

# Targets of Opportunity at Gemini Observatory

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### References



ToO observations started in 2005, to enable follow-up of SWIFT detections.

Useful papers:

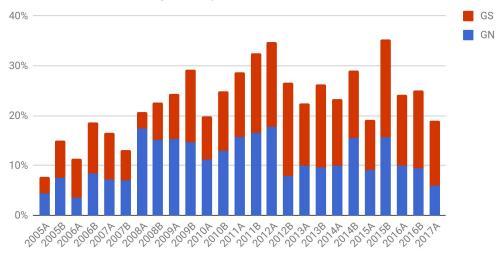
 Gillies et al. SPIE Proceedings Volume 6270, Observatory Operations: Strategies, Processes, and Systems; 62700W (2006); doi: 10.1117/12.672627

2. Roth et al. Observatory Operations: Strategies, Processes, and Systems III, Proc. of SPIE Vol. 7737, 77370N

### Demand

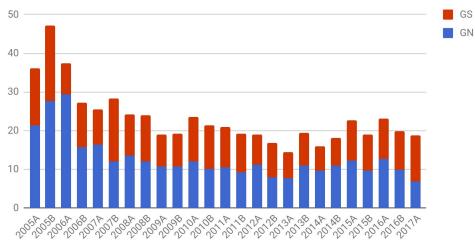


#### Fraction of ToO Programs per Semester



Semester

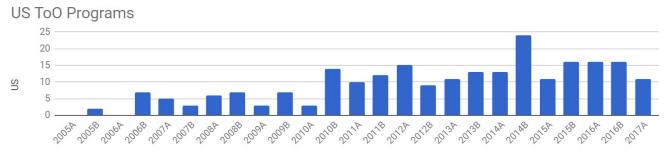
#### Average time Allocation per ToO Program



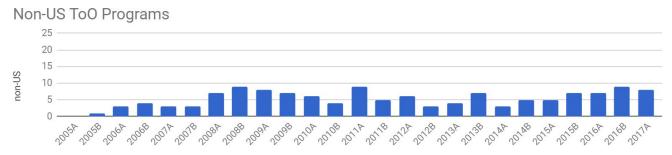
Semester

### Number of ToO Programs

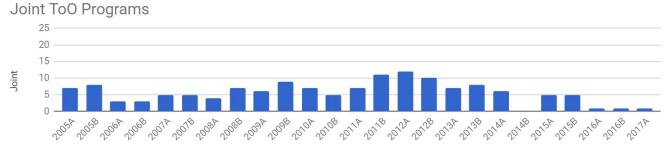




Semester



Semester

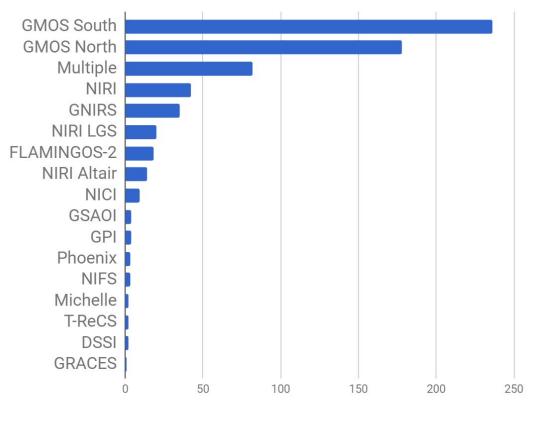


Semester

### Instruments used



Instruments used by ToO Programs



**ToO Programs** 

## How ToOs get into the database...

BERVATORY

"On Hold" "Prepared" 🖀 Templates Folder 🔛 [111] Spec rapid ToO manual trigger 🗄 🔚 [58] Spec standard ToO template **Observing Conditions** 🖻 📲 [64] Spec rapid ToO template GMOS-S Observing Conditions Target: GRB123556 GMOS-S Slew immediately! TOO: <Untitled> Constraint Log 📲 Observing Log 🖻 📌 Sequence 🗄 🔧 Sequence 🗄 🤧 Offset 🗄 😤 Offset Observe (1X) Observe (1X) 🌀 Flat 🌀 Flat 🚳 Arc Arc

The details of the trigger are formatted as an URL string which can be submitted to Gemini using any browser or URL tool such as wget. The following parameters are available.

