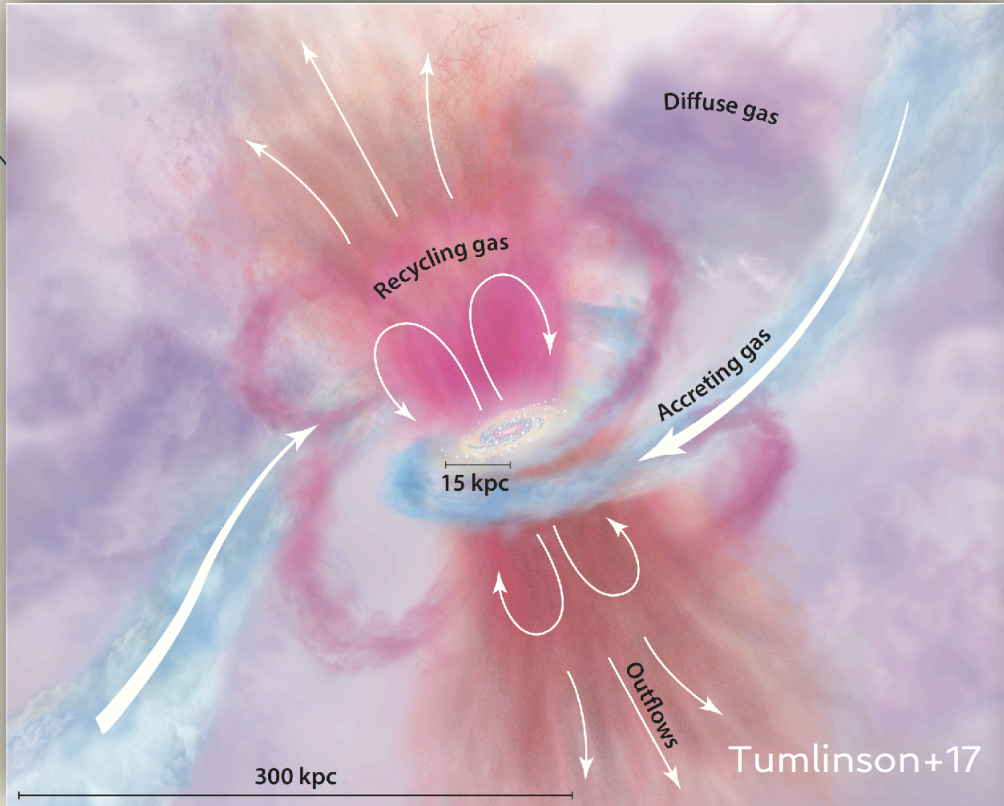


# GEOSTATISTICS OF GALAXIES – MAKING THE MOST OF METALLICITY MAPS

Benjamin Metha (he/him) | December 14, 2023



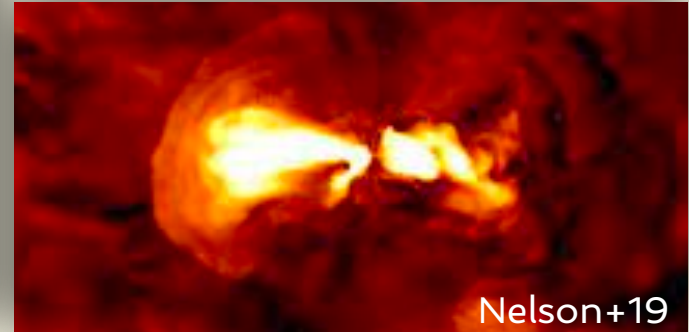
# GALAXIES ARE 3D AND COMPLICATED



Inflows/outflows



Mergers



Feedback/Turbulence

# OUR MODELS ARE 1D AND SIMPLE

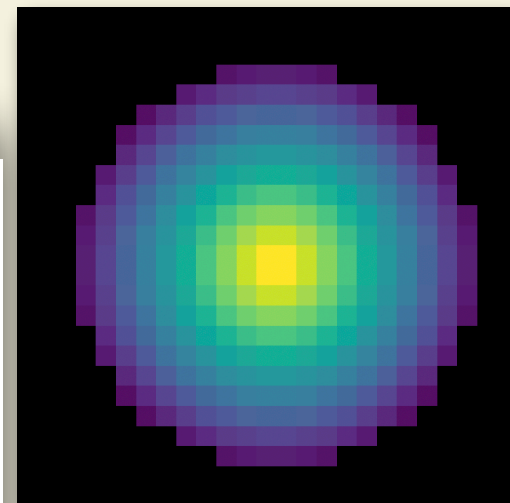
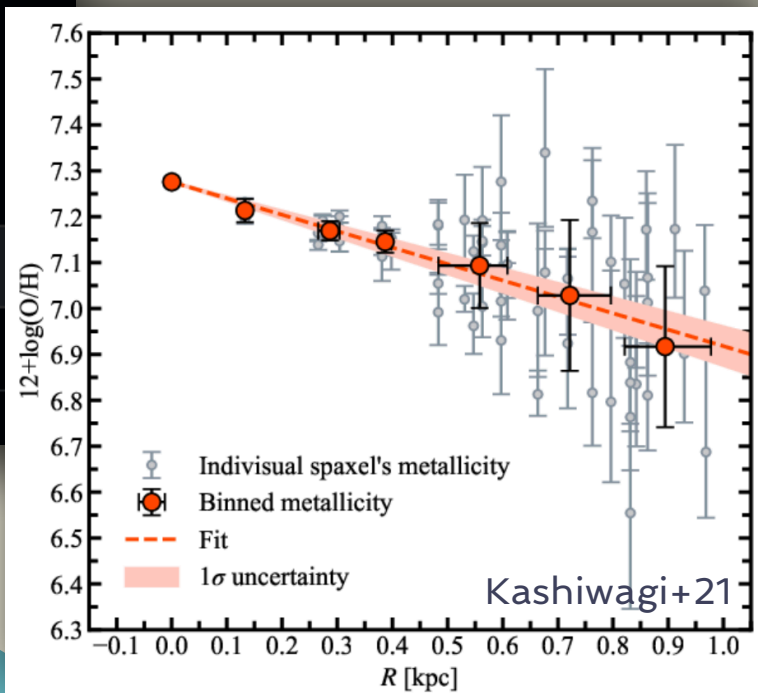


← Start New Search

QUICK FIELD: Author First Author

"metallicity gradient"

Your search returned **3,406** results

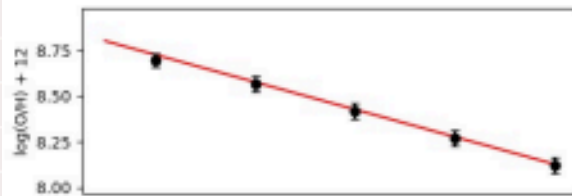
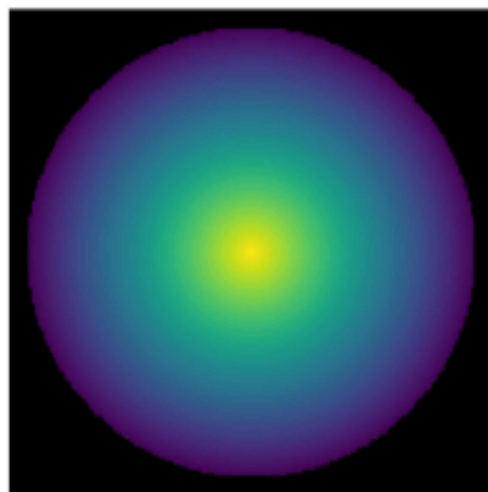


=> inside-out  
galaxy formation

# The Metallicity Gradient can't distinguish...

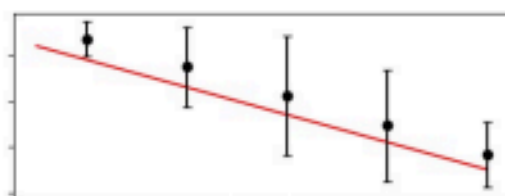
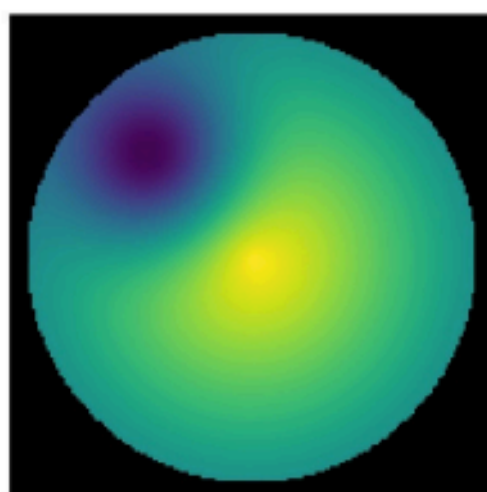
## a smooth, linear profile

indicative of secular evolution and inside-out star formation



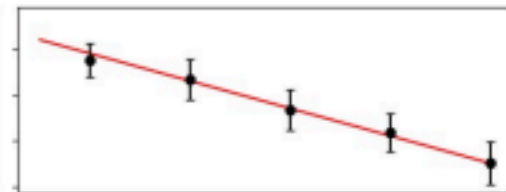
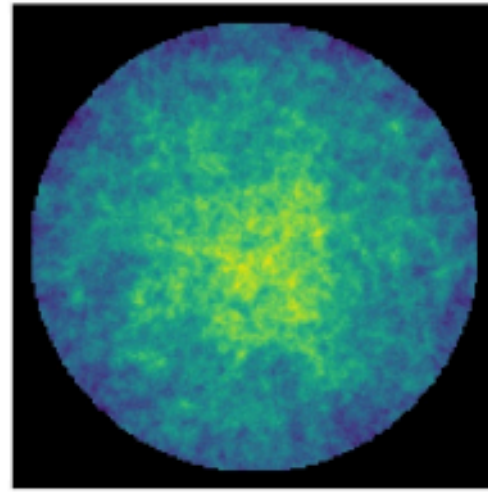
## large-scale structure

