

# Spectroscopic Data Reduction on your web browser

Simón Torres-Robledo<sup>1,2</sup>, César Briceño<sup>1,2</sup>. <sup>1</sup>SOAR Telescope, La Serena, Región de Coquimbo, Chile. <sup>2</sup>NSF's NOIRLab, Casilla 603, La Serena, Chile.



#### Abstract

The goodman-pipeline package provides all the functionality needed to reduce raw data obtained with the Goodman High Throughput Spectrograph on the Southern Astrophysical Research (SOAR) 4.1m telescope, produce 1-D wavelength calibrated spectra. However, many science use cases, especially those targeting transients and variable objects, depend critically on having fully reduced spectra available shortly

after the raw data are written to disk; the goodmanpipeline was not designed to fulfill this requirement. To address this issue we have developed a system that combines the data reduction elements of the goodmanpipeline with a Web API implementation within a webbased structure, to provide the user with immediate spectroscopic processing through a web browser.

This approach provides users with fully reduced spectra

just seconds after the shutter closes, which could be critical when making real time decisions such as whether to obtain more data of the same object, switch to a different configuration or instrument, submit an observation request to another facility, or simply discard the object as one of no interest for the specific science case, and move to a new target.

**User Interface** 

Built using Django because of its integrated security. It's

a responsive design using Bootstrap 4. For interactive

#### Architecture

The applications run in Docker containers and are built using Django and Django Rest Framework. The main components are:

Private Web API: Implements the data reduction routines developed in the goodman-pipeline package, as a web API. Does not require authentication, since it can only be accessed from another container within the same docker network. Also referred as data reduction API or reduction API, it is independent to allow expansion to other instruments.

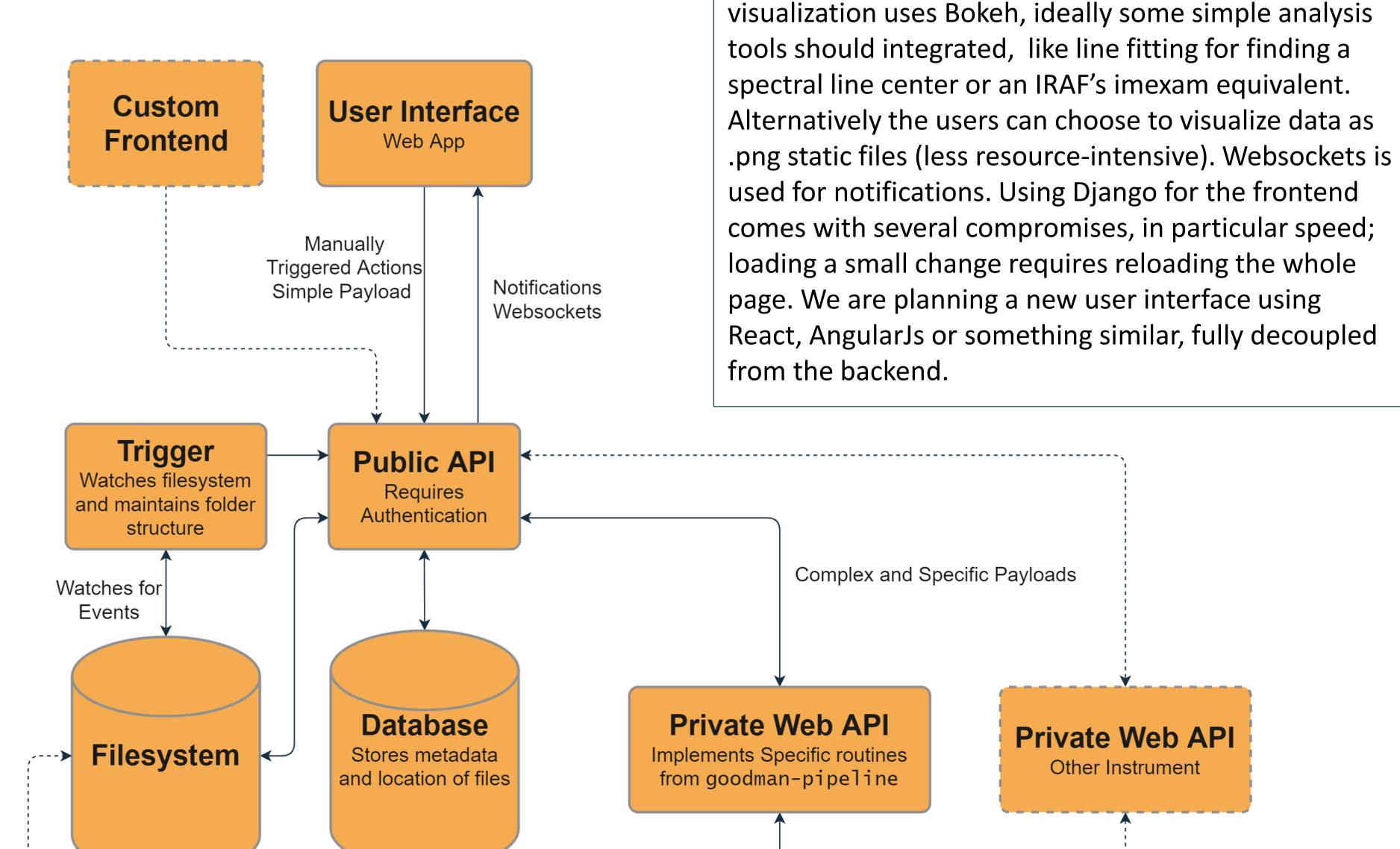
Public Web API: Provides several endpoints for very specific actions; it requires authentication. It doubles as a proxy for the private web API.