Programmatically via a URL string:

prog email password obsnum	<ul> <li>program id</li> <li>email address for user key</li> <li>password for user key associated with email, site specific</li> <li>id of the template observation to clone and update, must be 'On hold'</li> </ul>
target	- name of the target
ra	– target RA (J2000), format 'HH:MM:SS.SS'
dec	- target Dec(J2000), format 'DD:MM:SS.SSS'
mags	<ul> <li>target magnitude information (optional)</li> </ul>
note	<ul> <li>text to include in a "Finding Chart" note (optional)</li> </ul>



### ...and what happens next



### **Standard ToOs:**

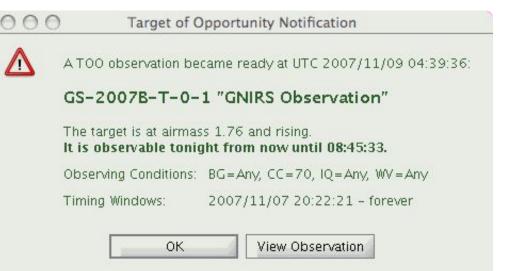
- Email notification to QCs
- Scheduled in the next plan

Observation GS-2016B-Q-41-28 (Standard ToO) is ready

The Standard Target of Opportunity Observation GS-2016B-Q-41-28, "GMOS-S 1" r-band acq" in program "DES Supernova Cosmology (South)", is ready to be observed.

### Rapid ToOs:

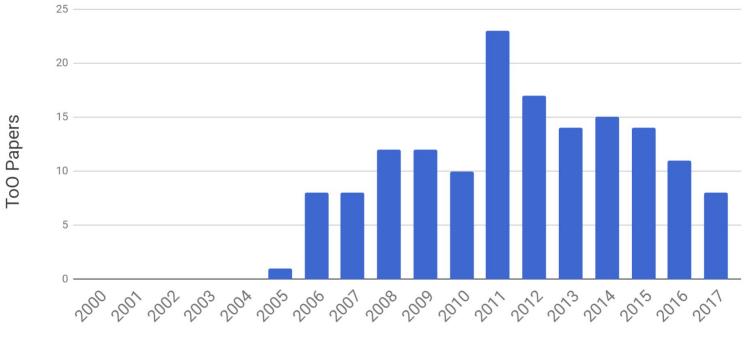
- Email notification to operations staff
- Popup+audible alert from the observer's OT