indicative of e.g. recent minor mergers, or extreme outflows

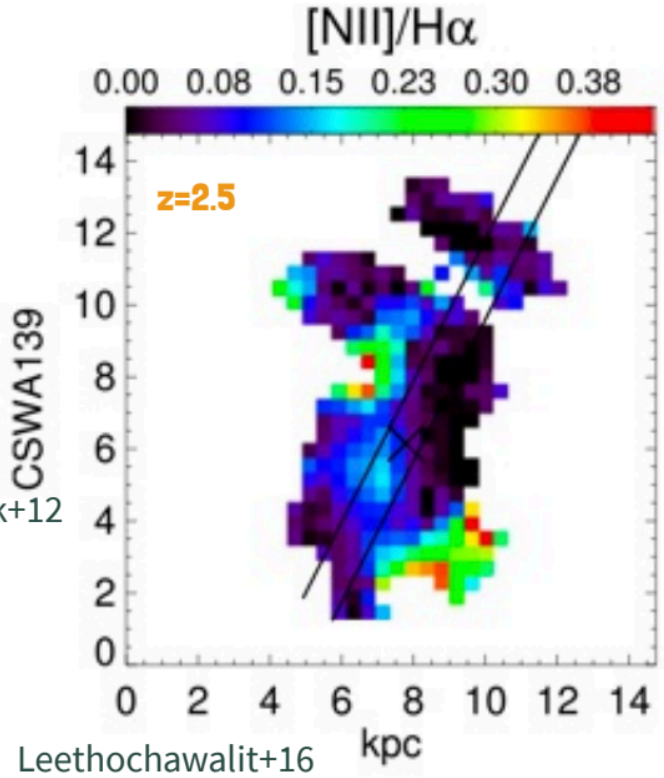
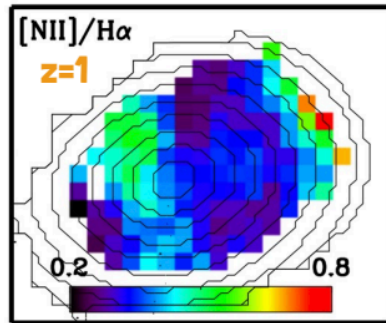
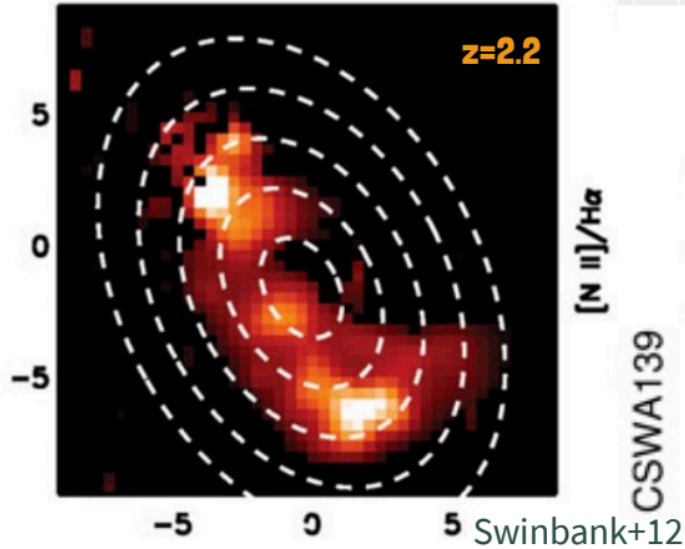
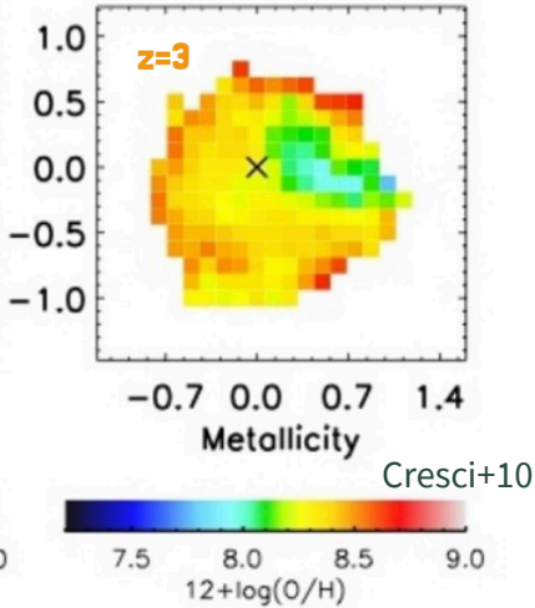


## small-scale substructure

indicative of inefficient mixing, constraining SNe-driven feedback



# Galaxies at cosmic noon aren't smooth!



# JWST-GLASS IS INCREDIBLE



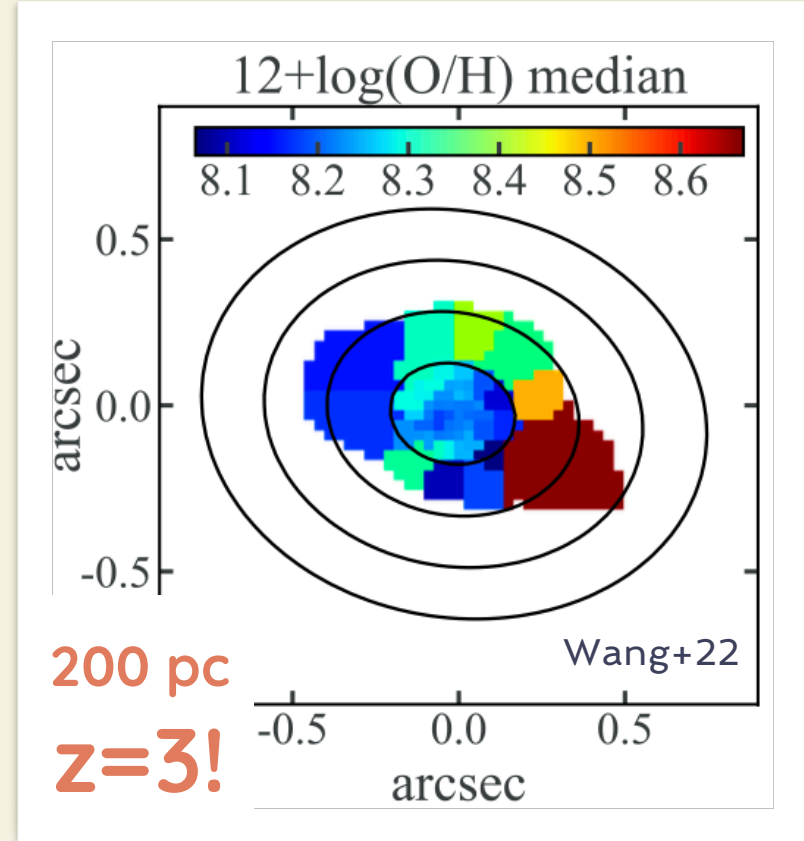
## JWST SENSITIVITY

NIRISS can do 65 mas resolved spectroscopy on  $m_{AB} < 24$  targets!



## GRAVITATIONAL LENSING

Magnification improves resolution and sensitivity.

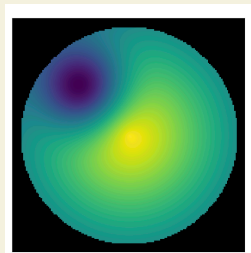
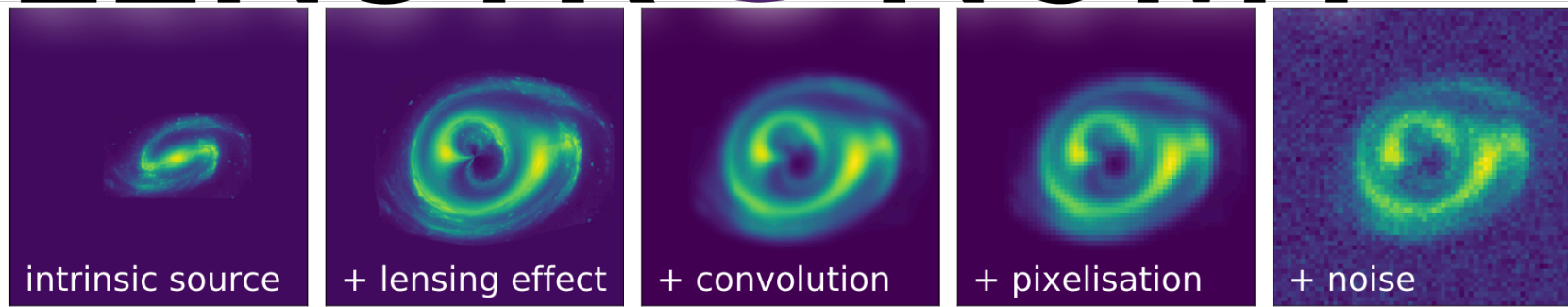




WHAT DO WE  
DO NOW?