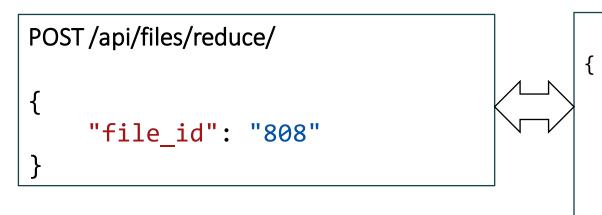
Events Trigger: Watches filesystem events, in particular, detects when a new file is written to a specific folder and triggers the appropriate actions, such as moving it into the corresponding new location.

## Operation

The Goodman Spectrograph is a highly configurable instrument. There are 20 pre-set modes that allow for full spectral range coverage at resolutions R~800-14000 depending on the grating. Though the Goodman-pipeline package is capable of handling data obtained with any of the pre-sets, we initially constrained the options to just 5 of the most used ones.

### **General Workflow** and Relations





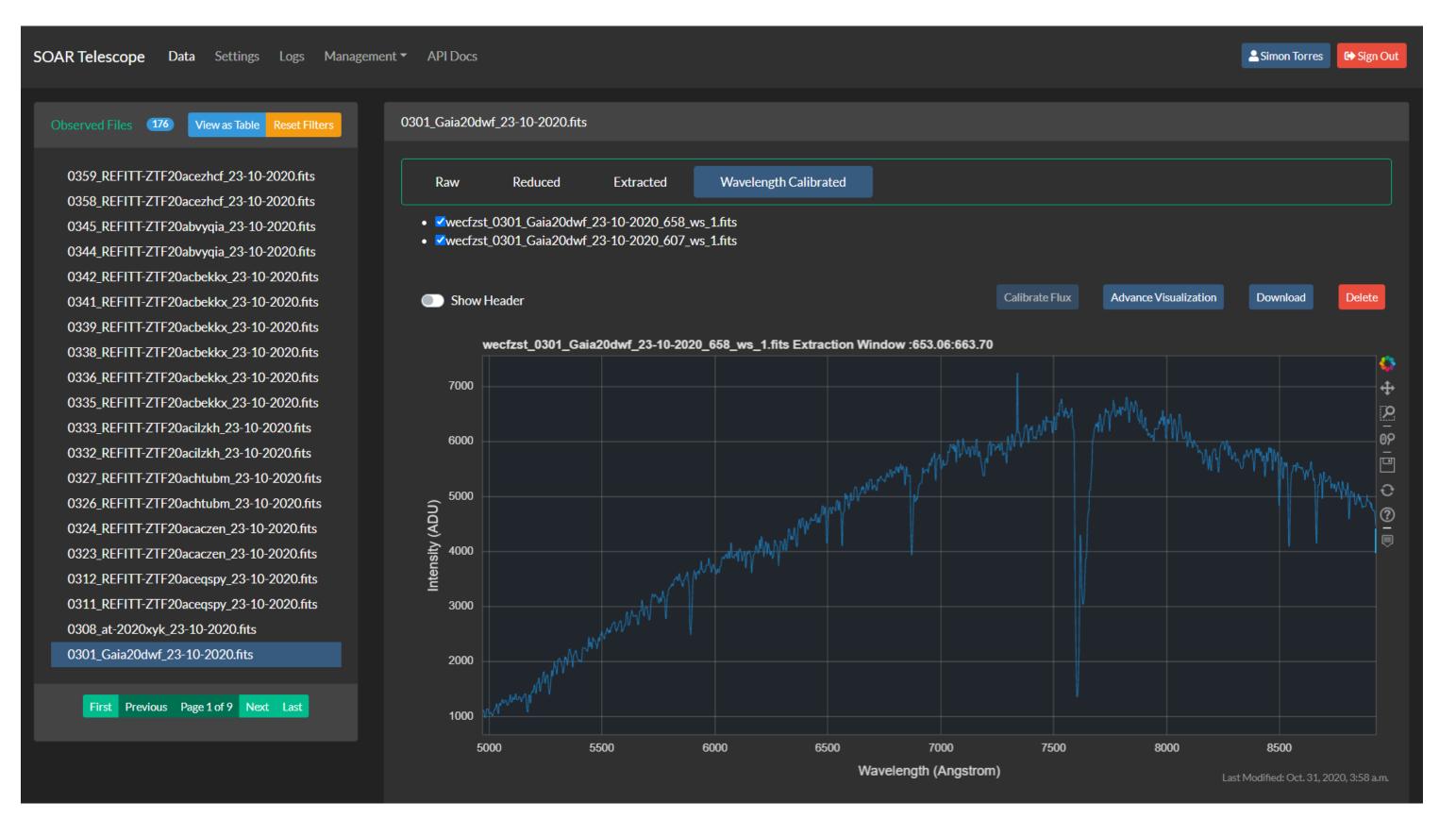
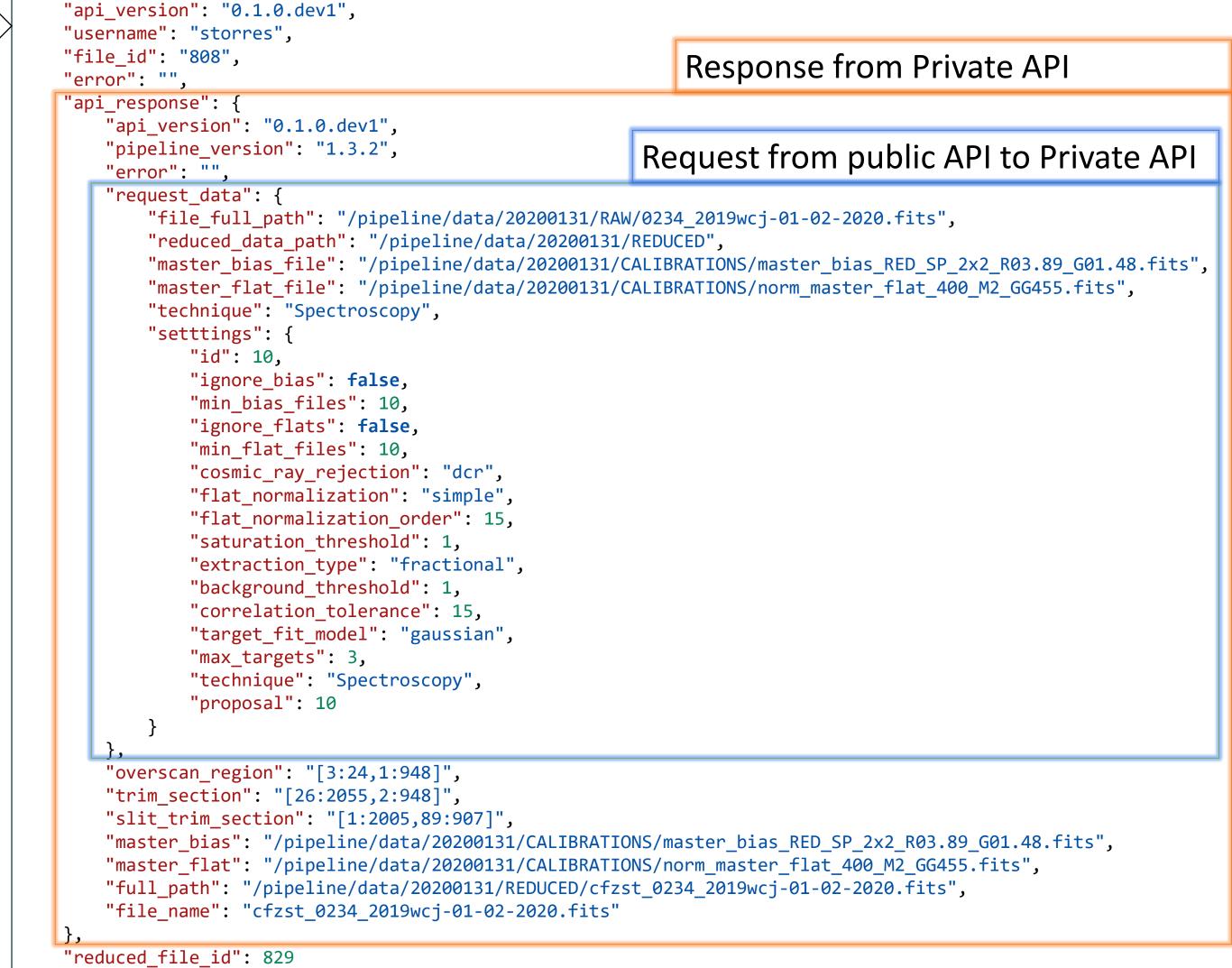