### **Publications**



Publications from ToO Programs

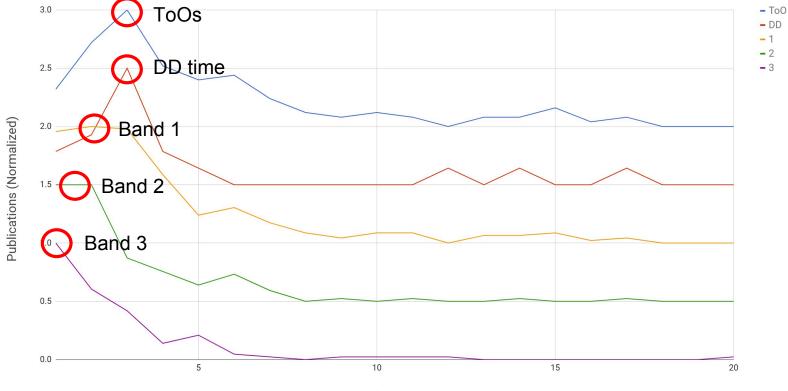


Year

# Impact



Normalised Cumulative Impact



Impact

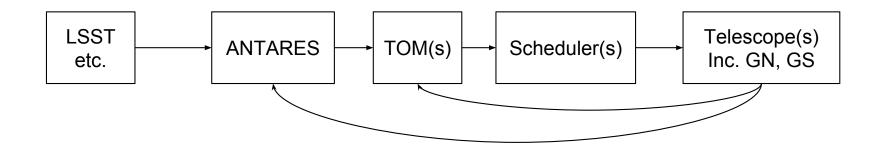
# **Recent Policy Development**

### MINI G

	MGEMINI
•	OBSERVATORY
Scienc	
Innouncements	Change page style Default  ▼
Contact	Policies for Competitive ToOs
ibrary	Gemini Observatory > Sciops > Observing With Gemini
10	Effective November 20, 2017, the policies described below will regulate the activation of Target of Opportunity (ToO)
ciops	programs, the priorities for their execution as well as the data access rights for the acquired data. The policies apply when multiple teams activate ToOs on the same target and on the same or subsequent nights.
Semini Home	
emini Home	Approval of ToO Triggers and DD Time
cience Visitors at Gemini	To be activated, Queue ToOs must be scientifically aligned with the nature of the event for which the trigger is
Observing With Gemini	requested. This is a general rule that applies to all ToO programs. DD proposals will not be accepted if they effectively duplicate, in their observational setup, an existing Queue ToO tha
PI's guide to Proposal	has been activated to observe the same target. DD proposals that are accepted prior to the activation of an existing
Routes and Observing modes	Queue ToO will be subject to the Prioritization and Data Sharing Policies discussed in II. and III. below.
Proposal Routes and	Prioritization for ToO Triggers and DD Time
Observing Modes Completion expectations	In case of multiple Queue ToO triggers on the same target and night, the Gemini Observatory will attempt to work with
Observing Overheads	the PIs to prioritize the observations so that the science is maximized. If no consensus can be reached amongst the PIs, or if no coordination is possible, the Gemini Observatory reserves the right to prioritize the observations according
Proposal Submission	to the following criteria, ranked in order of priority: 1) how well the scientific justification of the proposal aligns with the
2018A Call for Proposals	
Nightlog GN Nightlog GS	constraints, such as whether the Gemini observations are coordinated with data from other Observatories (e.g. Chandra, HST).
Phase II and S/W Tools	As a rule, all DD programs will be given lower priority than Queue ToO programs. However, DD programs that are awarded to extend a Queue ToO program that has used all of its allocation during the course of an observing
Changing Approved	campaign will have higher priority than other DD programs.
Programs Policies for	Paliaias an Data Access Diabts for TaO and DD Broavams
Competitive ToOs	Policies on Data Access Rights for ToO and DD Programs
Advice for Band 3 Programs	The following policies only apply when multiple ToOs (rapid or standard) are triggered on the same target and, for objects
What to expect	that vary significantly on timescales shorter than 24 hours, on the same night. A distinction is made between ToOs approved as part of the regular Queue, and ToO approved using DD time.
Telescone Time Charging	

### Future





### Future



Currently:

- Daily plans are created manually.
- Night-time observers switch plans as conditions change.
- Standard ToOs are planned in.
- Rapid ToOs are not.

Weaknesses:

- RToOs disrupt the plan
- Managing timing windows
- We don't cater for complicated follow-up cadences (eg. observe every N days, logarithmic cadences)

Mauna Kea - 20070923 start from Q79-61 after the IQ/LGS teams are done with their vilight N07B-Q-121 [89] 1 - 2 GMOS-N B600, 1.5arcsec, none OIWFS SDSS J214652.19+1112... rvations at the beginning of the night. :(I am putting this band4 observation as a filler. N07B-Q-33 [25] I - 3 GMOS-N B1200, none, 1.0arcsec OIWFS WD2326+049 servations. Please observe this tonight. This has higher priority than the rest of the observations s N07B-Q-79 [61] 82 - 140 NIRI H-cont PWFS2 XMMUJ2235.3-N07B-Q-79 [61] 82 - 140 NIRI H-cont PWFS2 XMMUU2235.3-2557 70 70 A you will probably be on time, if not early, when you finish this observation. Please do a speephot std (or any filler of you 100% N07B-Q-79 [77] 1 - 53 NIRI H-cont PWFS2 XMMUJ2235.3-2557 . No need to reslew. You will not finish this observation before hitting a rmass=2 51 - 134 NIRI 1 - 7 NIFS 1 - 6 NIFS H2 1-0 PWFS2 PP13S PWFS2 HH22-Spitzer PWFS2 HIP 30594 100% N07B-Q-39 [10] 100% N07B-Q-39 [13] 01:53 05:44 09:35 02:28 06:23 10:18 td for O39-10 and 13. 00:36 100% N07B-Q-39 [3] 1-3 NIFS PWFS2 OriBemm S 01:53 05:44 09:34 Stop when convinien

With more ToOs & follow-up this will become unmanageable!

Improvements being made under OCS Upgrade Project.

See Oct 2017 Gemini Focus & contact <a href="mailto:bmiller@gemini.edu">bmiller@gemini.edu</a> if you'd like to contribute ideas.