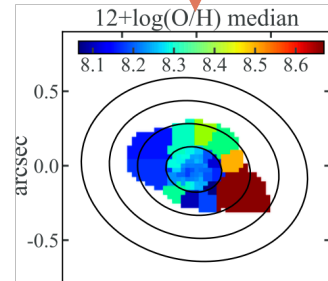
METALLICITY MODELLING IN:

# LENSTRONOMY



MODELS WITH  
2D STRUCTURE

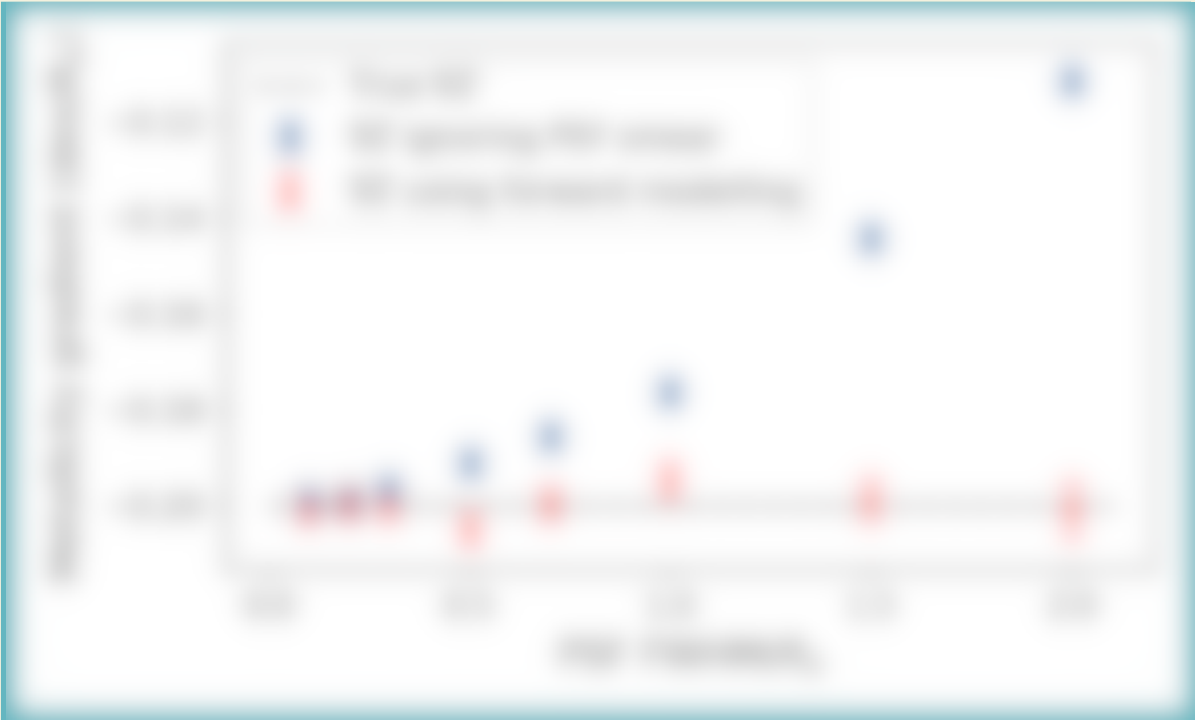
SUBMITTED!



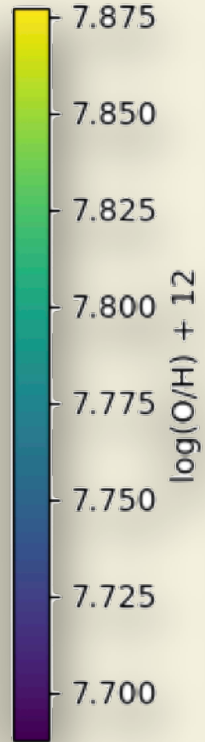
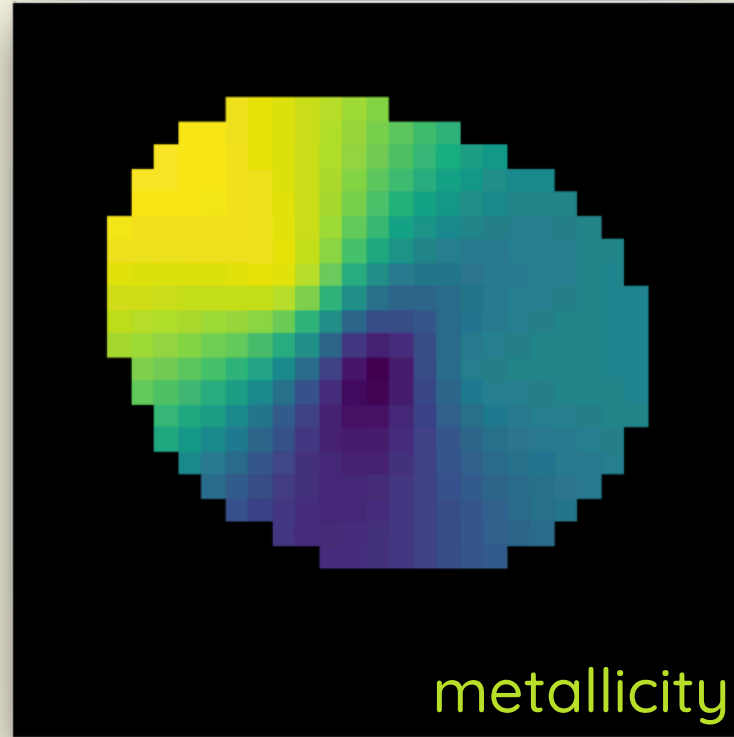
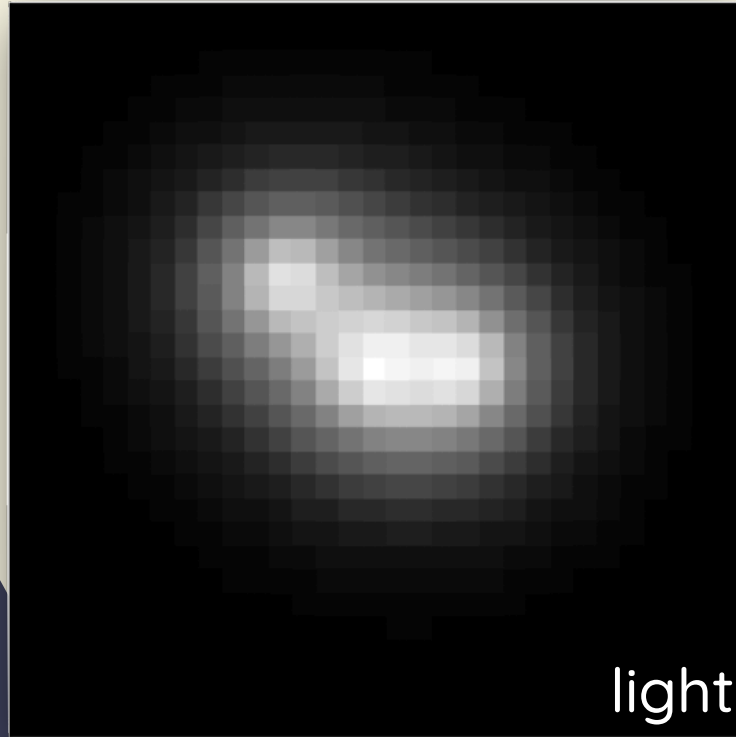
REAL DATA



