (dbl,ndetr=0 or dbl.rerr>0.3) and ' + \
        (dbl,ndeti>0 and dbl,ierr<0.1) and ' + \
        ' (db1,ndetz>0 and db1,zerr<0.1)' # pick ugr dropouts
db2where='(db1.id=db2.id)' # only pick dropouts that are found in ALLWISE WI
# Create the query string.
query = 'SELECT '+dblsel+', '+db2sel+' FROM (SELECT * FROM '+db1+' WHERE depthflag > 1) AS db1, '+ \
    (SELECT * FROM '+db2+' WHERE wise match=1) AS db2 '+\
    'WHERE ('+db2where+' and '+db1where+')'
print "Your query is:", query
print "Making query"
# Call the Query Manager Service
response = queryClient.query(token, adql = query, fmt = 'csv')
df = pd.read_csv(StringIO(response))
print len(df), "objects found."
Your guery is: SELECT dbl.fieldid,dbl.id,dbl.ra,dbl.dec,dbl.umag,dbl.gmag,dbl.rmag,dbl.imag,dbl.zmag,dbl.uerr,dbl.ger
r,dbl.rerr,dbl.ierr,dbl.zerr,dbl.depthflag,db2.wise_id,db2.wise_wlmag,db2.wise_wlerr,db2.wise_w2mag,db2.wise_w2err FR
OM (SELECT * FROM smash_drl.object WHERE depthflag > 1) AS dbl, (SELECT * FROM smash_drl.xmatch WHERE wise match=1) A
S db2 WHERE ((db1.id=db2.id) and (db1.ndetu=0 or db1.uerr>0.3) and (db1.ndetg=0 or db1.gerr>0.3) and (db1.ndetr=0 o
r dbl.rerr>0.3) and (dbl.ndeti>0 and dbl.ierr<0.1) and (dbl.ndetz>0 and dbl.zerr<0.1))
Making guery
5769 objects found.
CPU times: user 52.3 ms, sys: 11.1 ms, total: 63.3 ms
```





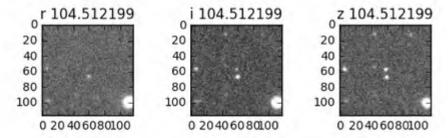
### Identifying *r*-dropouts

#### **Displaying the cutouts**

Now let's show the cutouts. The object in question is indeed invisible in the r-band image, but is visible in both i and z, and appears point-like.

```
In [12]: al=plt.subplot2grid((2,8),(0,0),rowspan=2,colspan=2)
imgplot = plt.imshow(rimg)
al.set_title('r '+idl.astype('string'))
a2=plt.subplot2grid((2,8),(0,3),rowspan=2,colspan=2)
imgplot = plt.imshow(iimg)
a2.set_title('i '+idl.astype('string'))
a3=plt.subplot2grid((2,8),(0,6),rowspan=2,colspan=2)
imgplot = plt.imshow(zimg)
a3.set_title('z '+idl.astype('string'))
```

Out[12]: <matplotlib.text.Text at 0x7fc06272fe90>



To go through the whole list of cutouts, the code from this notebook would be best put into a Python script and run from the command line, saving the images or making a figure showing all of the candidate objects at once.