Figure 1. Screenshot of the current user interface. The Observed Files section lists all the raw files that the user has authorization to see, calibration files are available for everyone. The large card to the right contains a preview of the raw image and all the files that were produced from it, i.e. the reduced, all the extracted and wavelength calibrated.



#### Contact

Simón Torres Robledo NSF's NOIRLab Email: simon.torres@noirlab.edu Website: https://github.com/simontorres Phone: +56 51 2205348

## Acknowledgments

The authors would like to acknowledge the important contributions from Bruno Quint, David Sanmartim and Tina Armond in the development of the goodman-pipeline package.

This research made use of Astropy, a community-developed core Python package for Astronomy (Astropy Collaboration, 2013 and 2018).

This work has been developed at the Southern Astrophysical Research (SOAR) telescope, which is a joint project of the Ministério da Ciência, Tecnologia, Inovaçãos e Comunicações do Brasil (MCTIC/LNA), the U.S. National Optical Astronomy Observatory (NOAO), the University of North Carolina at Chapel Hill (UNC), and Michigan State University (MSU).

#### References

- 1. Clemens, J. C., Crain, J. A., & Anderson, R. 2004, vol 5492 of
- Proceedings of SPIE, 331.
- 2. Astropy Collaboration et. al., 2013, A&A, 558, A33.
- Astropy Collaboration et. al., 2018, AJ, 156, 123.
- Marsh, T. R., 1989, PASP, 101:1032-1037.
- Mellado, P., 2019, ASPC, 523, 41.
- 6. Pych, W., 2004, PASP, 116, 148. 7. Torres-Robledo S., Briceño C., 2019, ASPC, 523, 203.
- 8. Torres-Robledo et. al., 2020, ASPC, 522, 533. 9. Van Dokkum, P. G., 2001, PASP, 113, 1420.

10. Young M. D., Gopu A., 2018, https://github.com/IUSCA/mean-stack