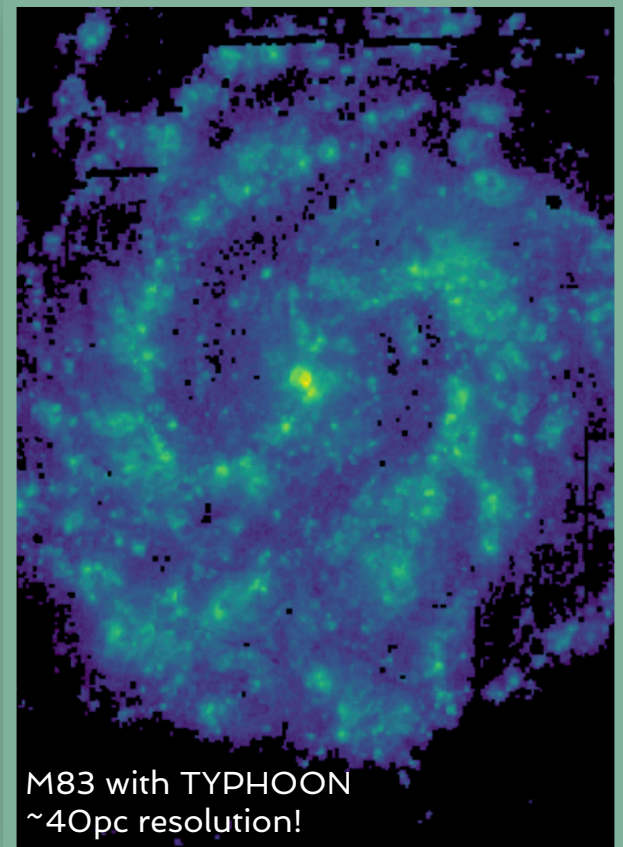
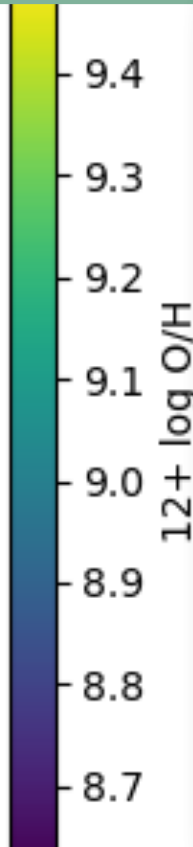
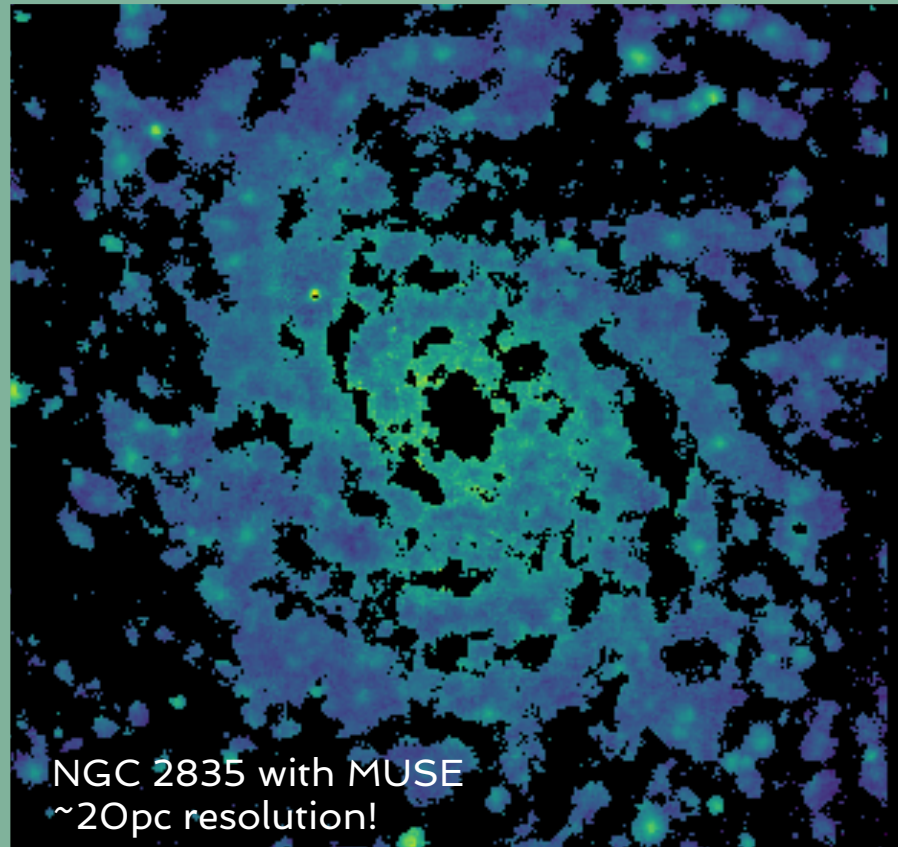
# CORRECTIONS FOR PSF SMEAR



# FITTING MORE FLEXIBLE MODELS

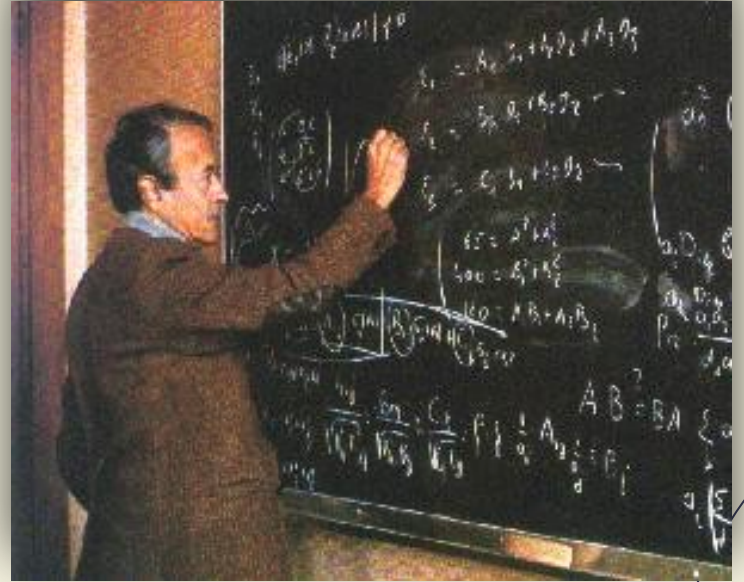


# LOCAL DATA LOOKS EVEN COOLER



# CAN WE DO EVEN BETTER?

“GEOSTATISTICS IS THE STUDY OF A **RANDOM PROCESS** THAT VARIES OVER A **SPATIAL DOMAIN** IN A PREDICTABLE WAY”



Georges Matheron. Image credit:

Centre de Geosciences

# HIERARCHICAL MODELLING

Observation error:  
known from telescope properties

$$Z_{obs}(x) = Z_{true}(x) + \epsilon(x)$$

$$Z_{true}(x) = \mu(x) + \eta(x)$$

Process mean:  
accounts for the metallicity gradient

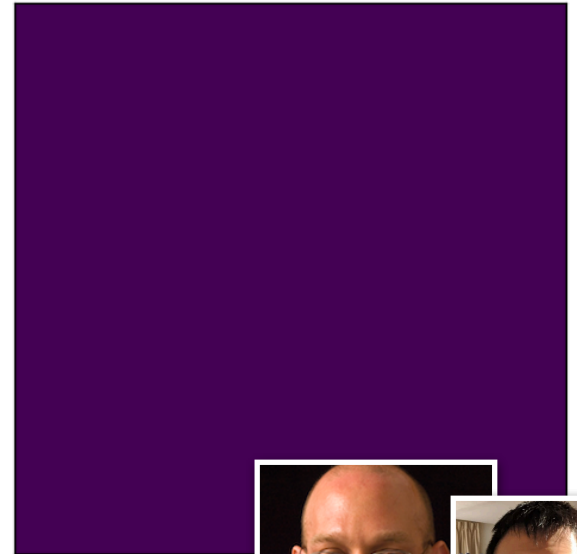
Random fluctuations:  
spatially-correlated deviations

IF WE CAN MODEL  $\epsilon(x)$ ,  
 $\mu(x)$ , AND  $\eta(x)$ , THEN WE  
CAN MODEL THE GALAXY!

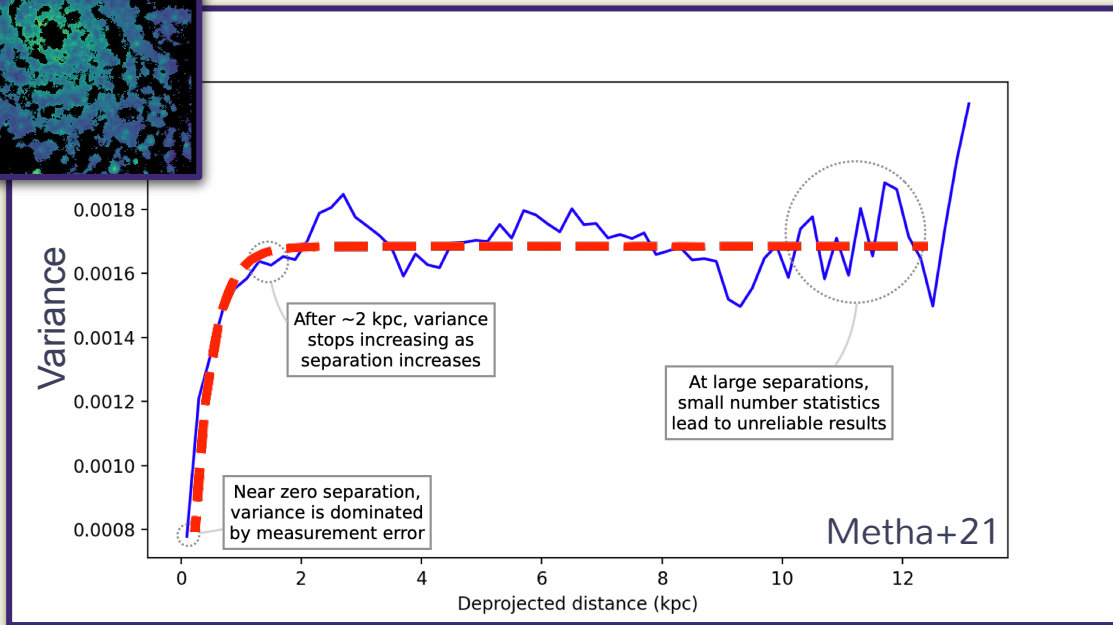
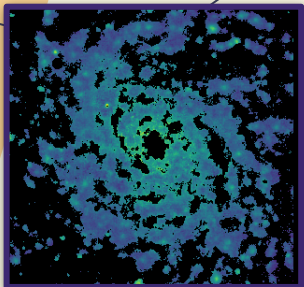
# TESTING AN ANALYTICAL MODEL

Krumholz & Ting (2018)

- Stochastic partial differential equations
- Assumptions:
  - Constant, uniform star formation
  - **Linear diffusion**
  - No stellar winds
- Predictions:
  - Spatial correlation of metallicity