# Discovering Hydra II dwarf

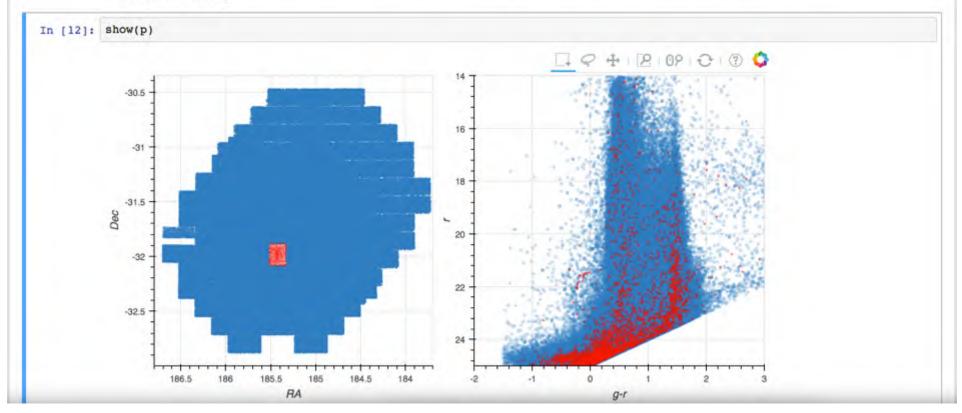
le	Edit		View	- le	150	rt	C	ell		Kern	e	Widge	ts	Help								Py	ython 2 C
+	90	Q	10	1	•	÷	H			C	Code		\$	63	Cell	foolbar	۵	ø	0	•			
			Que	y th	ne	SN	A	SH	D	R1 0	lata	base											
		1	We wi	que	нy	the	ave	rag	ed p	photo	metr	y table	from	the S	MASH	catalog	and s	selec	t Fie	eld 169, which we know a	contains the Hydra II dwarf.		
In	In [2]:		depti ranan decna mags dbase	<pre>ield = 169  # SMASH Field Number to query epth = 1  # minimum depth aname = 'ra' ecname = 'dec' ags = 'gmag,rmag' base='smash_dr1.object' id = 'fieldid' 'Create the query string.</pre>																			
			querj	-	C	sel (de (ab (gm (gm (gm (f	ect re ptl s() ag ag mag	(' ifl is is be j-ri	+ra +fi ag rp) no twe mag de	name d+' > % < ( t nt en ! ) be pth	<pre># ', = \ 1) a 0.5) 111) 9 and etween</pre>	*ed\' and and d 25) en -1	AND		v	+',de; ) & \	othfl	ag 1	ros	n '+dbase+ \			
		<pre>print "Your query is: ", query Your query is: select ra,dec,gmag,rmag,depthflag from smash dr1.object where (fieldid = '169' AND (depthflag &gt; 1) an</pre>														and							
		(abs(sharp) < 0.5) and (gmag is not null) and (gmag between 9 and 25) and ((gmag-rmag) between -																					
In	[3]		df =	pd.	re	qu qu ad_	ery	C1 (S	ien tri	t.qu ngIQ	D(re	(tokes		dql	= que	ry, fo	nt =	'csv	·)				
		1	<pre>print len(df), "objects found." Making query 297788 objects found. CPU times: user 307 ms, sys: 57.1 ms, total: 364 ms Wall time: 12.5 s</pre>											toti		54 ms							



## **Discovering Hydra II dwarf**

### The plots

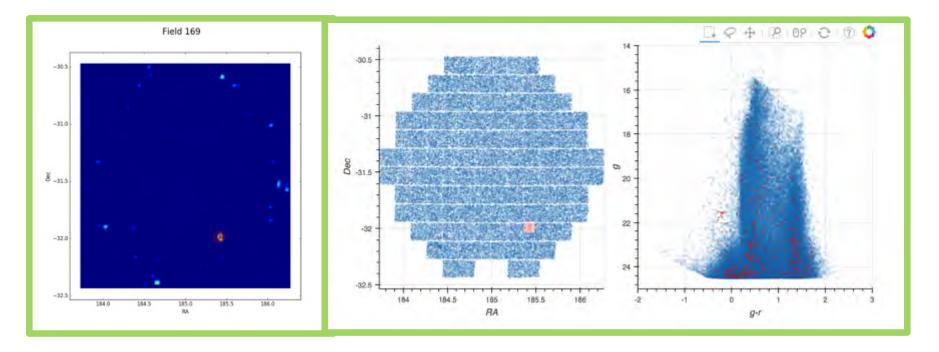
Finally, we render the plot. The figures are interactive, with ability to pan, zoom, and select samples of data that are then updated in the other plot. With the large number of points used here, the interaction can be a little slow, depending on browser and hardware. Try Box Select on the clump of points at lower left, where Hydra II is lurking.







# Automation of workflow



From Poster 154.25



# Coming in 2017

- Authentication
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