### **Automated Scheduling**

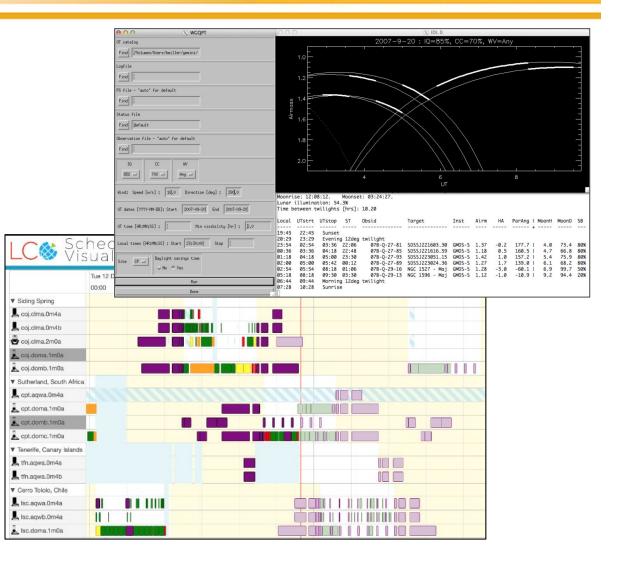


Automated scheduling is needed to increase flexibility and reduce workloads.

The schedule will adapt as new ToOs arrive (or conditions change).

We are experimenting with algorithms.

Will hopefully be handled by a ToO network scheduler (see Blum presentation).



# Thank you



### Targets of Opportunity at Gemini Observatory

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Narrative notes & Takeaway points to accompany slide set

#### Slide 4

Takeaway - after a ramp early in the program, both the proposal load (currently 10% per telescope) and the average length of a program allocation (currently 8-10 hours) have stabilized.

### Slide 5

Total number of US ToO programs (combination of Rapid and Standard) has risen fairly continually. Other partners' ToO numbers have varied slowly with time. Joint ToO programs have recently dropped to zero.

#### Slide 6

Takeaway: the instrument distribution for ToO programs has been quite similar to that for other programs. Dominated by GMOS, but that is partly because GMOS combines all optical modes in one instrument while IR modes are distributed across multiple cryostats.

#### Slide 7

Both methods for triggering rely on a template observation being present in the database. By-hand method requires the PI to go into the program, set the target coordinates and set the observation "ready", the programmatic method allows this to be done by a script - e.g. for the case when you need the fastest response.

#### Slide 8

Standard ToOs don't need immediate observations. They are scheduled in by the QCs, either the following night or subsequent nights. QCs get an automatic email when the observation is set "ready".

Rapid ToOs need fast response - either immediate (drop everything, abort the current observation and observe this - in which case the RToO program is charged for any time lost to the currently executing program) or as soon as the current observation is complete.

Rapid ToOs result in an audible alert in the control room and a popup in the observer's Observing Tool. Hard for observers to miss!

### Slide 9

Publications from ToO programs appear to have peaked in 2011, even allowing for publication lag and the fact that 2017 publications were not complete at the time of the search. This plot is representative; but for simplicity, it includes publications resulting from a single ToO program only.

### Slide 10

This chart uses D.Crabtree's "impact" parameter for publications, which ratios the number of citations of each publication to the average of Astronomical Journal papers from the same year. It shows that the peak impact for publications arising from Band 1, 2 and 3 programs (again,

single program papers) increases from band 3 through band 1, and that DD and ToO publications peak at a higher impact still.

#### Slide 11

Points the reader to the new statement of policy on data access rights in the case of "competitive ToOs" - on which multiple programs trigger at the same (or close to the same) time. This policy became necessary because of the LIGO neutron star merger event in 2017.

#### Slide 12

In the future, we're envisaging the telescope (or telescopes, as we anticipate being in a network with multiple smaller and larger apertures) are sent observations by a software scheduler which in turn receives requests from a TOM (Target Observation Manager); this is a program-specific agent which takes a pre-filtered alert stream from the ANTARES broker and generates its own specific observation requests.

#### Slide 13

Our current system won't be up to handling a 30-50% ToO rate per night that one might anticipate in the LSST era. This slide, and the next and final one, outline some changes we'll need to make. Input is invited to our OCS upgrades program, which this work will be done in parallel with.