# SEMIVARIOGRAMS REVEAL SUBSTRUCTURE



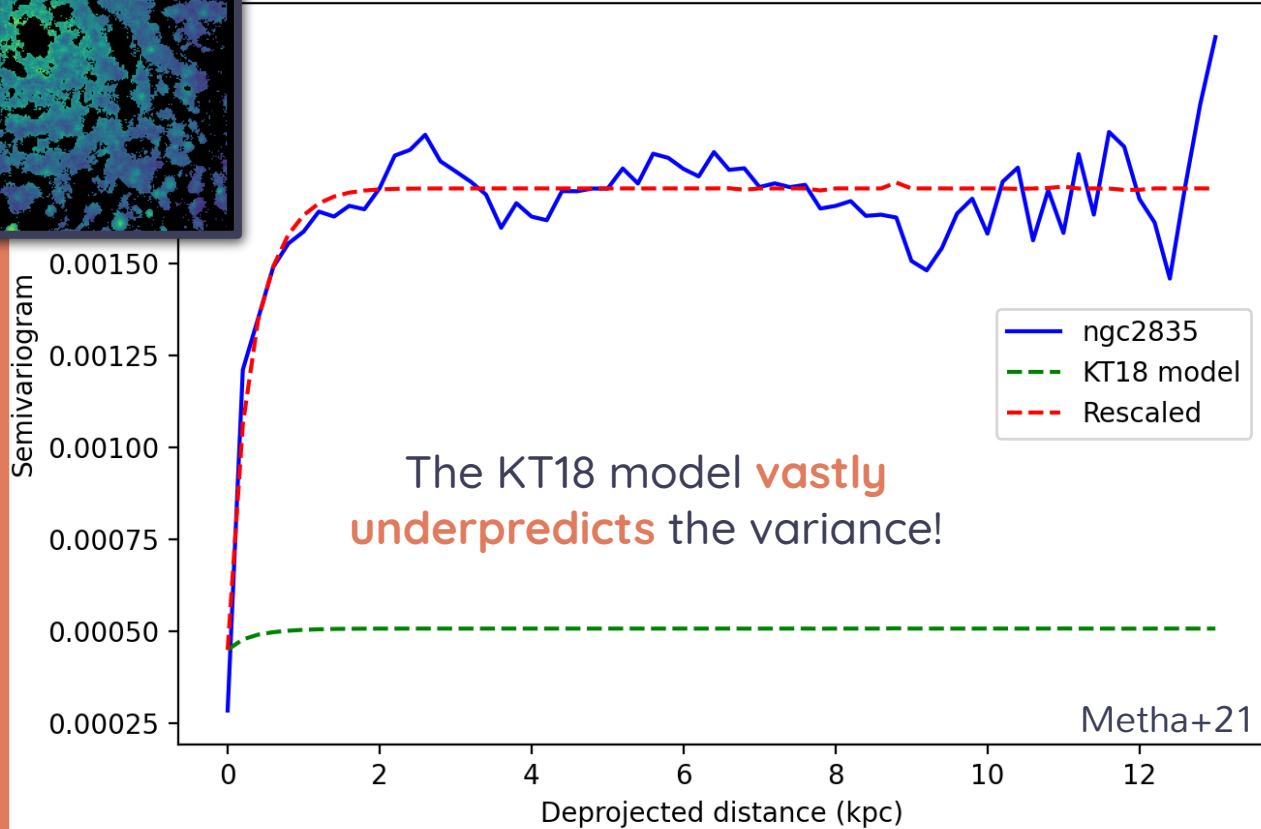
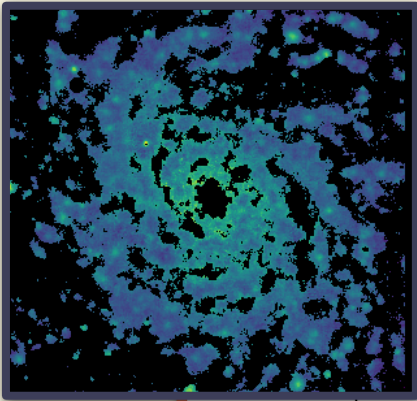
WORKS ON **BAD DATA**

Naturally separates structures from uncorrelated **noise**.

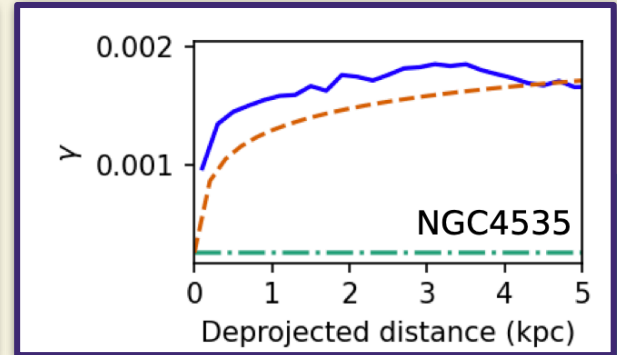
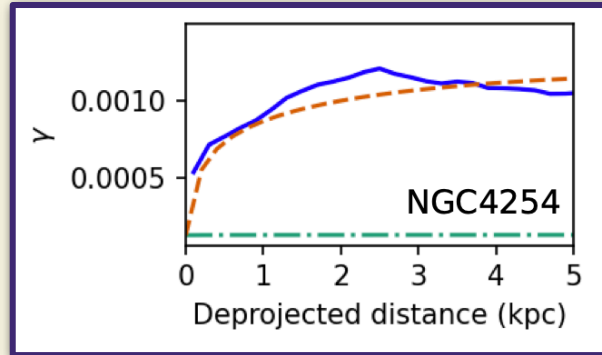
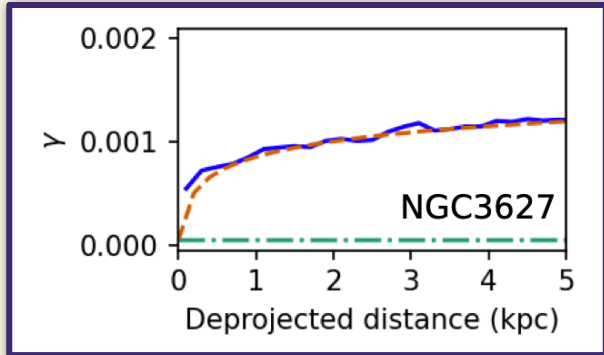
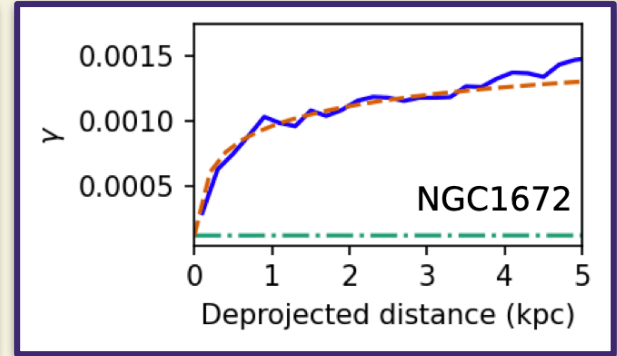
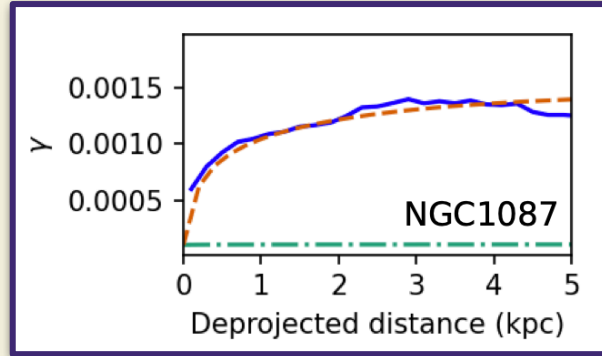
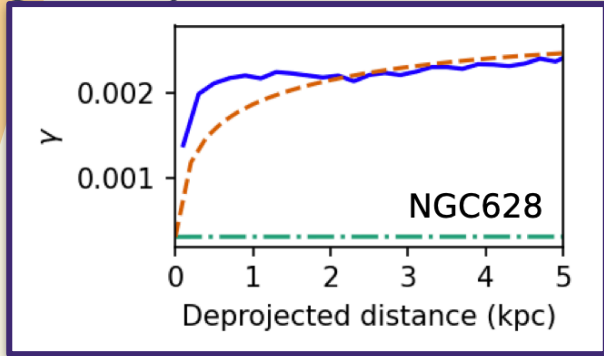
**HIGHLY INFORMATIVE**

Shows **covariance structure**, like a power spectrum.





# FOR 6 OTHER GALAXIES



- Empirical Semivariogram
- - - KT18 model
- - - Rescaled KT18 model



GALAXIES AREN'T  
VERY WELL MIXED!

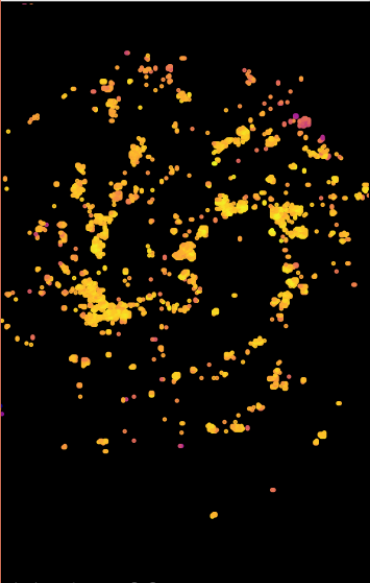
# GEOSTATISTIC MODELS → MORE DETAILS

Metal rich



Metal poor

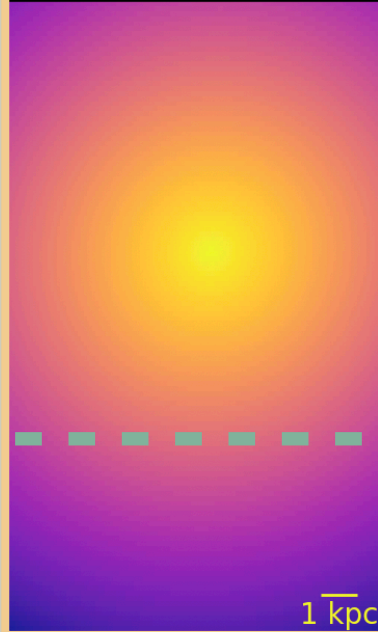
Observations



Metha+22

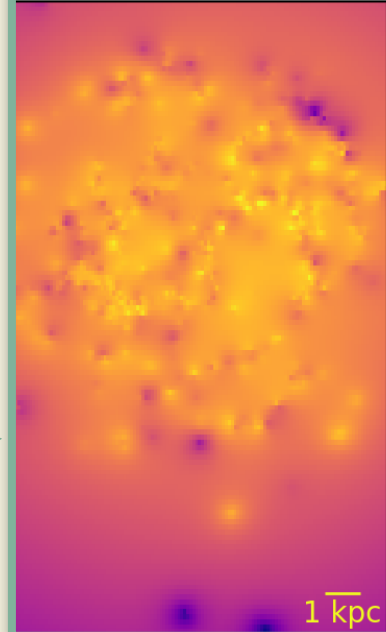
1 kpc

$\nabla Z$  predictions



1 kpc

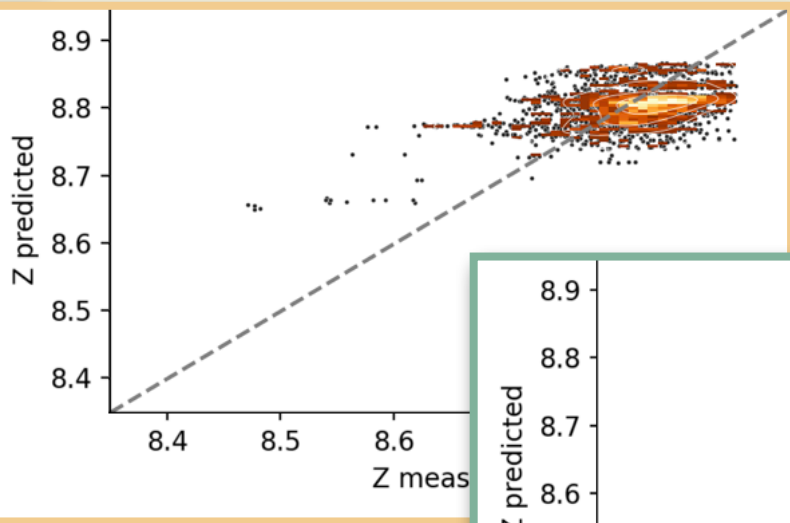
Geostatistical Model



1 kpc

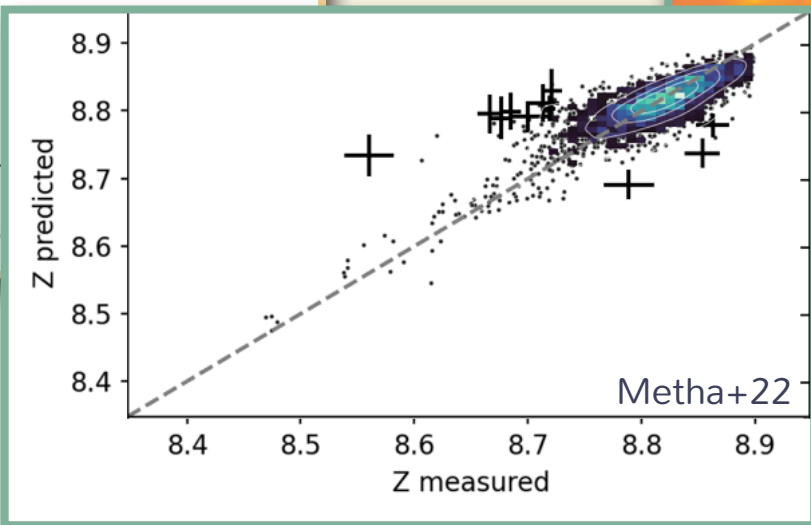
# GEOSTATISTIC MODELS → BETTER PREDICTIONS

$\nabla Z$  predictions



1 kpc

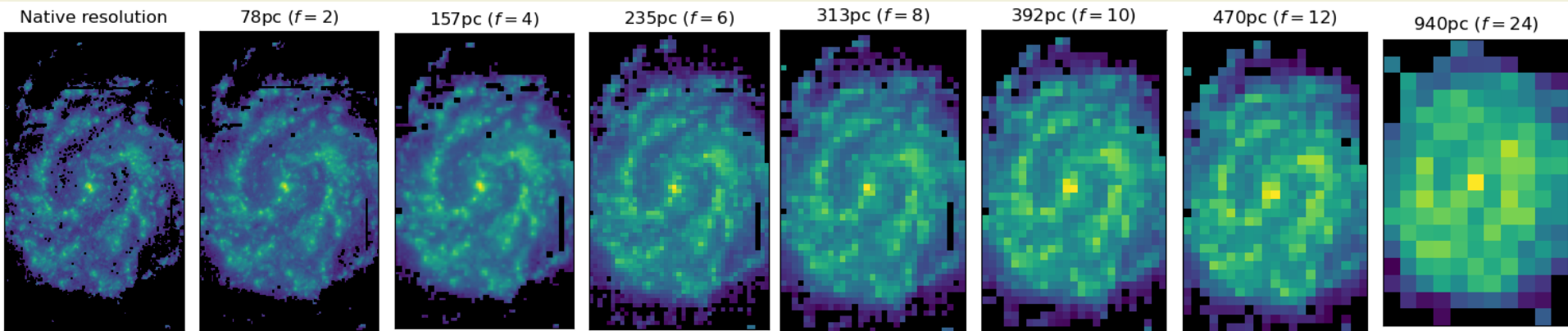
Geostatistical Model



Metha+22

1 kpc

# RESOLUTION LIMITS



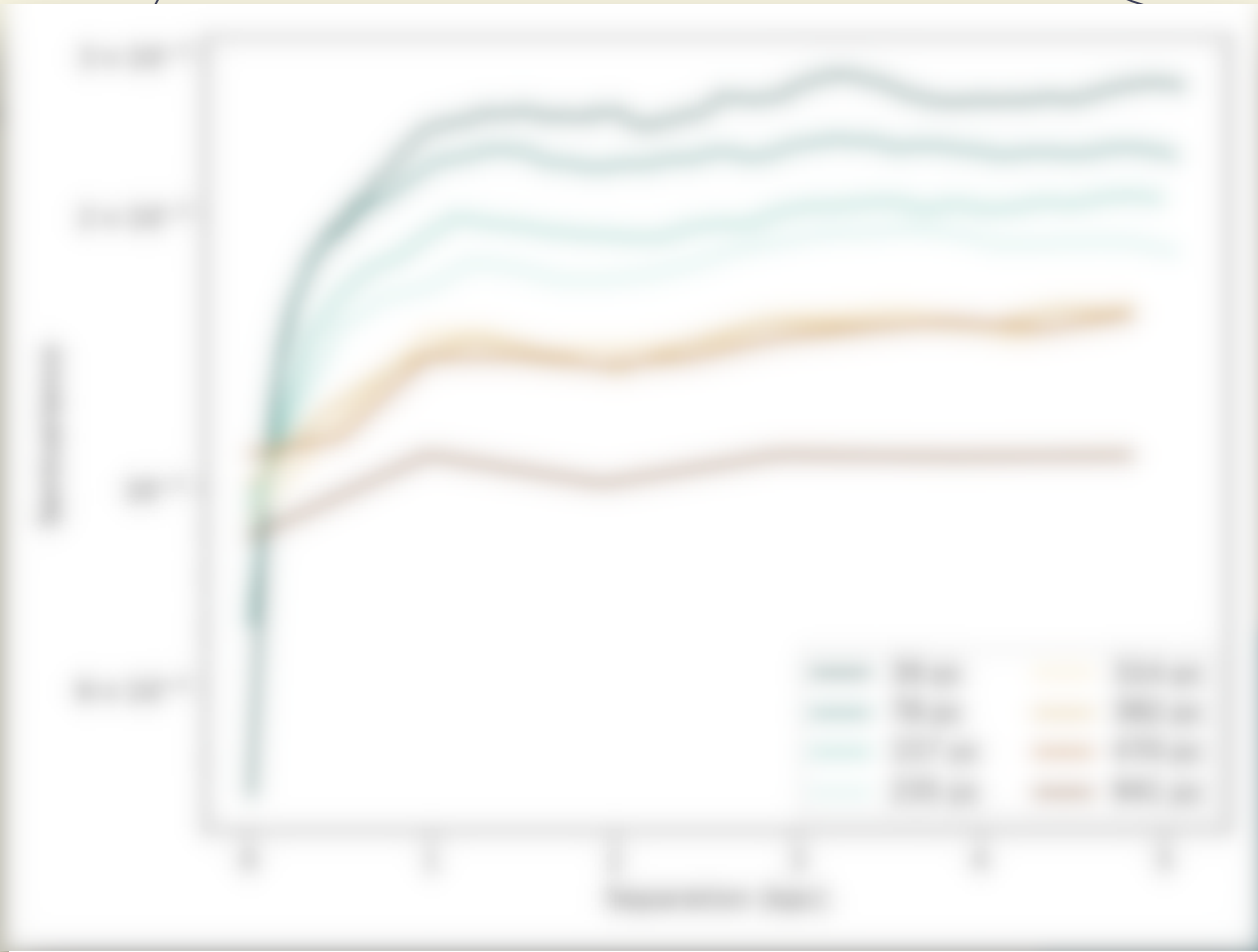
MUSE quality



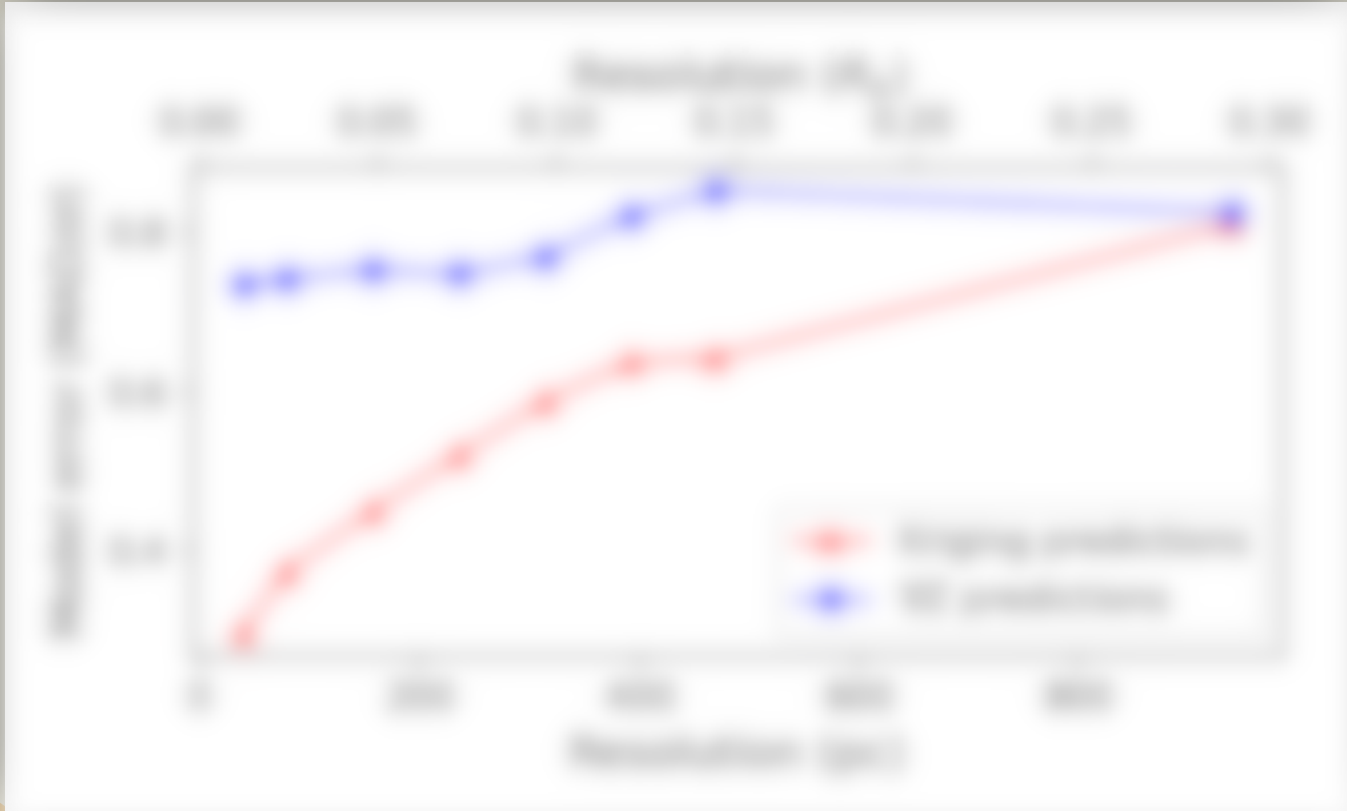
CALIFA/SAMI/  
MANGA quality



# RESOLUTION LIMITS

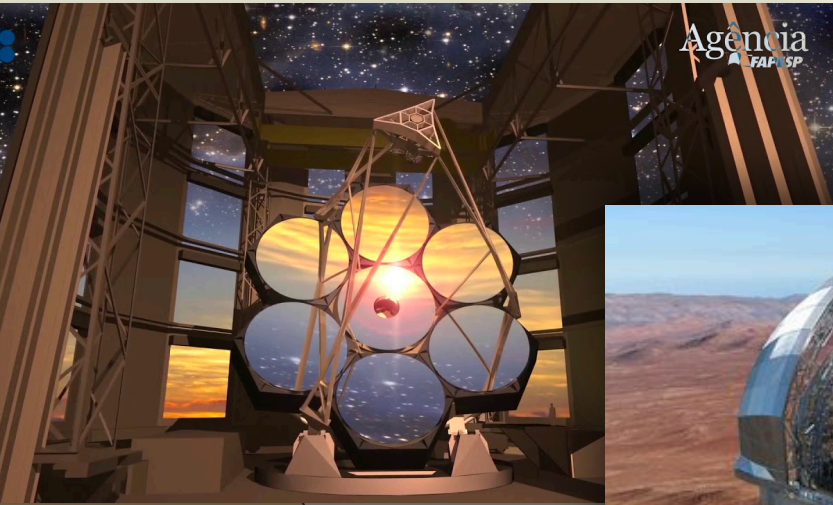


# PREDICTION ACCURACY vs RESOLUTION





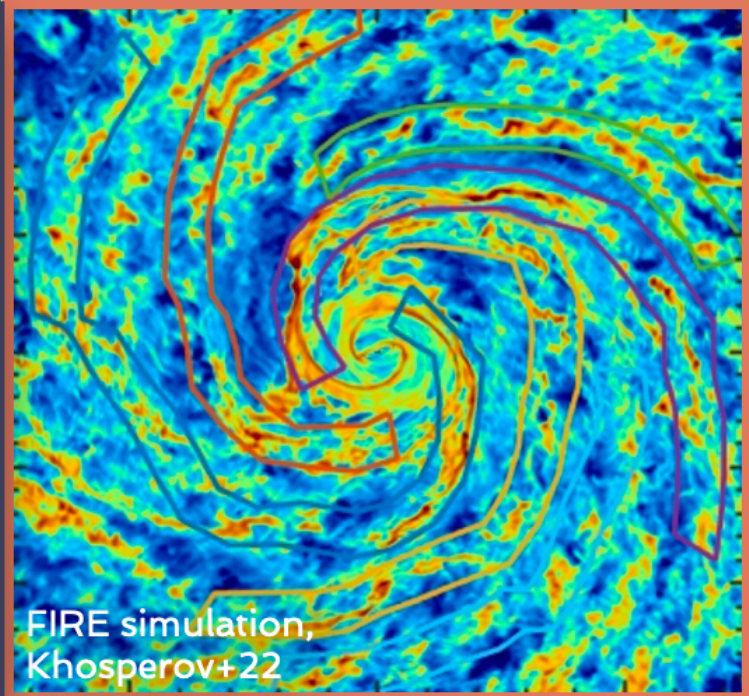
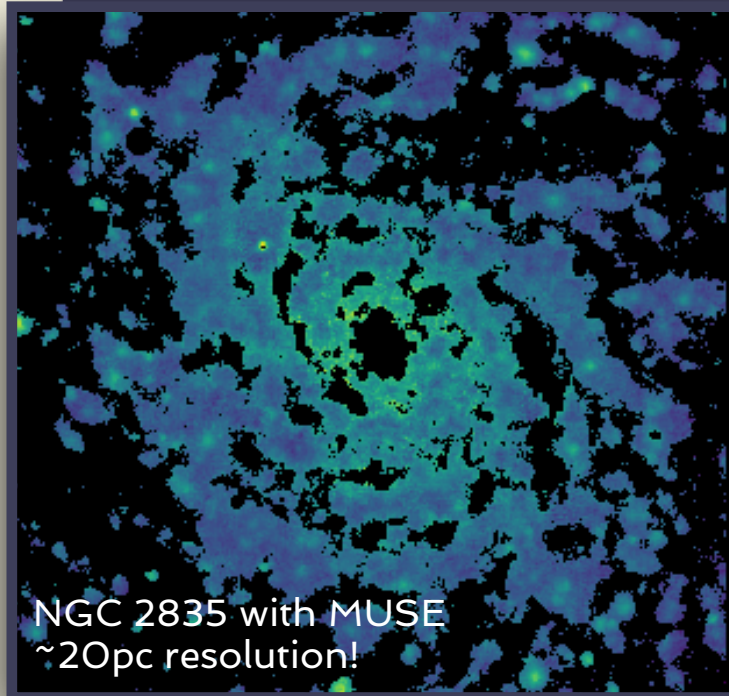
# THE ELTs WILL SOLVE ALL MY PROBLEMS



**IFU Resolution: 4-6 mas!**

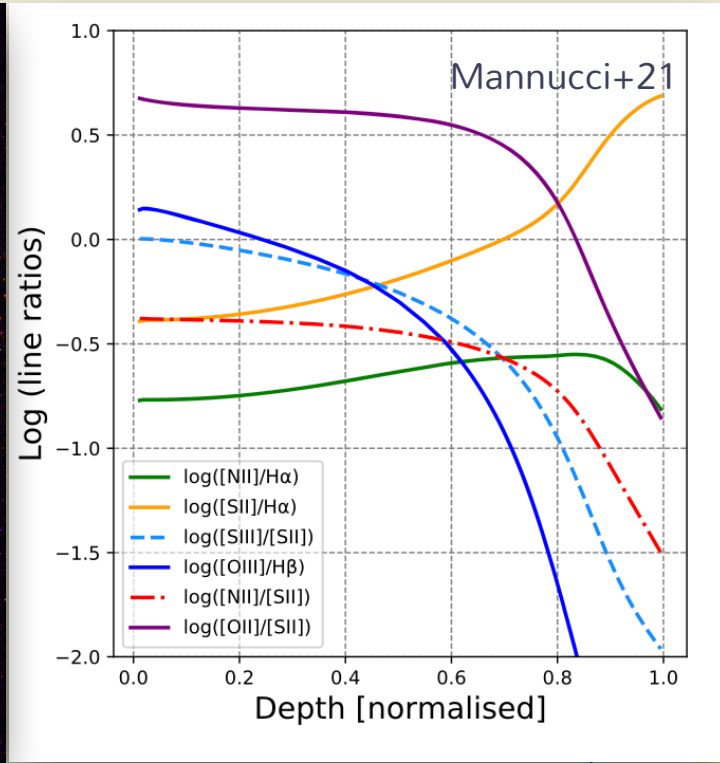
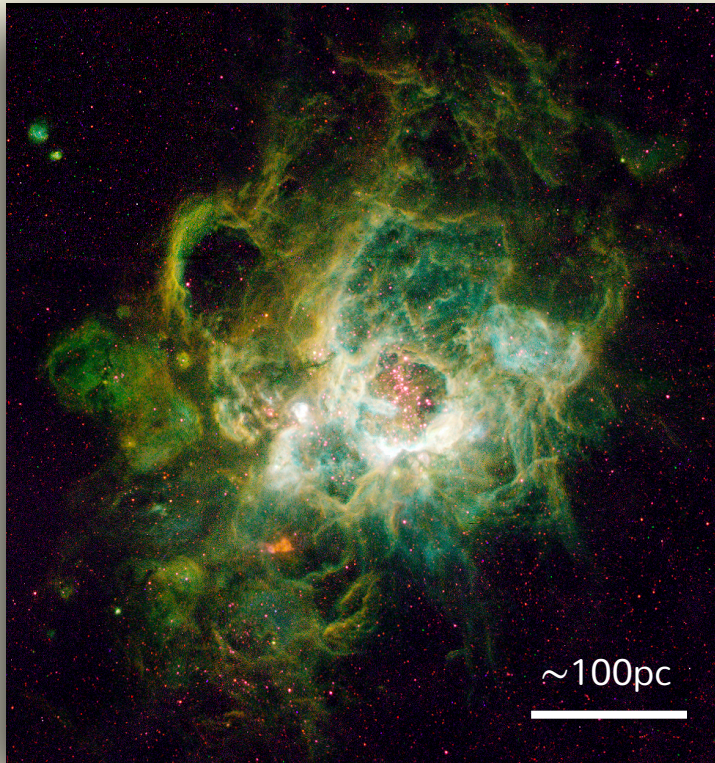
# PARSEC RESOLUTION FOR LOCAL GALAXIES

⇒ Direct comparisons to zoom-in simulations like FIRE!



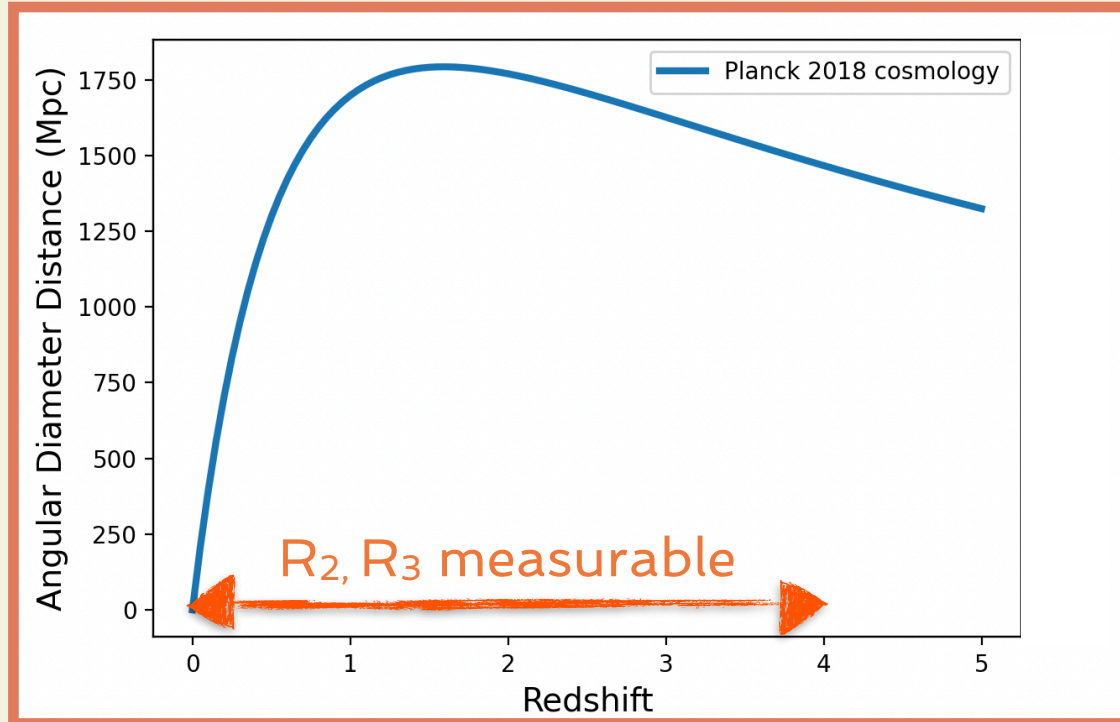
# PARSEC RESOLUTION METALLICITY

May need new diagnostics.



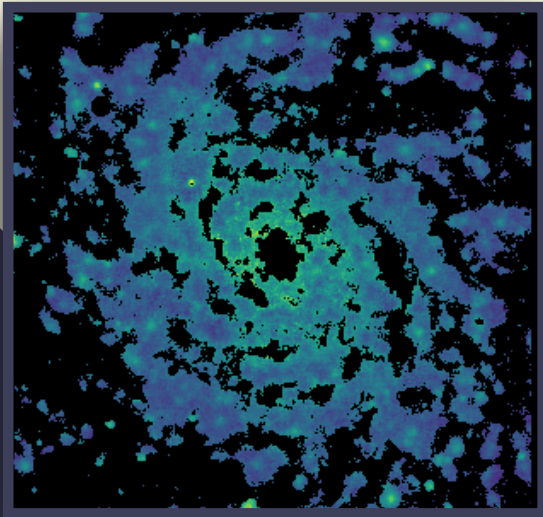
# GEOSTATISTICS AT HIGH REDSHIFT!

GMTIFS, HARMONI and IRIS will give **better than 100pc resolution** at **all redshifts**.

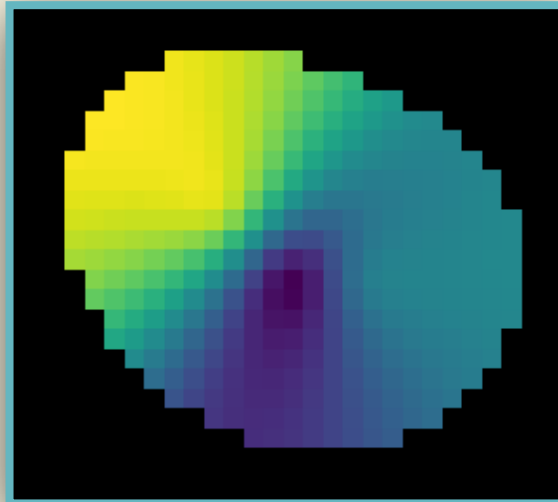


# SUMMARY

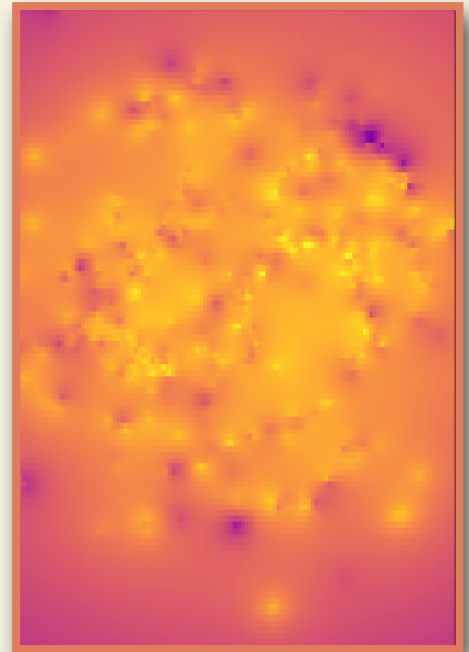
High resolution  
metallicity maps contain  
a wealth of information!



Lenstronomy's new  
module can correct for  
PSF smear and fit new  
models to JWST data



Geostatistical methods  
are ideal for ELTs' data



# THANKS!

DO YOU HAVE ANY QUESTIONS?

[methab@student.unimelb.edu.au](mailto:methab@student.unimelb.edu.au)



CREDITS: This presentation template based on a template by Slidesgo, and includes icons by Flaticon and The Noun Project